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Normal Modes of a Loss Cone

Plasma Slab with Steep Density Gradient

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The main text of this report (excluding the appendices), with minor changes, has been accepted for publication in Physics of Fluids.

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ABSTRACT

The electrostatic flute-like normal modes were found for a collisionless loss cone plasma in a uniform magnetic field with density varying sinusoidally in one direction in space. A local method, appropriate to kL >> 1, where k is the wave number and L the scale length, and a nonlocal method, appropriate to kL \leq 1, were both used; the ion Larmor radius a_i was not assumed to be small. The usual drift cone mode was found when $a_i << L$; other instabilities were found for $a_i \geq L$. In principle, the methods could be applied to other configurations of plasma.

Research sponsored by the Energy Research and Development Administration Contract E(04-3)-34-PA128.

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I. Introduction

In analyzing the kinetic instabilities of a non-uniform plasma the assumption is frequently made^{1,2} that the Larmor radius is small compared to the scale length of the plasma, although in many experiments³ the ion Larmor radius a, is comparable to the scale length L. We have found that it is often possible to find the normal modes, and even to use a local approximation, without assuming a, << L. The procedure starts with the integrodifferential Vlasov-Poisson dispersion relation for the electrostatic modes of a non-uniform collisionless plasma. Two different approximate methods can be used to solve this integral equation, reducing it to either a differential equation or a matrix equation. Each method is valid in a different regime, but neither method requires a, << L. The two methods have been used to find the properties of the most unstable flute-like mode of a deuterium loss cone plasma with a density varying sinusoidally in space in a direction perpendicular to a uniform magnetic field, for a variety of densities, density gradients, and mirror ratios. (The sinusoidally varying density was used because it simplifies the calculations and because it has been used in computer simulations of the drift cone mode^{4,5}.) The numerical results of the two methods are in good agreement (typically within 10%) in the regime where both are expected to be valid, thus increasing our confidence in them.

In Sec. II the integrodifferential dispersion relation will be derived for an arbitrary configuration of plasma and magnetic field, as well as for the special case of a uniform magnetic field, and we will discuss the reasons for the usual assumption that $a_i << L$, and under what circumstances this assumption can be dispensed with. In Sec. III the two methods of solving the integral equation will be described and

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applied to a loss cone plasma with sinusoidally varying density. The range of validity of each method will be found. In Sec. IV the numerical results obtained by the two methods will be given, and the numerical techniques used will be briefly described. In Sec. V the physical significance of these results (particularly for $a_i \ge L$) will be discussed, and a summary and conclusions will be presented in Sec. VI.

II. Dispersion Relation

The linearized Vlasov equation for a plasma with an electrostatic perturbation is

$$\frac{\mathrm{d}f_{1s}}{\mathrm{d}t}(\underline{x},\underline{y},t) = -\frac{q_s}{m_s} \quad \nabla \phi(\underline{x},t) \cdot \frac{\partial f_{0s}}{\partial \underline{y}}(\underline{x},\underline{y}) \tag{1}$$

where f_{1s} is the first-order perturbation in the distribution function, f_{0s} is the zero-order equilibrium distribution function (satisfying $\partial f_{0s}/\partial t = 0$), q_s is the charge and m_s is the mass for species s; and ϕ is the perturbed potential. Eq. (1) can be integrated over the zero-order particle orbits to obtain

$$f_{1s}(\bar{x},\bar{y},t) = \frac{q_s}{m_s} \int_{-\infty}^{t} d\tau \, \nabla \phi[\bar{x}'_s(\tau),\tau] \cdot \frac{\partial f_{0s}}{\partial \bar{y}_s} \left[\bar{x}'_s(\tau)\bar{y}_s'(\tau)\right]$$
(2)

where $x_s'(\tau)$ is the unperturbed orbit of a particle of species s, and $v_s'(\tau) \equiv dx_s'/d\tau$, with $x_s'(t) = x$ and $v_s'(t) = y$.

Eq. (2) is combined with the Poisson equation

$$\sum_{s} 4\pi q_{s} \int f_{1s}(x, y, t) dy + \nabla^{2} \phi(x, t) = 0$$

to yield the well-known result⁶

$$\sum_{s} \frac{4\pi q_{s}^{2}}{m_{s}} \int dy \int_{-\infty}^{t} d\tau \nabla \phi(\mathbf{x}_{s}, \tau) \cdot \frac{\partial f}{\partial \mathbf{v}_{s}'}(\mathbf{x}_{s}, \mathbf{v}_{x}') + \nabla^{2} \phi(\mathbf{x}, t) = 0 \quad (3)$$

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Eq. (3) can be fourier transformed in x: (See Appendix A)

$$\int d\underline{k} \quad \tilde{\phi} \quad (\underline{k}) \left(\int d\underline{k}' \exp(i\underline{k}' \cdot \underline{x}) \left\{ \sum_{s} \frac{4\pi q_{s}^{2}}{m_{s}} \int d\underline{y} \quad \int_{-\infty}^{0} d\tau \right\}$$
$$i\underline{k} \quad \cdot \quad \tilde{\partial} \frac{\tilde{f}_{s}}{\partial \underline{y}_{s}} \left(\underline{k} - \underline{k}', \underline{y}_{s}' \right) \exp[i\underline{k}' \cdot (\underline{x}_{s}' - \underline{x}) - i\omega\tau] \left\{ -k^{2} \right\} = 0$$
(4)

where
$$\tilde{f}_{s}(k, v) \equiv (2\pi)^{-3} \int dx f_{0s}(x, v) \exp(ik \cdot x)$$

 $\tilde{\phi}(k) \equiv (2\pi)^{-3} \int dx \phi(x, 0) \exp(-ik \cdot x)$

and $\phi(x,t)$ is assumed to vary in time as exp(-i ω t).

Eq. (4) can be written in compact form as

$$\int d\mathbf{k} \exp(\mathbf{i}\mathbf{k}\cdot\mathbf{x}) \tilde{\phi}(\mathbf{k}) D(\mathbf{x},\mathbf{k},\omega) = 0$$

or

$$D(\mathbf{x}, \mathbf{i}\nabla, \boldsymbol{\omega})\phi(\mathbf{x}) = 0$$
⁽⁵⁾

where D/k^2 is the usual dielectric function, defined by

$$D(\underline{x},\underline{k},\omega) \equiv -k^{2} + \sum_{s} D_{s}(\underline{x},\underline{k},\omega)$$

$$D_{s}(\underline{x},\underline{k},\omega) \equiv \frac{4\pi q_{s}^{2}}{m_{s}} \exp(-i\underline{k} \cdot \underline{x}) \int d\underline{k}' \exp(i\underline{k}' \cdot \underline{x})$$

$$\int d\underline{v} \int_{-\infty}^{0} d\tau \ i\underline{k} \cdot \frac{\partial f}{\partial \underline{v}_{s}'} (\underline{k}-\underline{k}',\underline{v}_{s}') \exp[i\underline{k}' \cdot (\underline{x}_{s}'-\underline{x}) - i\omega\tau]$$
(6)

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If the magnetic field varies over a distance comparable to a Larmor radius, then it is difficult to calculate the zero-order orbits $x_s'(\tau)$, and to find an equilibrium distribution function $f_{0s}(\underline{x},\underline{v})$. In a low- β plasma, however, it is possible for the scale length of the magnetic field (i.e. $B_0/|\nabla B_0|$) to be much greater than the scale length of the plasma, so we can assume the magnetic field $\underline{B}_0(x)$ to be uniform in space, or varying gradually, in calculating the particle orbits and the constants of motion, without assuming that $f_{0s}(\underline{x},\underline{v})$ varies gradually as a function of \underline{x} . For uniform \underline{B}_0 , the constants of motion are v_1 , v_{\parallel} , and $\underline{x}_{gc'} \equiv \underline{x}_1 + \underline{v} \times \hat{B}_0 \omega_{cs}^{-1}$ (where \hat{B}_0 is a unit vector parallel to \underline{B}_0), so

$$f_{0s}(x, v) = g_{s}(v_{\perp}, v_{\parallel}, x_{gc^{\perp}})$$

and (see Appendix A)

$$\tilde{f}_{s}(\underline{k},\underline{v}) = \exp[-i\underline{k}\cdot(\underline{v} \times \hat{B}_{0}) \omega_{cs}^{-1}]\tilde{g}_{s}(\underline{v}_{\underline{i}},\underline{v}_{\parallel},\underline{k})$$
(7)

where

$$\tilde{g}_{s}(v_{1},v_{\parallel},\underline{k}) \equiv (2\pi)^{-3} \int d\underline{x} \ g_{s}(v_{1},v_{\parallel},\underline{x}) \ \exp(i\underline{k} \cdot \underline{x})$$
The particle orbits are
$$\underline{x}_{s}'(\tau) = \underline{x} + \frac{\underline{v}_{1}}{\underline{\omega}_{cs}} \sin \underline{\omega}_{cs} \tau + \frac{\underline{v} \times \hat{B}_{0}}{\underline{\omega}_{cs}} (1 - \cos \underline{\omega}_{cs} \tau)$$

$$+ v_{\parallel} \hat{B}_{0} \tau \qquad (8)$$

[For non-uniform $B_0(x)$, the particle drifts would have to be included in Eq.(8) as well, and the constants of motion replaced by adiabatic constants of motion.]

Putting Eqs. (7) and (8) into Eq. (6), and using the Bessel function identity 7

$$\exp(i \ a \ \sin \theta) = \sum_{\ell} J_{\ell}(a) \ \exp(i\ell\theta)$$

we can do the integration over τ and over the azimuthal direction of \underline{v} to obtain (see Appendix A)

$$D_{s}(\underline{x},\underline{k},\omega) = \frac{4\pi q_{s}^{2}}{m_{s}} \exp(-i\underline{k} \cdot \underline{x}) \int d\underline{k}' \exp(i\underline{k}' \cdot \underline{x}) \int_{-\infty}^{\infty} dv_{\parallel}$$

$$\int_{0}^{\infty} 2\pi v_{\underline{i}} dv_{\underline{i}} \sum_{\ell} [\hat{k}_{\underline{i}} \cdot (\hat{k}_{\underline{i}}' + i\hat{B}_{0} \times \hat{k}_{\underline{i}}')]^{\ell} J_{\ell}(\frac{\underline{k}_{\underline{i}}v_{\underline{i}}}{\omega_{cs}}) J_{\ell}(\frac{\underline{k}_{\underline{i}}'v_{\underline{i}}}{\omega_{cs}})$$

$$(\omega - \ell \omega_{cs} - k_{\parallel}v_{\parallel})^{-1} \left[\frac{i\underline{k} \cdot (\hat{B}_{0} \times \underline{k}')}{\omega_{cs}} - \tilde{g}_{s}(v_{\underline{i}}, v_{\parallel}, \underline{k} - \underline{k}') - \frac{\ell \omega_{cs}}{v_{\underline{i}}} - \frac{\delta}{\partial v_{\underline{i}}} - \tilde{g}_{s}(v_{\underline{i}}, v_{\parallel}, \underline{k} - \underline{k}')\right]$$

$$(9)$$

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where $\mathbf{k}_{\parallel} \equiv \mathbf{k} \cdot \hat{\mathbf{B}}_{0}, \ \mathbf{k}_{\perp} \equiv \mathbf{k} - \mathbf{k}_{\parallel} \hat{\mathbf{B}}_{0}, \ \mathbf{\hat{k}}_{\perp} \equiv \mathbf{k}_{\perp}/\mathbf{k}_{\perp}$.

If the scale length of the plasma is much greater than a Larmor radius, then only $k' \approx k$ will contribute significantly to the integral over k' in Eq.(9), and $D_s(x,k,\omega)$ can be expressed in terms of $g_s(v_1,v_{\parallel},x)$, and its gradient at the same point x in space⁸: (see Appendix B)

$$D_{s}(\underline{x},\underline{k},\omega) = -\frac{4\pi q_{s}^{2}}{m_{s}} \int dv_{\parallel} \int 2\pi v_{\perp} dv_{\perp} \sum_{\ell} J_{\ell}^{2} (k_{\perp} v_{\perp} / \omega_{cs})$$

$$(\omega - \ell \omega_{cs} - k_{\parallel} v_{\parallel})^{-1} [\underline{k} \cdot (\hat{B}_{0} \times \nabla g_{s}) / \omega_{cs} + k_{\parallel} \partial g_{s} / \partial v_{\parallel}$$

$$+ (\ell \omega_{cs} / v_{\perp}) \partial g_{s} / \partial v_{\perp}] \qquad (10)$$

If a suitable $g_{s}(v_{1},v_{\parallel},x)$ is chosen, however, it may be possible to integrate Eq. (9) over k' analytically; in this case there is no need to assume that the Larmor radius is small compared to the scale length.

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III. Methods of Solution

We will consider only modes with $k_{\parallel} = 0$, so the integration over v_{\parallel} can be done by replacing $g_{s}(v_{\perp},v_{\parallel},x)$ by

$$g_{s^{\perp}}(v_{\perp},\tilde{x}) \equiv \int dv_{\parallel} g_{s}(v_{\perp},v_{\parallel},\tilde{x})$$

[with $\tilde{g}_{s!}(v_1, k)$ similarly defined] in Eq. (9). We will further assume that $g_{s!}$ is Maxwellian and that there are no temperature gradients

$$g_{s^{\perp}}(v_{1},x) = (2\pi v_{s}^{2})^{-1} \exp(-v_{1}^{2}/2v_{s}^{2})n_{s}(x).$$

A loss-cone distribution can be constructed by subtracting two 9 Maxwellians 9

$$g_{s!}(v_{1},x) = [2\pi v_{s}^{2}(1-R^{-1})]^{-1} n_{s}(x)$$

$$[exp(-v_{1}^{2}/2v_{s}^{2})-exp(-Rv_{1}^{2}/2v_{s}^{2})]$$
(11)

where R is the mirror ratio. The two components can be formally treated as two different species, with thermal velocities $v_s^{}$ and $v_s^{}$ R^{-1/2}. For the rest of this paper, unless otherwise noted, all species will be assumed to have Maxwellian distributions, with the understanding that loss cone species will be formally treated as two different Maxwellian species.

Finally, we assume a slab geometry, with density $n_s(x)$ depending only on x and \underline{B}_0 in the z direction. Then Eq. (9) becomes (see Appendix C)

$$D_{s}(\tilde{x}, \tilde{k}, \omega) = \frac{4\pi q_{s}^{2}}{m_{s}} \int_{-\infty}^{\infty} dk'' \exp(ik''x) \sum_{\ell} \exp(i\ell\alpha)$$

$$\exp[-(k^{2}+k'^{2})a_{s}^{2}/2]I_{\ell} (k k'a_{s}^{2}) (\omega-\ell\omega_{cs})^{-1}$$

$$(-ik_{y}k''/\omega_{cs} + \ell \omega_{cs}/v_{s}^{2}) \tilde{n}_{s}(k'') \qquad (12)$$

where k " $\equiv k \frac{1}{x} - k_x$, k_y ' $\equiv k_y$, $\alpha \equiv i \log_e \left(\frac{k \frac{1}{x} + ik_y}{k'}\right) - i \log_e \left(\frac{k \frac{1}{x} + ik_y}{k}\right)$ is the angle between the vectors k and k',

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$$a_{s} \equiv v_{s}/\omega_{cs},$$

and $\tilde{n}_{s}(k'') \equiv (2\pi)^{-1} \int_{-\infty}^{\infty} dx \exp(-ik''x)n_{s}(x).$

For a sinusoidal density profile,

$$n_{s}(x) = (1 + \Delta_{s} \cos k_{0}x) n_{s0}$$
(13)

we have $\tilde{n}_{s}(k'') = (2\pi)^{-1} n_{s0}[\delta(k'') + (\Delta_{s}/2)\delta(k'' + k_{0}) + (\Delta_{s}/2)\delta(k'' - k_{0})].$ (There is no requirement that Δ_{s} be small, in fact usually we will use $\Delta_{s} \approx 1$; see Fig. 1.)

Then the integral over k" in Eq. (12) becomes a sum over three terms, $k'' = 0, -k_0 + k_0$. Since we are assuming no equilibrium electric fields, we must have

$$\sum_{s} q_{s} \int_{\tilde{u}} dv f_{0s}(\tilde{x}, \tilde{v}) = 0 \text{ for all } \tilde{x}$$

or (see Appendix D)

$$\sum_{s} q_{s} n_{s0} \Delta_{s} \exp(-k_{0}^{2} a_{s}^{2}/2) = 0$$
 (14a)

and

$$\sum_{s} q_{s} n_{s0} = 0 \tag{14b}$$

The eigenmodes $\phi(x)$ satisfying Eq. (5) must be of the form

$$\phi(\mathbf{x}) = \phi(\mathbf{x}) \exp(\mathbf{i}\mathbf{k}_{\mathbf{y}} \mathbf{y})$$

since $D(x,k,\omega)$ has no dependence on y, and we have been assuming $k_{\parallel} = 0$. We can then write Eq. (5) as

$$\int dk_{x} \exp(ik_{x} x) D(x,k_{x},\omega) \tilde{\phi}(k_{x}) = 0$$





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$$D(x,-i\frac{\partial}{\partial x},\omega)\phi(x) = 0$$
(15)

where $k_{\ensuremath{\mathbf{y}}}$ is understood to be a parameter, and

$$\tilde{\phi}(k_x) \equiv (2\pi)^{-1} \int_{-\infty}^{\infty} dx \phi(x) \exp(-ik_x x)$$

We want to find the normal modes $\phi_n(x)$ and the corresponding frequencies ω_n which satisfy Eq. (15) with appropriate boundary conditions.

A. Local Method

One approach is to expand $D(x,k_x)$ in a power series around some $x = x_L$, $k_x = k_L$, and $\omega = \omega_L$ (all complex, in general) such that

$$D(x_{L},k_{L},\omega_{L}) = \frac{\partial D}{\partial x} (x_{L},k_{L},\omega_{L}) = \frac{\partial D}{\partial k_{x}} (x_{L},k_{L},\omega_{L}) = 0$$
(16)

(Note that a wave packet localized around x_L , with well-defined wave number k_L , would have a frequency close to ω_L , and would have zero group velocity and acceleration.)

Then

$$D(\mathbf{x}, \mathbf{k}_{\mathbf{x}}, \omega) = \frac{1}{2} (\mathbf{x} - \mathbf{x}_{\mathbf{L}})^{2} \frac{\partial^{2} D}{\partial \mathbf{x}^{2}} + \frac{1}{2} (\mathbf{k}_{\mathbf{x}} - \mathbf{k}_{\mathbf{L}})^{2} \frac{\partial^{2} D}{\partial \mathbf{k}_{\mathbf{x}}^{2}} + (\mathbf{x} - \mathbf{x}_{\mathbf{L}}) (\mathbf{k}_{\mathbf{x}} - \mathbf{k}_{\mathbf{L}}) \frac{\partial^{2} D}{\partial \mathbf{x} \partial \mathbf{k}_{\mathbf{x}}} + (\omega - \omega_{\mathbf{L}}) \frac{\partial D}{\partial \omega} + \dots$$
(17)

where the derivatives of D are evaluated at $x_L^{}, k_L^{}, \omega_L^{}$, and Eq. (15) becomes

$$A \frac{d^2\psi}{dx^2} + Bx \frac{d\psi}{dx} + C x^2\psi + \lambda\psi = 0$$
(18)

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where

A =
$$-1/2 \frac{\partial^2 D}{\partial k_x^2}$$
, B = $-i \frac{\partial^2 D}{\partial k_x \partial x}$, C = $1/2 \frac{\partial^2 D}{\partial x^2}$,

or

$$\lambda \equiv (\omega - \omega_L) \frac{\partial D}{\partial \omega}$$
, all evaluated at x_L, k_L , and ω_L ;

and $\psi(\mathbf{x}) \equiv \exp(-ik_{L}\mathbf{x})\phi(\mathbf{x} + \mathbf{x}_{L})$.

This approach is similar to that used by other authors 10^{-12} , who have, however, taken $k_L = 0$ to satisfy Eq. (16), and hence B = 0 in Eq. (18). Furthermore, the truncation of Eq. (17) has usually been justified by assuming small Larmor radius compared to scale length, and we wish to emphasize that this assumption is not the only one which will justify Eq. (17).

Taking $\psi(x) \rightarrow 0$ as $x \rightarrow \pm \infty$ as the boundary conditions, Eq. (18) has the solutions (see Appendix E)

$$\psi_n(x) = H_n(n^{1/2}x) \exp(-\beta x^2/2)$$
(19a)

$$\lambda_{n} = (2nn + \beta)A \tag{19b}$$

where

$$n \equiv [(B/2A)^2 - C/A]^{1/2}, \beta \equiv n + B/2A,$$

and H_n is the order n Hermite polynomial, n=0,1,2... Then

$$\phi_n(x) = H_n[\eta^{1/2}(x-x_L)] \exp(-\beta(x-x_L)^2/2 + ik_L x)$$
 (20a)

$$\omega_{n} = \omega_{L} - (n\eta + \beta/2) \left(\frac{\partial D}{\partial \omega}\right)^{-1} \frac{\partial^{2} D}{\partial k_{x}^{2}}$$
(20b)

A typical normal mode $\phi_n(x)$ for n=0 is shown in Fig. 1.

These solutions are valid provided we can neglect the higher derivatives of D in Eq. (17), and provided the Stokes lines of D have the proper

and

topological behavior in the complex x-plane with respect to the real axis to allow the boundary conditions to be satisfied.¹³ We conjecture that Eq.(20) is a valid solution for the first few modes (n=0,1,2...) provided (see Appendix F) 1/2

$$\frac{1}{L(k_{L}^{2} + k_{y}^{2})} >> 1$$
(21)

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(where L is the scale length of the plasma) except perhaps for special unusual cases. For instabilities requiring $k_{\underline{i}}a_{\underline{i}} \gtrsim 1$ (e.g. drift cone¹⁴ or drift cyclotron¹⁵), $a_{\underline{i}} << L$ is sufficient for inequality (21) to be satisfied, but it is certainly not necessary.

Since ω_n depends linearly on n, the fastest growing mode must be the n = 0 mode for some set (x_L, k_L, ω_L) satisfying Eq. (16), or else it cannot be found by the local method at all. [It has turned out in almost every case we have examined for which inequality (21) is satisfied, that the fastest growing mode can in fact be found by the local method using Eqs.(16) and (20), and is thus the n = 0 mode, but it is possible to construct models where this will not be true.] The fastest growing mode will then be

$$\phi(\mathbf{x}) = \exp[-\beta(\mathbf{x} - \mathbf{x}_{L})^{2}/2 + i\mathbf{k}_{L}\mathbf{x}]$$
(22a)

$$\omega = \omega_{\rm L} - \frac{\beta}{2} \left(\frac{\partial D}{\partial \omega}\right)^{-1} \frac{\partial^2 D}{\partial k_{\rm x}^2} \approx \omega_{\rm L}$$
(22b)

where we require that x_L , k_L , ω_L satisfy

$$\operatorname{Im}\left[\delta\left(\frac{\partial D}{\partial \omega}\right)^{-1} \frac{\partial^2 D}{\partial k_x^2}\right] > 0$$

as well as Eq. (16).

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A large class of solutions (x_L, k_L, ω_L) satisfying Eq. (16) have $k_L = 0$, since this guarantees that $\partial D/\partial k_x = 0$, and it is then only necessary to find x_L and ω_L satisfying D = 0 and $\partial D/\partial x = 0$. In this case $\partial^2 D/\partial k_x \partial x = 0$, and $\delta = \beta = (\frac{\partial^2 D}{\partial x_L^2} / \frac{\partial^2 D}{\partial x_L^2})^{1/2}$

$$\partial x^2 \partial k_2$$

B. Nonlocal Method

When inequality (21) is not satisfied a different approach is required. We Fourier transform $D(x,k_x,\omega)$

$$D(x,k_{x},\omega) = (2\pi)^{-1} \int_{-\infty}^{\infty} dk'' \exp(ik''x) \tilde{D}(k'',k_{x},\omega)$$

where

$$\tilde{D}(k'',k_{x},\omega) \equiv \int_{-\infty}^{\infty} dx \exp(-ik''x) D(x,k_{x},\omega)$$
(23)

Putting Eq. (23) into Eq. (15), multiplying by exp(-ik'x) for arbitrary k', and integrating over x, we obtain

$$\int dk'' \tilde{D}(k'',k'-k'',\omega) \tilde{\phi}(k'-k'') = 0$$
 (24)

for all k'.

For the sinusoidal density profile given by Eq. (13), using Eqs.(6) and (12), Eq. (23) becomes

$$\tilde{D}(k'',k_{x},\omega) = G_{0}(k_{x},\omega) \ \delta(k'') + G_{+1}(k_{x},\omega) \ \delta(k''-k_{0}) + G_{-1}(k_{x},\omega) \ \delta(k'' + k_{0})$$
(25)

-13-

where

$$\begin{split} & G_{0} \equiv -k^{2} + \sum_{s} G_{0,s}, \\ & G_{\pm 1} \equiv \sum_{s} G_{\pm 1,s}, \\ & G_{0,s}(k_{x},\omega) \equiv \frac{\omega_{ps}^{2}}{v_{s}^{2}} \exp(-k^{2}a_{s}^{2}) \sum_{\ell} I_{\ell}(k^{2}a_{s}^{2})(\omega/\omega_{cs} - \ell)^{-1} \ell \\ & G_{\pm 1,s}(k_{x},\omega) \equiv \frac{\Delta_{s}\omega_{ps}^{2}}{2v_{s}^{2}} \exp[-(k^{2}+k_{\pm 1}^{2})a_{s}^{2}/2] \\ & \sum_{\ell} \exp(i\ell\alpha_{\pm 1}) I_{\ell}(k_{\pm 1}a_{s}^{2})(\omega/\omega_{cs} - \ell)^{-1} (\ell \pm ik_{y}k_{0}a_{s}^{2}) \\ & k_{\pm 1}^{2} \equiv (k_{x} \pm k_{0})^{2} + k_{y}^{2}, \\ & a_{\pm 1} \equiv i \log_{e} (\frac{k_{x}\pm k_{0} + ik_{y}}{k_{\pm 1}}) -i \log_{e}(\frac{k_{x} + ik_{y}}{k}) \text{ is the angle between} \\ & k \text{ and } k_{\pm 1}, \text{ and} \\ & \omega_{ps}^{2} \equiv 4\pi n_{s0} q_{s}^{2}/m_{s}. \end{split}$$

Then the integral in Eq. (24) becomes a discrete sum

$$G_{0}(k',\omega) \tilde{\phi}(k') + G_{+1}(k'-k_{0},\omega) \tilde{\phi}(k'-k_{0}) + G_{-1}(k'+k_{0},\omega)$$
$$\tilde{\phi}(k'+k_{0}) = 0$$
(26)

If we require $\phi(\mathbf{x})$ to be periodic with the same periodicity

 k_0 as the density [appropriate for a simulation with periodic boundary conditions; non-periodic $\phi(x)$ will be considered later], then $\tilde{\phi}(k_x)$ will vanish except at $k_x = pk_0$ for integer p:

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$$\tilde{\phi}(k_x) = \sum_{p = -\infty}^{\infty} \phi_p \delta(k_x - pk_0)/2\pi$$

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where

$$\phi_{p} \equiv (k_{0}/2\pi) \int_{0}^{2\pi/k_{0}} dx \phi(x) \exp(-ipk_{0}x)$$

$$\phi(x) = \sum_{p} \phi_{p} \exp(ipk_{0}x) \qquad (27)$$

and

[It should be kept in mind that in general
$$\phi(x)$$
 is complex, so there
is no relation between ϕ_p and ϕ_{-p} , as there would be if $\phi(x)$ were always
real.]

Then, setting $k' = pk_0$, Eq. (26) becomes

$$A_{p,p-1}(\omega)\phi_{p-1} + A_{p,p}(\omega)\phi_{p} + A_{p,p+1}(\omega)\phi_{p+1} = 0$$
 (28)

where

$$A_{p,p'}(\omega) \equiv G_{(p-p')}(p'k_0,\omega).$$

Eq. (28) must be satisfied for all integers p, since Eqs. (24) and (26) are satisfied for arbitrary k'. So, we require

$$\det A_{p,p'}(\omega) = 0$$
(29)

If, for the modes we are interested in, ϕ_p is negligibly small for all $p \ge p_{max}$, and all $p \le p_{min}$ for some finite p_{max} and p_{min} , then we can truncate the matrix $A_{p,p}$, at $p > p_{max}$ and $p < p_{min}$. We can then solve Eq. (29) to find the normal mode frequencies ω (in general there will be $p_{max}-p_{min} + 1$ of them for each cyclotron harmonic), use Eq. (28) to find the ϕ_p for each normal mode, and then use Eq. (27) to find $\phi(x)$.

This method is easily generalized to periodic density profiles $n_s(x)$ other than the sinusoidal profile given by Eq. (13). For

$$n_{s}(x) = \sum_{j=-\infty}^{\infty} n_{s,j} \exp(ijk_{0}x)$$
(30)

Eq. (25) becomes

$$\tilde{\tilde{D}}(k'',k_{x},\omega) = \sum_{j=-\infty}^{\infty} G_{j}(k_{x},\omega) \,\delta(k'' - jk_{0})$$
(31)

with G_j defined like G_{±1} in Eq. (25) but with $\Delta_s/2$ replaced by n_{s,j}/n_{s0}, and ±k₀ replaced by + jk₀. Eq. (28) then becomes

$$\sum_{\mathbf{p}'} \mathbf{A}_{\mathbf{p},\mathbf{p}'} \boldsymbol{\phi}_{\mathbf{p}'} = 0$$

which again leads to Eq. (29). The advantage in using a sinusoidal density profile is that the matrix $A_{p,p}$, is tri-diagonal, and Eq. (29) is easier to solve numerically than it would be for an arbitrary matrix. Adding terms with $2k_0$, for example, would result in a penta-diagonal matrix. On the other hand, an isolated slab of plasma with $a_i/L-1$ would be more realistically modelled by a periodic density profile including higher harmonics of k_0 .

It is also not difficult to generalize this method to potentials $\phi(\mathbf{x})$ which are not periodic in x. It can be shown using Floquet's theorem that, except for a measure zero subset of the parameter space, all normal mode potentials must be of the form (see Appendix G)

$$\phi(\mathbf{x}) = \exp(\mathbf{i}K\mathbf{x}) \sum_{\mathbf{p}} \phi_{\mathbf{p}} \exp(\mathbf{i}\mathbf{p}\mathbf{k}_{0}\mathbf{x})$$
(32)

for some K, $0 < K < k_0$. Then we can still use Eqs. (28) and (29), but we must redefine $A_{p,p'}$ as

$$A_{p,p'}(\omega) \equiv G_{(p-p')}(p'k_0 + K,\omega)$$
 (33)

The nonlocal method will only be practical if we can use $p_{max}^{-p}min$ not too large; otherwise the numerical solution of Eq. (29) will be too difficult. In order to estimate the conditions under which the nonlocal method is practical, we will temporarily assume that inequality

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(21) is at least marginally satisfied, so that Eq. (20a) is at least qualitatively correct for the fastest growing mode. Then, using the definition of ϕ_p in Eq. (27), we find

$$\phi_p \propto \exp[-(pk_0 - k_L)^2 / 2\beta - ipk_0 x_L]$$

so we must choose $P_{max} - P_{min}$ to be at least a few times $|\beta|^{1/2} k_0^{-1}$. We estimate that β is on the order of $k_0 (k_L^2 + k_y^2)^{1/2}$ from the definitions following Eqs. (18) and (19), since the scale length of D in x will be about k_0^{-1} , while the "scale length" of D in k_x will be about $(k_L^2 + k_y^2)^{1/2}$, and we can estimate the relative magnitudes of the various second derivatives of D from these scale lengths. It then follows that

$$p_{max} - p_{min} \ge k_0^{-1} (k_L^2 + k_y^2)^{1/2}$$
 (34)

and we need $p_{max} - p_{min} >> 1$ if inequality (21) is satisfied. So the nonlocal method is not a practical way to find the fastest growing mode in those cases where the local method is valid, whereas it is practical when the local method is marginally valid. We cannot say in general whether the nonlocal method is practical when inequality (21) is not even marginally satisfied. However, in all such cases which we have investigated numerically, $P_{max} - P_{min}$ can be chosen small (of order unity). Thus the local and nonlocal methods complement each other; each method works best in the regime where the other method is least efficient or least valid.

IV. Numerical Techniques

To find the normal modes and frequencies using the nonlocal method, the roots of Eq. (29) were found numerically in the region of interest of the complex ω plane, using a standard root-finding routine^{16,17}

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(modified by the authors) based on Muller's method.¹⁸ Usually a small value of $k_{y_{i}}$ (typically 0.5) was used initially, since the root-finder was most efficient near the real axis (rarely missing roots with Im $\omega/\omega_{ci} \leq 1$) and at small $k_{y_{i}}$ the roots tended to be near the real axis. Then each root found at the initial $k_{y_{i}}$ was followed out to larger $k_{y_{i}}$; in this way it was possible to find roots far from the real axis at large $k_{y_{i}}$, which the root-finder would very likely miss if $k_{a_{i}}$ were large initially. For each root found, the eigenvector $(\phi_{pmin}, \dots \phi_{pmax})$ was calculated from Eq. (28), and $\phi(x)$ [from Eq. (27)] was plotted out if desired. In practice P_{min} was always set equal to $-P_{max}$, since the matrix elements $A_{p,p}$, for p,p' < 0 could be found using the identity $A_{p,p'} = A_{-p'}, -p$ [following from the definitions of $A_{p,p'}$ and $G_0, G_{\pm 1}$ after Eqs. (28) and (25)]; thus little time would be saved by using $|P_{min}| \neq |P_{max}|$ if $P_{min} < 0 < P_{max}$. (Some time would be saved by using $|P_{min}| \leqslant |p_{max}|$ to find modes with $\phi_{-P_{max}} < \phi_{Pmax}$.) 10

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Since the time required to calculate det $A_{p,p}$, is proportional to p_{max} , and the number of roots is proportional to p_{max} , the time required to find all the roots in a given range of k_y and ω was roughly proportional to p_{max}^2 . A typical run for a loss cone deuterium plasma with $\omega_{pe} \approx \omega_{ce}$ and $k_0 a_i \approx 1$, in the range of k_y and ω typical of the drift cone instability ($0.5 \leq k_y a_i \leq 20$, $|\omega| / \omega_{ci} < 2.5$ at $k_y a_i = 0.5$), using $p_{max} = 3$, took one minute on the CDC 7600 computer. However, for $p_{max} \geq 5$ the roots were so close together, especially near the cyclotron harmonics, that the root finder could not find many of them. In order to avoid this problem and to reduce the computation time, an approximation was sometimes employed in which the asymptotic form for the Bessel function¹⁹

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$$e^{-x} I_{\ell}(x) \approx (2\pi x)^{-1/2} \exp(-\ell^2/2x), (x >> \ell >>1)$$

was used, and the sum over & replaced by an integral, for the ion Then the definitions after Eq. (25) become, for the ion terms (see Appendix H

$$G_{0,s}(k_{x},\omega) = -(\omega_{ps}^{2}/v_{s}^{2}) [1 + \zeta_{0} Z(\zeta_{0})]$$

$$G_{\pm 1,s}(k_{x},\omega) = -(\omega_{ps}^{2} \Delta_{s}/2v_{s}^{2}) \exp(-k_{0}^{2}a_{s}^{2}/2][1 + \zeta_{\pm 1}Z(\zeta_{\pm 1})]$$
(35)

where

terms.

$$\zeta_0 = \omega/(\sqrt{2}kv_s)$$

$$\zeta_{\pm 1} = (\omega \pm ik_y k_0 a_s v_s)/(\sqrt{2} kv_s)$$

and Z(ζ) is the usual plasma dispersion function 20

$$Z(\zeta) \equiv \pi^{-1/2} \int dt \exp(-t^2)/(t-\zeta)$$

This approximation is valid when Im $\omega \ge \omega_{ci}$, ka_i \ge 1, and and is equivalent to using straight line orbits for the ions $k_0 \lesssim k$, when integrating the Vlasov equation over the zero-order orbits. When this approximation is used, the modes associated with the ion cyclotron harmonics (Bernstein waves) are eliminated, allowing higher pmax to be used without confusing the root-finder (no problems were encountered up to $p_{max} = 9$, as long as the Z(ζ) routine had high enough precision) and greatly reducing the computation time for a given p_{max} (typically by a factor of 10).

The electron terms were evaluated in the limit of zero Larmor radius, ka_e << 1:

$$G_{0,e}(k_{x},\omega) = -\omega_{pe}^{2}k^{2}/(\omega_{ce}^{2} - \omega^{2})$$

$$G_{\pm 1,e}(k_{x},\omega) = -(\omega_{pe}^{2}\Delta_{e}/2) [k k_{\pm 1}/(\omega_{ce}^{2} - \omega^{2}) \pm i k_{0}k_{y}/|\omega_{ce}|\omega]$$
(36)

and Δ_{e} was related to Δ_{i} (usually set equal to 0.99) by Eq. (14a), so

$$\Delta_{e} = \Delta_{i} \sum_{s = ions} (n_{s}/n_{e}) \exp(-k_{0}^{2}a_{s}^{2}/2)$$
(37)

where the sum is over ion species [recall that a loss cone ion distribution, given by Eq. (11), is treated as two Maxwellian ion species].

Numerical solutions were obtained using the local method also. The first two parts of Eq. (16), D = 0, and $\partial D/\partial x = 0$, were solved simultaneously for complex ω_L and x_L , using a Newton-Raphson²¹ root finder (k_L was assumed to be zero, so that $\partial D/\partial k_{y} = 0$ was always satisfied). This root-finder (in contrast to the root-finder used for the nonlocal method) worked well only when the initial values of ω_{I} and x_{I} were reasonably close to the correct values. Initial values were taken either from the results of the nonlocal method, or from previous calculations of $\omega_{I_{L}}$ and $x_{I_{L}}$ using slightly different parameters. Because the root-finder was very quick, and the equations to be solved were relatively simple, it was possible to cover the parameter space fairly densely, using the local method. In fact the entire parameter space could be covered in less time using the local method than it took to solve a single case using the nonlocal method, and the results were more accurate using the local method when inequality (21) was well-satisfied. On the other hand, the nonlocal method was needed to obtain initial values of ω_{L} and x_{L} in a given regime of parameter space, as well as to find modes with $k_{I} \neq 0$, and modes with inequality (21) not satisfied.

Abridged flow charts and complete listings for the program ROOTS and its subroutines (used for the nonlocal method) and of the program LOCAL and its subroutines (used for the local method) are given in Appendix I. ŝ

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V. Results

Runs were made for both methods using a mirror ratio R = 3, a mass ratio $m_i/m_e = 3700$ (appropriate for a deuterium plasma), a density gradient k_0a_i ranging from 0.1 to 6.3 $[a_i \text{ refers to } v_i/\omega_{ci} \text{ where } v_i$ is the thermal velocity of the first Maxwellian component in Eq. (11); the total ion thermal velocity is thus $v_i(1+R^{-1})^{1/2}$], and density $\omega_{pi}^2/\omega_{ci}^2$ ranging from 3 to 10^4 . In each case, for the mode with the highest growth rate, ω and k_y were found and $\phi(x,y)$ was plotted. Sample results using the nonlocal method for one case, $\omega_{pi}^2/\omega_{ci}^2 = 30$, $k_0a_i = 1.13$, and $p_{max} = 3$, are shown in Fig. 2, a plot of ω vs. k_y , (note that there are $p_{max}-p_{min}$ +1 = 7 brances for each cyclotron harmonic in Fig. 2, as discussed in Sec. IIB.)and Fig. 3, a contour plot of $\phi(x,y)$. When inequality (21) was satisfied and when p_{max} was high enough so that $\phi_p \approx 0$ for $|p| \geq p_{max}$, the results using each of the two methods were found to be in excellent agreement, within 1% for ω, k_y , and x_L , and within 10% for the half-width of $\phi(x)$. (see Appendix J)

Figure 4 shows the growth rate of the fastest growing mode as a function of $\omega_{pi}^2/\omega_{ci}^2$ and k_0a_i , using either the local or the nonlocal results, whichever was most accurate for each set of parameters. Figure 5 shows the minimum p_{max} needed in order to use the nonlocal method, as a function of $\omega_{pi}^2/\omega_{ci}^2$ and k_0a_i . Figure 6 shows $|\phi(x)|^2$ and $|\phi_p|^2$ for various values of k_0a_i , for $\omega_{pi}^2/\omega_{ci}^2 = 1000$.

VI. Intepretation

The parameter space is divided into several regimes in each of which the fastest growing mode has qualitatively different characteristics (shown in Fig. 7) and is driven unstable by a different mechanism.

At $k_{0a}^{a} < 1.25$ (region A in Fig. 7), the most unstable mode is a drift cone mode. Its characteristics are in qualitative agreement with





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Figure 3. Contour plot of the real part of $\phi(x,y)$ for the fastest growing mode in Fig. 2, at $\omega/\omega_{ci} = 0.953 + 0.873i$, $k_y a_i = 3.55$. The plasma density is at a maximum at $k_o x = 0$ and 2π , and at a minimum at $k_o x = \pi$. Re $\phi(x,y)$ is normalized so that it ranges from -1 to +1. Contours are shown for 0.8, 0.6, 0.4, 0.2, and 0 (solid curves) and for -0.2, -0.4, -0.6, and -0.8. The figure would have to be compressed vertically by a factor of 2.56 to make the vertical and horizontal scales the same.



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Figure 5. The minimum p_{max} needed to find the fastest growing mode using the nonlocal method is shown as a function of $(\omega_{pi}/\omega_{ci})^2$ and $k_0 a_i$, for $\Delta_i = 0.99$, R = 3, $m_i/m_e = 3700$.



Figure 6. $|\phi(\mathbf{x})|^2$ and $|\phi_p|^2$ are shown for the fastest growing mode for various values of $k_{o}a_i$, using $(\omega_{pi}/\omega_{ci})^2 = 1000$, $\Delta_i = 0.99$, R = 3, and $m_i/m_e = 3700$. The plasma guiding center density is at a maximum at $k_o \mathbf{x} = 0$ and 2π , and at a minimum at $k_o \mathbf{x} = \pi$. For $k_o a_i > 1.80$, the plasma density (as opposed to guiding center density) has its maximum at $k_o \mathbf{x} = \pi$, and its minimum at $k_o \mathbf{x} = 0$ and 2π . The twelve cases shown are located respectively in regions A, A, B, B, B, C, D, E, E, F, F, and G of Fig. 7.



Figure 7. Regions of the parameter space are shown in which the fastest growing mode has different characteristics. The usual drift cone mode occurs in region A; other types of instabilities occur in the other regions, discussed in Sec. VI. R = 3, $\Delta_i = 0.99$, $m_i/m_e = 3700$.

the mode described by Post and Rosenbluth¹⁴ for an infinite medium with constant density gradient; it is modified by the fact that $\phi(x)$ is centered near the edge of the plasma (typically $k_0 x_L/2\pi \approx 0.6$) where the density gradient $(a_i/n)dn/dx$ (a small parameter in Ref. 14) is relatively large. For $k_0 a_i \geq 0.4$, the ion density gradient begins to compete with the loss cone as a source of free energy for the instability, and the mode has some of the characteristics of the Mikhailovskii drift cyclotron instability¹⁵.

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At $k_0 a_i > 1.25$ and $\omega_{pi}^2 / \omega_{ci}^2 \ge 50$ (region B in Fig. 7), a new mode appears, and quickly becomes more unstable than the drift cone mode. This mode has $\phi(x)$ centered at $k_0 x/2\pi = 0.5$ (the point of minimum ion guiding center density), and has a purely imaginary frequency. It is an ion two-stream instability, driven by the double-humped ion velocity distribution $f_{0i}(x,y)$ at $k_{0x}/2\pi = 0.5$ [since $n_i(x)$ is double-humped, and $f_{0i}(x,y) = g_i(v_1,v_{\parallel})n_i(x_{gc})$]. It is thus an artifact of the periodic guiding center density profile $n_i(x)$, and would not appear in a plasma with a single-humped density profile.

The ion-two-stream instability is most unstable at $k_0a_i \approx 1.35$, and disappears at $k_0a_i > 1.80$. The drift cone (or drift cyclotron) instability also disappears at about this value of k_0a_i , since $\Delta_e=0$ at $k_0a_i=1.80$ for R=3 [in order to satisfy Eq.(14b)], and the drift cone mode depends on an electron density gradient. But for $\omega_{pi}^2/\omega_{ci}^2 \ge 200$, another instability appears at $k_0a_{i} \ge 1.70$, and remains the most unstable mode until $k_0a_i \approx 2.5$ (region C of Fig. 7). This mode has $\phi(x)$ centered at $k_0x/2\pi = 0$, the point of maximum ion guiding center density, and is a purely growing ion instability, with $k_L \ge k_y$; the maximum growth rate occurs when $k_y = 0$. The instability arises because the projection of the ion

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velocity distribution in the x-direction $\int f_{01}(x,y) dv_y dv_z$, has its loss cone deepened slightly when $d^2n_1/dx^2 < 0$ (true at x = 0), and this makes the ion distribution unstable by the Penrose criterion²² to modes with wave vector in the x-direction. Eq. (16) can be solved analytically in the limits $k_y = 0$, Re $\omega_L = 0$, $k_L v_i \ge |\omega_L| \ge \omega_{ci}$, which are applicable in region C when $\omega_{pi} \gg \omega_{ci}$. In these limits, $D(x,k_x,\omega)$ can be written in terms of the plasma dispersion function $Z(\zeta)$, using Eqs. (11), (23), (25), and (35), and $Z(\zeta)$ can be approximated by $i\pi^{1/2}$, since $\zeta = \omega_L/k_L v_i < 1$.²⁰ Then Eq. (16) has the solutions (see Appendix K)

$$\omega_{\rm L} = 0.307 \text{ i } \omega_{\ell h} (1 - R^{-1})^{-1/2} (e_2 - e_1)^{3/2} [R^{1/2} (e_2 + 1) - (e_1 + 1)]^{-1}$$

$$k_{\rm L} = 0.577 (\omega_{\ell h} / v_1) (1 - R^{-1})^{-1/2} (e_2 - e_1)^{1/2}$$

$$x_{\rm L} = 0$$
(38)

where

and

 $e_1 \equiv \exp(-k_0^2 a_1^2/2), e_2 \equiv \exp(-k_0^2 a_1^2/2R),$

$$\omega_{\text{lh}} \equiv \omega_{\text{pi}} [1 + \omega_{\text{pe}}^2 (1 + \Delta_{\text{e}})/\omega_{\text{ce}}^2]^{-1/2}$$

is the lower hybrid frequency at x = 0,

and $\Delta_{e} = (Re_{1}-e_{2})/(R-1)$ from Eq. (14a) assuming $\Delta_{i} = 1$. The growth rate is greatest for a given ω_{lh} when $k_{0}a_{i} = [2R \ln R/(R-1)]^{1/2}$. For R = 3, this corresponds to $k_{0}a_{i} = 1.80$, and Eq. (38) gives $\omega_{L} = 0.06 \ i \ \omega_{lh}$ and $k_{L}/k_{0} = 0.27 \ \omega_{lh}/\omega_{ci}$. Since Eq. (38) is derived using the local method and depends on inequality (21) for validity, we might expect that this mode will no longer exist when $\omega_{pi}^{2}/\omega_{ci}^{2} \leq 200$, since then $k_{L}/k_{0} \leq 2$. The nonlocal method confirms this expectation and gives results in good agreement with Eq. (38) when $\omega_{pi}^{2}/\omega_{ci}^{2} \geq 200$

Since the mechanism for this instability is the deepening of the loss cone due to $d^2n/dx^2 < 0$, it does not depend on the periodic nature of the density profile, and we would expect it to occur even in an isolated slab of plasma. However, it would not occur in a plasma column with cylindrical symmetry; this may be seen by applying the Penrose criterion²² at x = 0 to a plasma with distribution function $f_{0i}(x,y) = g_i(v_1,v_{\parallel},x_{gc})$ given by Eq. (11), and $n_i(x) = n_0 \exp(-x^2/L - y^2/L)$. Hence the instability of region C, like that of region B, is an artifact of the model density profile used. (see Appendix L)

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When $k_0 a_i > 1.80$, $\Delta_e < 0$ (for R = 3), i.e. the electron density gradient is in the opposite direction of the ion guiding center density gradient, and a drift cone or drift cyclotron instability occurs with $\phi(x)$ centered at $0 < k_0 x/2\pi < 0.5$, rather than at $0.5 < k_0 x/2\pi < 1.0$, as occurs when $\Delta_e > 0$. This instability has higher growth rate than any of the purely growing ion instabilities when $2.5 \ge k_0 a_i \ge 4.0$ (regions D and E). At $k_0 a_i \approx 2.5$ (region D) finite $k_L \sim k_y$ must be used in Eq. (16) to describe the fastest growing mode by the local method, but at $k_0 a_i > 2.5$ (region E) $k_L = 0$ may be used. The ion density gradient is more important than the loss cone in driving this instability, but the loss cone is crucial in determining the electron density gradient Δ_e (which would be positive and much smaller in absolute value if there were no loss cone), on which the real part of the frequency (and indirectly the growth rate) depends.

For $k_0 a_i \ge 4.0$, the electron density gradient is unimportant because Δ_e (which decreases exponentially with $k_0 a_i$) is negligibly small. Ion instabilities of the two-stream (or many-stream) type again become important (for $\omega_{pi}^2/\omega_{ci}^2 \ge 50$), driven by the ion density profile [which for $k_0 a_i >> 1$ results in $f_{0i}(v_y)$ consisting of several

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streams], and by the loss cone. These instabilities are purely growing at higher densities and lower k_0a_i (region F). At lower densities and higher k_0a_i , when ion cyclotron effects (i.e. $\operatorname{Im} \omega \leq \omega_{ci}$) become important, the real frequencies are not zero (region G). Because $f_{0s}(x, y)$ is not very dependent on x when $k_0a_i >> 1$, these modes tend to have $\phi(x)$ very spread out in x, with $k_y = k_x \approx k_0$. Hence they cannot be described by the local method, only by the nonlocal method.

Fastest-growing modes which cannot be described by the local method also exist at lower $k_{0}a_{i}$, at sufficiently low densities (regions H,I,J,K). For regions H,I, and J, Re $\omega \neq 0$; in region K, Re $\omega = 0$. In regions I and K, like regions G and F, the instability is due entirely to the ions, and the electron density gradient (which is very small in these regions) plays no role. In regions H and J, the electron density gradient is important, and both the ions and electrons contribute to the mode. This may be seen by comparing the perturbed ion and electron densities for the fastest-growing mode in the different regions. In regions F, G, I, and K, the perturbed ion density is much greater, while in regions H and J, the perturbed ion and electron densities are comparable. VII. Summary and Conclusions

Two methods were derived for finding the linear normal modes of electrostatic perturbations in a non-uniform Vlasov plasma. The methods complement each other, since the local method, using Eqs. (16) and (22) is valid only when the wavelength is much smaller than the scale length, while the nonlocal method, using Eqs. (27), (28), and (29), is most efficient when the wavelength is comparable to the scale length. Neither method depends on the Larmor radius being small compared to the scale length. Both methods were employed (each in its own regime of validity)

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to find the normal modes for the drift cone and related instabilities in a loss cone plasma with density varying sinusoidally in space in a direction perpendicular to a uniform magnetic field, over a wide range of densities and density gradients.

An immediate application of these results is to computer simulations of the drift cone mode using periodic boundary conditions and a slab geometry, with sinusoidal density profile. Analytic solutions for the linear normal modes and frequencies for this model are useful in order to check the linear behavior of the simulation, before we explore the nonlinear behavior. Such simulations should have $k_0^{a_1} < 1.25$; otherwise we will introduce new instabilities that depend on the model, and would not occur in a cylindrical plasma column, for example, as discussed in Sec. VI.

If other models (including more realistic models of experiments) behave in a similar fashion to the model we have considered, then no qualitatively new behavior (not predicted by the small Larmor radius approximation) would occur if $(a_i/n) dn/dx < 1$. The condition is satisfied for the bulk of the plasma in the 2XIIB experiment.³

More generally, we have shown that it is possible to use a local method to find the normal modes and their growth rates, even when the Larmor radius is comparable to the scale length, provided the wavelength is much less than the scale length [and assuming we can find the equilibrium distribution $f_{0s}(x,y)$]. When the wavelength is comparable to the scale length, it is possible to use a matrix equation (not tri-diagonal in general) truncated at fairly low index p. Although the model we have solved does not resemble any experiment, it should be possible to use these methods to find the normal modes of experimental plasmas with Larmor radius comparable to scale length.

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Historical Note

This report is the culmination of many years of work.

In 1969, Birdsall and Fuss (at Lawrence Livermore Lab.) began computer simulation of the drift-cyclotron loss-cone mode, using particles on a two-dimensional periodic mesh, starting with a sinusoidal density profile [Eq.(13) with $\Delta_i \approx 1$, $k_0 a_i/2\pi \approx 0.1$, 0.2]. The checks between simulation and small-amplitude dispersion results (ω , k), calculated by N. Lindgren (Univ. of Calif., Berkeley) from existing local approximation theory¹⁴, were fair, considering the difference in the model and theory.

Langdon (Univ. of Calif., Berkeley) then provided the exact or nonlocal theory, essentially Eq.(29), from which Birdsall and Fuss obtained a few sample solutions for ω , k (showing the multiple roots, as in Fig. 2) and ϕ (x,y) (showing the localization, as in Fig. 3), and these only for a ring distribution of velocities,

$$g_{i!}(v_{i}, x) = (2\pi v_{o})^{-1} \delta(v_{i}-v_{o}) n_{i}(x),$$

rather than the loss cone distribution given by Eq.(11). While these results were presented in talks^{4,5} and progress reports, they were incomplete.

Since 1973, Gerver derived the local method [eqs.(16) and (20)], and used these and the nonlocal dispersion relation, Eq.(29), with loss cone distributions of the form given by Eq.(11), to explore the broad range of parameters displayed in Fig. 4 through 7. This required considerable refinement of the numerical techniques used earlier, including ⁶ the use of straight-line orbits [Eq.(35)], and extensive modifications of the root-finding routine. The physical identification of the various modes and the interpretations are also Gerver's.

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Acknowledgments

We wish to express our appreciation to A.N. Kaufman and W.B. Kunkel for helpful discussions and suggestions, and to Richard Meyers for help in debugging the computer programs.
Derivation of Eq. (4):

Starting with Eq. (3) and the identities

$$\phi(\underline{x}_{s}',t) = \int d\underline{k} \ \tilde{\phi}(k) \exp(i\underline{k} \cdot \underline{x}_{s}' - i\omega t)$$

$$f_{0s}(\underline{x}_{s}',\underline{v}_{s}') = \int d\underline{k}'' \ \tilde{f}_{s}(\underline{k}'',\underline{v}_{s}') \exp(-i\underline{k}'' \cdot \underline{x}_{s}')$$

we obtain

$$\int d\mathbf{k}'' \int d\mathbf{k} \left[\sum_{\mathbf{s}} \frac{4\pi q_{\mathbf{s}}^2}{m_{\mathbf{s}}} \int d\mathbf{y} \int_{-\infty}^{0} d\tau \ \mathbf{i} \mathbf{k} \cdot \frac{\partial \tilde{\mathbf{f}}_{\mathbf{s}}}{\partial \mathbf{y}_{\mathbf{s}}'} \left(\mathbf{k}'', \mathbf{y}_{\mathbf{s}}' \right) \right]$$

$$\tilde{\phi}(\mathbf{k}) \exp(\mathbf{i} \mathbf{k} \cdot \mathbf{x}_{\mathbf{s}}' - \mathbf{i} \mathbf{k}'' \cdot \mathbf{x}_{\mathbf{s}}' - \mathbf{i} \omega \mathbf{t}) - \mathbf{k}^2 \ \tilde{\phi}(\mathbf{k}) = 0$$

$$\int d\mathbf{k} \ \tilde{\phi}(\mathbf{k}) \left(\int d\mathbf{k}' \exp(-\mathbf{i} \mathbf{k}'' \cdot \mathbf{x} + \mathbf{i} \mathbf{k} \cdot \mathbf{x}) \left\{ \sum_{\mathbf{s}} \frac{4\pi q_{\mathbf{s}}^2}{m_{\mathbf{s}}} \right\}$$

$$\int d\mathbf{y} \ \int_{-\infty}^{0} d\tau \ \mathbf{i} \mathbf{k} \cdot \frac{\partial \tilde{\mathbf{f}}_{\mathbf{s}}}{\partial \mathbf{y}_{\mathbf{s}}} \left(\mathbf{k}'', \mathbf{y}_{\mathbf{s}}' \right) \exp[\mathbf{i} (\mathbf{k}'' - \mathbf{k}) \cdot (\mathbf{x} - \mathbf{x}_{\mathbf{s}}') - \mathbf{i} \omega \tau]$$

$$-\mathbf{k}^2 = 0$$

Let $k' \equiv k - k''$, and we obtain Eq. (4).

Derivation of Eq. (7):

Starting with the identity

$$\tilde{f}_{s}(\underline{k}, \underline{v}) = \int d\underline{x} f_{s}(\underline{x}, \underline{v}) \exp(i\underline{k} \cdot \underline{x})$$

$$= \int d\underline{x} g_{s}(\underline{v}_{\perp}, \underline{v}_{\parallel}, \underline{x}_{gc}) \exp(i\underline{k} \cdot \underline{x})$$

$$= \int d\underline{x}_{gc} g_{s}(\underline{v}_{\perp}, \underline{v}_{\parallel}, \underline{x}_{gc}) \exp(i\underline{k} \cdot \underline{x}_{gc})$$

$$\exp[-i\underline{k} \cdot (\underline{v} \times \hat{B}_{0}) \omega_{cs}^{-1}]$$

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$$= \exp[-i\underline{k} \cdot (\underline{v} \times \hat{B}_{0}) \omega_{cs}^{-1}] \tilde{g}_{s} (v_{\perp}, v_{\parallel}, \underline{k})$$

using $x_{gc} = x + v \times \hat{B}_{0} \omega_{cs}^{-1}$

Derivation of Eq. (9):

We put Eq. (7) into Eq. (6) to obtain $D_{s}(\underline{x},\underline{k},\omega) = \frac{4\pi q_{s}^{2}}{m_{s}} \exp(-i\underline{k} \cdot \underline{x}) \int d\underline{k}' \exp(i\underline{k}' \cdot \underline{x}) \cdot (\underline{k}' \cdot \underline{k}') \cdot (\underline{k}' \cdot \underline{k}' \cdot \underline{k}' \cdot \underline{k}') \cdot (\underline{k}' \cdot \underline{k}' \cdot \underline{k}' \cdot \underline{k}') \cdot (\underline{k}' \cdot \underline{k}' \cdot \underline{k}' \cdot \underline{k}' \cdot \underline{k}') \cdot (\underline{k}' \cdot \underline{k}' \cdot \underline{k}' \cdot \underline{k}' \cdot \underline{k}' \cdot \underline{k}') \cdot (\underline{k}' \cdot \underline{k}' \cdot \underline$

We evaluate $\frac{\partial}{\partial v_s}$ { }

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$$\frac{\partial}{\partial \underline{v}_{s}} \{ \} = \exp[-i(\underline{k} - \underline{k}') \cdot (\underline{v} \times \hat{B}_{0})\omega_{cs}^{-1}]$$

$$\{\frac{\partial \tilde{g}_{s}}{\partial \underline{v}_{s}} - \frac{i\tilde{g}_{s}}{\omega_{cs}} \frac{\partial}{\partial \underline{v}_{s}} [(\underline{k} - \underline{k}') \cdot (\underline{v}_{s}' \times \hat{B}_{0}]\}$$
(A2)

where the arguments of \tilde{g}_s and its derivatives are always taken to be $(v_1, v_{\parallel}, k - k')$.

$$\frac{\partial \tilde{g}_{s}}{\partial v_{s}'} = \hat{v}_{\perp s} \frac{\partial \tilde{g}_{s}}{\partial v_{\perp}} + \hat{B}_{0} \frac{\partial \tilde{g}_{s}}{\partial v_{\parallel}}$$
(A3)

$$\frac{\partial}{\partial \mathbf{v}_{s}'} [(\mathbf{k} - \mathbf{k}') \cdot (\mathbf{v}_{s}' \times \hat{\mathbf{B}}_{0})] = \frac{\partial}{\partial \mathbf{v}_{s}'} \{\mathbf{v}_{s}' \cdot [\hat{\mathbf{B}}_{0} \times (\mathbf{k} - \mathbf{k}')]\}$$
$$= \hat{\mathbf{B}}_{0} \times (\mathbf{k} - \mathbf{k}')$$
(A4)

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where $\hat{\mathbf{v}}_{\underline{\mathbf{s}}}' \equiv \underline{\mathbf{v}}_{\underline{\mathbf{s}}}' / |\underline{\mathbf{v}}_{\underline{\mathbf{s}}}'|$, and $\underline{\mathbf{v}}_{\underline{\mathbf{s}}}' \equiv \underline{\mathbf{v}}_{\underline{\mathbf{s}}}' - (\underline{\mathbf{v}}_{\underline{\mathbf{s}}}' \cdot \hat{\underline{\mathbf{B}}}_{0})\hat{\underline{\mathbf{B}}}_{0}$

Putting Eqs. (A2), (A3) and (A4) into Eq. (A1), we find

$$D_{s}(\underline{x},\underline{k},\omega) = \frac{4\pi q_{s}^{2}}{m_{s}} \exp(-i\underline{k} \cdot \underline{x}) \int d\underline{k}' \exp(i\underline{k}' \cdot \underline{x})$$

$$d\underline{y} \int_{-\infty}^{0} d\tau \exp[-i(\underline{k} - \underline{k}') \cdot (\underline{y}_{s}' \times \hat{B}_{0})\omega_{cs}^{-1}]$$

$$\{i\underline{k} \cdot \hat{v}_{\underline{i},s}' \frac{\partial \tilde{g}_{s}}{\partial v_{\underline{i}}} + \underline{k} \cdot [\hat{B}_{0} \times (\underline{k} - \underline{k}')]\tilde{g}_{s} \omega_{cs}^{-1}$$

$$+ i k_{\parallel} \frac{\partial \tilde{g}_{s}}{\partial v_{\underline{i}}}\} \exp[i\underline{k}' \cdot (\underline{x}_{s}' - \underline{x}) - i\omega\tau] \qquad (A5)$$

We can use the fact that $\underline{k} \cdot (\hat{B}_0 \times \underline{k}) = 0$ to replace $\underline{k} \cdot [\hat{B}_0 \times (\underline{k} - \underline{k}')]$ by $\underline{k} \cdot (\hat{B}_0 \times \underline{k}')$ in Eq. (A5). We then put in the expression for the particle orbits, Eq. (8), and obtain

$$D_{s}(\mathbf{x},\mathbf{k},\omega) = \frac{4\pi q_{s}^{2}}{m_{s}} \exp(-i\mathbf{k} \cdot \mathbf{x}) \int d\mathbf{k}' \exp(i\mathbf{k}' \cdot \mathbf{x})$$

$$\int d\mathbf{v} \exp[i(\mathbf{k}' - \mathbf{k}) \cdot (\mathbf{v} \times \hat{\mathbf{B}}_{0})\omega_{cs}^{-1}] \int_{-\infty}^{0} d\tau$$

$$\exp[i\mathbf{k} \cdot \mathbf{v}_{\perp} \omega_{cs}^{-1} \sin \omega_{cs}\tau + i\mathbf{k} \cdot (\mathbf{v} \times \hat{\mathbf{B}}_{0}) \omega_{cs}^{-1}$$

$$(1 - \cos \omega_{cs}\tau) + i\mathbf{k}_{\parallel}\mathbf{v}_{\parallel} - i\omega\tau]$$

$$\{[i\mathbf{k} \cdot \hat{\mathbf{v}}_{\perp} \cos \omega_{cs}\tau + i\mathbf{k} \cdot (\hat{\mathbf{v}}_{\perp} \times \hat{\mathbf{B}}_{0}) \sin \omega_{cs}\tau] \frac{\partial \tilde{\mathbf{g}}_{s}}{\partial \mathbf{v}_{\perp}}$$

$$- \mathbf{k} \cdot (\hat{\mathbf{B}}_{0} \times \mathbf{k}') \omega_{cs}^{-1} \tilde{\mathbf{g}}_{s} + i\mathbf{k}_{\parallel} \frac{\partial \tilde{\mathbf{g}}_{s}}{\partial \mathbf{v}_{\parallel}} \}$$

$$= \frac{4\pi q_{s}^{2}}{m_{s}} \exp(-i\mathbf{k} \cdot \mathbf{x}) \int d\mathbf{k}' \exp(i\mathbf{k}' \cdot \mathbf{x}) \int d\mathbf{v}$$

$$\exp[i\mathbf{k}' \cdot (\mathbf{v} \times \hat{\mathbf{B}}_{0}) \omega_{cs}^{-1}] \{[\mathbf{k} \cdot (\mathbf{k}' \times \hat{\mathbf{B}}_{0}) \omega_{cs}^{-1} \tilde{\mathbf{g}}_{s}$$

$$+ i\mathbf{k}_{\parallel} \frac{\partial \tilde{\mathbf{g}}_{s}}{\partial \mathbf{v}_{\parallel}}]\mathbf{I}_{1} + i \frac{\partial \tilde{\mathbf{g}}_{s}}{\partial \mathbf{v}_{\perp}} \mathbf{I}_{2}\}$$
(A6)

where
$$I_1 \equiv \int_{-\infty}^{0} d\tau \exp[i\mathbf{k} \cdot \mathbf{y}_{\perp} \omega_{cs}^{-1} \sin \omega_{cs} \tau]$$

$$\frac{-i\mathbf{k} \cdot (\mathbf{y} \times \mathbf{B}_0) \omega_{cs}^{-1} \cos \omega_{cs} \tau + i\mathbf{k}_{\parallel} \mathbf{v}_{\parallel} \tau - i\omega \tau]}{I_2 \equiv \int_{-\infty}^{0} d\tau \exp[i\mathbf{k} \cdot \mathbf{y}_{\perp} \omega_{cs}^{-1} \sin \omega_{cs} \tau - i\mathbf{k} \cdot (\mathbf{y} \times \hat{\mathbf{B}}_0) \omega_{cs}^{-1}]}$$

$$\cos \omega_{cs} \tau + i\mathbf{k}_{\parallel} \mathbf{v}_{\parallel} \tau - i\omega \tau] [\mathbf{k} \cdot \hat{\mathbf{v}}_{\perp} \cos \omega_{cs} \tau + \mathbf{k} \cdot (\hat{\mathbf{v}}_{\perp} \times \hat{\mathbf{B}}_0) \sin \omega_{cs} \tau]$$

We define an azimuthal angle ϕ by

$$\underbrace{\mathbf{k}}_{\mathbf{v}} \cdot \underbrace{\mathbf{v}}_{\mathbf{I}} = \mathbf{k}_{\mathbf{I}} \mathbf{v}_{\mathbf{I}} \cos \phi$$
$$-\underbrace{\mathbf{k}}_{\mathbf{v}} \cdot (\underbrace{\mathbf{v}}_{\mathbf{x}} \times \widehat{\mathbf{B}}_{\mathbf{0}}) = \underbrace{\mathbf{k}}_{\mathbf{I}} \mathbf{v}_{\mathbf{I}} \sin \phi$$

Then

$$I_{1} = \int_{-\infty}^{0} d\tau \exp[ik_{\perp}v_{\perp} \omega_{cs}^{-1} \sin(\phi + \omega_{cs}\tau) + ik_{\parallel}v_{\parallel}\tau - i\omega\tau]$$

$$I_{2} = \int_{-\infty}^{0} d\tau \exp[ik_{\perp}v_{\perp} \omega_{cs}^{-1} \sin(\phi + \omega_{cs}\tau) + ik_{\parallel}v_{\parallel}\tau - i\omega\tau]$$

$$k_{\perp} \cos(\phi + \omega_{cs}\tau)$$

These integrals can be evaluated by using the Bessel function identity

$$\sum_{\ell} J_{\ell}(\mathbf{x}) \exp(i\ell\theta) = \exp(i \mathbf{x} \sin\theta)$$

$$I_{1} = \sum_{\ell} J_{\ell}(\mathbf{k}_{\perp}\mathbf{v}_{\perp}/\omega_{cs}) \int_{-\infty}^{0} d\tau \exp(i\ell\phi + i\ell\omega_{cs} - i\omega\tau + i\mathbf{k}_{\parallel}\mathbf{v}_{\parallel}\tau]$$

$$= i \sum_{\ell} J_{\ell}(\mathbf{k}_{\perp}\mathbf{v}_{\perp}/\omega_{cs}) \exp(i\ell\phi) (\omega - \ell\omega_{cs} - \mathbf{k}_{\parallel}\mathbf{v}_{\parallel})^{-1}$$
(A8)

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$$I_{2} = (k_{\perp}/2) \sum_{\ell} J_{\ell} (k_{\perp} v_{\perp} / \omega_{cs}) \int_{-\infty}^{0} d\tau \sum_{\pm} \exp[i(\ell \pm 1)\phi]$$

$$+ i(\ell \pm 1)\omega_{cs}\tau - i\omega\tau + ik_{\parallel} v_{\parallel}\tau]$$

$$= (k_{\perp}/2) \sum_{\ell} \sum_{\pm} J_{\ell\pm1} (k_{\perp} v_{\perp} / \omega_{cs}) \int_{-\infty}^{0} d\tau \exp[i\ell\phi]$$

$$+ i\ell\omega_{cs}\tau - i\omega\tau + ik_{\parallel} v_{\parallel}\tau]$$

$$= i(k_{\perp}/2) \sum_{\ell} \sum_{\pm} J_{\ell\pm1} (k_{\perp} v_{\perp} / \omega_{cs}) \exp(i\ell\phi)$$

$$(\omega - \ell\omega_{cs} - k_{\parallel} v_{\parallel})^{-1}$$

$$= i(\omega_{cs} / v_{\perp}) \sum_{\ell} \ell J_{\ell} (k_{\perp} v_{\perp} / \omega_{cs}) \exp(i\ell\phi) (\omega - \ell\omega_{cs} - k_{\parallel} v_{\parallel})^{-1}$$
(A9)

where we have used the Bessel function identity

$$(2\ell/x)J_{\ell}(x) = J_{\ell+1}(x) + J_{\ell-1}(x)$$

The dy in Eq. (A6) can be replaced by $v_1 dv_1 d\phi$, and the integration over ϕ performed. We define α as the angle between k_1 ' and k_1

$$\exp(i\alpha) = \hat{k}_{\underline{i}} \cdot (\hat{k}_{\underline{i}}' + i \hat{B}_{0} \times \hat{k}_{\underline{i}}')$$
(A10)

Then

$$-\underline{k}' \cdot (\underline{v} \times \hat{B}_{0}) = \underline{k}_{\underline{i}}' \underline{v}_{\underline{i}} \sin(\phi - \alpha)$$

and we can use the identity (A7) to evaluate

$$\exp[i\underline{k}' \cdot (\underline{\mathbf{y}} \times \underline{\mathbf{B}}_{0}) \omega_{cs}^{-1}] = \exp[-i\underline{\mathbf{k}}_{\mathbf{L}}' \underline{\mathbf{v}}_{\mathbf{L}} \omega_{cs}^{-1} \sin(\phi - \alpha)]$$

$$= \sum_{m} J_{m}(\underline{\mathbf{k}}_{\mathbf{L}}' \underline{\mathbf{v}}_{\mathbf{L}} / \omega_{cs}) \exp[-i\underline{\mathbf{m}}(\phi - \alpha)] \qquad (A11)$$

The ϕ dependence of I_1 and I_2 , given by Eqs. (A8) and (A9), appears only in the factor $\exp(i\ell\phi)$. The only other ϕ dependence of the integrand in Eq. (A6) is in the term evaluated in Eq.(A11). So the integration over ϕ that must be done in Eq. (A6) is

$$\int d\phi \, \exp[-im(\phi\alpha)] \, \exp(i\ell\phi) = 2\pi \, \delta_{\ell m} \, \exp(i\ell\alpha) \tag{A12}$$

Putting Eqs. (A8) through (A12) into Eq. (A6), we obtain Eq. (9).

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Derivation of Eq. (10):

In Eq. (9), $\tilde{g}_{g}(v_{1},v_{\parallel},k-k')$ is negligible for $|k-k'| \gg L^{-1}$, where L is the scale length, and for $v_{1} \gg v_{s}$, where v_{s} is the thermal velocity. If $a_{s} \ll L$, where $a_{s} \equiv v_{s}/\omega_{cs}$ is the Larmor radius, then $\tilde{g}_{s}(v_{1},v_{\parallel},k-k')$ is negligible unless $(k-k') \cdot (v \times \tilde{B}_{0})\omega_{cs}^{-1} \cdots 1$, so we can make the approximation

$$\exp[-i(\underline{k} - \underline{k}') \cdot (\underline{v} \times \hat{B}_0) \omega_{cs}^{-1}] \approx 1$$
(B1)

in Eq. (A5) without changing the integral very much. We use the definition of \tilde{g}_s after Eq. (7) to obtain the identities

$$\int d\mathbf{k}' \exp[i(\mathbf{k}' - \mathbf{k}) \cdot \mathbf{x}] \; \partial \tilde{g}_{s}(\mathbf{k} - \mathbf{k}') / \partial \mathbf{v}_{\parallel}$$

$$= \; \partial g_{s}(\mathbf{x}) / \partial \mathbf{v}_{\parallel}$$

$$\int d\mathbf{k}' \; \exp[i(\mathbf{k}' - \mathbf{k}) \cdot \mathbf{x}] \partial \tilde{g}_{s}(\mathbf{k} - \mathbf{k}') / \partial \mathbf{v}_{t}$$

$$= \; \partial g_{s}(\mathbf{x}) / \partial \mathbf{v}_{t}$$

$$\int d\mathbf{k}' \; [\hat{B}_{0} \; \mathbf{x} \; i(\mathbf{k}' - \mathbf{k})] \; \exp[i(\mathbf{k}' - \mathbf{k}) \cdot \mathbf{x}] \tilde{g}_{s}(\mathbf{k} - \mathbf{k}') = \; \hat{B}_{0} \; \mathbf{x} \; \nabla g_{s}(\mathbf{x})$$

which can be used together with Eq. (B1) to perform the integration over \underline{k} ' in Eq. (A5), yielding

$$D_{s}(\underline{x},\underline{k},\omega) = (4\pi q_{s}^{2}/m_{s}) \exp(-i\underline{k} \cdot \underline{x}) \int dv \int_{-\infty}^{0} d\tau$$

$$\exp[i\underline{k} \cdot (\underline{x}_{s}' - \underline{x}) - i\omega\tau] [i\underline{k} \cdot \hat{v}_{\underline{i}s}' (\partial g_{s}/\partial v_{\underline{i}})$$

$$+ i\underline{k} \cdot (\hat{B}_{0} \times \nabla g_{s}) \omega_{cs}^{-1} + i\underline{k}_{\parallel} (\partial g_{s}/\partial v_{\parallel})] \qquad (B2)$$

Eq. (8) can be put into Eq. (B2), and the integration over τ and ϕ performed just as in Appendix A, yielding Eq. (10). The factor $\exp[i\underline{k}' \cdot (\underline{v} \times \hat{B}_0) \omega_{cs}^{-1}]$ which appears in Eq. (A6) must be replaced by $\exp[i\underline{k} \cdot (\underline{v} \times \hat{B}_0) \omega_{cs}^{-1}]$, which is justified by Eq. (B1); otherwise the derivation is completely analogous. Derivation of Eq. (12):

For a Maxwellian g_{s1}

$$g_{s!}(v_{1},x) = (2\pi v_{s}^{2})^{-1} \exp(-v_{1}^{2}/2v_{s}^{2})n_{s}(x)$$

and a slab $n_{s}(x)$ independent of y and z, we have

$$\int dv_{\parallel} \tilde{g}_{s}(v_{1}, k - k') = (2\pi v_{s}^{2})^{-1} \exp(-v_{1}^{2}/2v_{s}^{2})\tilde{n}_{s}(k_{x}' - k_{x})$$
$$\delta(k_{y} - k_{y}')\delta(k_{z}')$$

where $k_z = 0$ by assumption, \tilde{g}_s is defined after Eq. (7) and \tilde{n}_s is defined after Eq. (12). Then Eq. (9) becomes (after integrating over k_v' , k_z' and v_{\parallel})

$$D_{s}(x,k,\omega) = (4\pi q_{s}^{2}/m_{s}) \int_{-\infty}^{\infty} dk_{x}' \exp[i(k_{x}' - k_{x})x]$$

$$\int_{0}^{\infty} 2\pi v_{1} dv_{1} \sum_{\ell} \exp(i\ell\alpha) J_{\ell}(k_{1}v_{1}/\omega_{cs})J_{\ell}(k_{1}'v_{1}/\omega_{cs})$$

$$(\omega - \ell\omega_{cs})^{-1} (2\pi v_{s}^{2})^{-1} \exp(-v_{1}^{2}/2v_{s}^{2})$$

$$[ik_{y}(k_{x} - k_{x}') + \ell\omega_{cs}/v_{s}^{2}] \tilde{n}_{s}(k_{x}' - k_{x})$$

We define k" $\equiv k_x' - k_x$, change the variable of integration from k_x' to k", and use

$$v_{s}^{-2} \int_{0}^{\infty} v_{!} \exp(-v_{!}^{2}/2v_{s}^{2}) J_{\ell}(kv_{!}/\omega_{cs}) J_{\ell}(k'v_{!}/\omega_{cs}) dv_{!}$$

= $\exp[-(k^{2} + k'^{2}) v_{s}^{2}/2\omega_{cs}^{2}] I_{\ell}(kk' v_{s}^{2}/\omega_{cs}^{2})$

[trom identity 6.633.2 in Gradshteyn and Kyzhik²³]to obtain Eq. (12).

APPENDIX D

Derivation of Eq. (14):

$$\rho_0(\mathbf{x}) \sum_{\mathbf{s}} q_{\mathbf{s}} \int d\mathbf{y} \ f_{0\mathbf{s}}(\mathbf{x}, \mathbf{y}) = 0 \text{ for all } \mathbf{x}$$
(D1)

for no net equilibrium charge density.

$$f_{0s}(x, v) = g_{s}(v_{1}, v_{\parallel}, x_{gc})$$
(D2)

where $\tilde{x}_{gc} \equiv \tilde{x} + \tilde{y} \times \hat{B}_0 \omega_{cs}^{-1}$

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so $x_{gc} = x + v_y / \omega_{cs}$ $y_{gc} = y - v_x / \omega_{cs}$

Using Eq. (D2), the left hand side of Eq. (D1) becomes

$$\rho_{0}(\mathbf{x}) = \sum_{s} q_{s} \int dv_{x} dv_{y} g_{s!}(v_{!}, \mathbf{x}_{gc})$$
$$= \sum_{s} q_{s} (2\pi v_{s}^{2})^{-1} \int dv_{x} dv_{y} \exp(-v_{!}^{2}/2v_{s}^{2}) n_{s}(\mathbf{x}_{gc})$$

for Maxwellian $g_{\underline{i}s}$. For a plasma slab, $n_s(\underline{x})$ independent of y and z,

$$\rho_0(\mathbf{x}) = \sum_{\mathbf{s}} q_{\mathbf{s}} (2\pi v_{\mathbf{s}}^2)^{-1} \int dv_{\mathbf{x}} dv_{\mathbf{y}} \exp(-v_{\underline{t}}^2/2v_{\mathbf{s}}^2) n_{\mathbf{s}} (\mathbf{x} + v_{\mathbf{y}}/\omega_{\mathbf{cs}})$$
(D3)

For a sinusoidal density profile,

$$n_{s}(x) = n_{0s}(1 + \Delta_{s} \cos k_{0} x)$$

Eq. (D3) becomes

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$$\rho_{0}(x) = \sum_{s} q_{s} (2\pi v_{s}^{2})^{-1} \int dv_{x} dv_{y} \exp(-v_{x}^{2}/2v_{s}^{2})$$
$$\exp(-v_{y}^{2}/2v_{s}^{2}) n_{0s} [1 + \Delta_{s} \cos(k_{0}x + k_{0}v_{y}/\omega_{cs})]$$

Integrating over $v_{\mathbf{x}}$

$$\rho_{0}(x) = \sum_{s} n_{0s} q_{s} (2\pi v_{s}^{2})^{-1/2} \int dv_{y} \exp(-v_{y}^{2}/2v_{s}^{2})$$

$$[1 + \Delta_{s} \cos(k_{0}x + k_{0}v_{y}/\omega_{cs})]$$

$$= \sum_{s} n_{0s} q_{s} (2\pi v_{s}^{2})^{-1/2} \int dv_{y} \exp(-v_{y}^{2}/2v_{s}^{2})$$

$$[1 + \Delta_{s} \exp(ik_{0}x)\exp(ik_{0}v_{y}/\omega_{cs})/2$$

$$+ \Delta_{s} \exp(-ik_{0}x)\exp(-ik_{0}v_{y}/\omega_{cs})/2]$$

Since $\rho_0(x)$ vanishes for all x,

$$\int_{0}^{2\pi/k_{0}} \rho_{0}(x) dx = 2\pi k_{0}^{-1} \sum_{s} n_{0s} q_{s} = 0 \text{ or Eq. (14b) and}$$

$$\int_{0}^{2\pi/k_{0}} \rho_{0}(x) \exp(\pm i k_{0} x) dx =$$

$$\pi k_{0}^{-1} \sum_{s} n_{0s} q_{s} \Delta_{s} (2\pi v_{s}^{2})^{-1/2} \int dv_{y} \exp(-v_{y}^{2}/2v_{s}^{2}) \exp(\pm i k_{0} v_{y}/\omega_{cs})$$

$$= \pi k_{0}^{-1} \sum_{s} n_{0s} q_{s} \Delta_{s} \exp(-k_{0}^{2} v_{y}^{2}/2\omega_{cs}^{2}) = 0$$

or Eq. (14a).

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Solution of Eq. (18): $A \frac{\partial^2 \psi}{\partial x^2} + Bx \frac{\partial \psi}{\partial x} + Cx^2 \psi + \lambda \psi = 0$ (E1) Assume $\psi = (\sum_{m=0}^{n} a_m x^m) \exp(-\beta x^2/2)$ (E2) Then $\frac{\partial \psi}{\partial x} = (\sum_{m=0}^{n-1} (m+1)a_{m+1}x^m)\exp(-\beta x^2/2)$ $-(\beta \sum_{m=0}^{n} a_{m} x^{m+1}) \exp(-\beta x^{2}/2)$ $= \left[\sum_{m=0}^{n-1} (m+1)a_{m+1}x^{m} - \beta \sum_{m=1}^{n+1} a_{m-1}x^{m}\right] \exp(-\beta x^{2}/2)$ $= \{\sum_{m=1}^{n+1} [(m+1)a_{m+1} - \beta a_{m-1}] x^{m} + a_{1}\} \exp(-\beta x^{2}/2)$ (E3) $\partial^2 \psi / \partial x^2 = \left(\sum_{m=2}^{n+2} \{ (m+1) [(m+2)a_{m+2} - \beta a_m] \right)$ $-\beta(m a_m - \beta a_{m-2})\} x^m$ + $[2(3a_3 - \beta a_1) - \beta a_1] x + 2a_2 - \beta a_0) \exp(-\beta x^2/2)$ $= \{\sum_{m=2}^{n+2} [(m^2 + 3m + 2)a_{m+2} - \beta(2m + 1)a_m \}$ + $\beta^2 a_{m-2}]x^m + 3(2a_3 - \beta a_1) x + (2a_2 - \beta a_0) \} exp(-\beta x^2/2)$ (E4) $x (\partial \psi / \partial x) = \left[\sum_{m=2}^{n+2} (ma_m - \beta a_{m-2})x^m + a_1x\right] \exp(-\beta x^2/2)$ (E5)

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$$x^{2}\psi = (\sum_{m=2}^{n+2} a_{m-2}x^{m}) \exp(-\beta x^{2}/2)$$
 (E6)

Putting Eqs. (E4), (E5) and (E6) into Eq. (E1) and dividing through by $\exp(-\beta \ x^{2}/2)$ yields

$$\sum_{m=2}^{n+2} x^{m} \{A[m^{2} + 3m + 2)a_{m+2} - \beta(2m + 1)a_{m} + \beta^{2} a_{m-2}] + B[ma_{m} - \beta a_{m-2}] + C a_{m-2} + \lambda a_{m}\} + [3A(2a_{3} - \beta a_{1}) + B a_{1} + \lambda a_{1}] x + [A(2a_{2} - \beta a_{0}) + \lambda a_{0}] = 0$$
(E7)

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This can be satisfied for all x only if each term in the polynomial vanishes. So

$$A(2a_{2} - \beta a_{0}) + \lambda a_{0} = 0$$
(E8)

$$3A(2a_{3} - \beta a_{1}) + Ba_{1} + \lambda a_{1}$$

$$= 6Aa_{3} + (B + \lambda - 3A\beta)a_{1} = 0$$
(E9)

$$A(m^{2} + 3m + 2)a_{m+2} + [Bn - \beta A(2m + 1) + \lambda]a_{m}$$

$$+ (C - \beta B + \beta^{2}A) a_{m-2} = 0 \text{ for } n + 2 \ge m \ge 2$$
(E10)
Setting m = n + 2 in Eq. (E10) yields (since $a_{m} = 0$ for m > n)

$$C - \beta B + \beta^{2}A = 0$$
(E11)

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Then Eq. (E10) becomes

$$A(m+2)(m+1)a_{m+2} + [Bm - \beta A(2m+1) + \lambda]a_m = 0$$
 (E12)

where
$$\beta = (B/2A) + [(B/2A)^2 - C/A]^{1/2}$$
 (E13)

[the + sign is taken in Eq. (E13) so that Re $\beta > 0$ and $\psi(x) \rightarrow 0$ as $x \rightarrow \pm \infty$].

Taking m=n in Eq. (E12) yields

$$Bn - \beta A(2n + 1) + \lambda = 0$$

or

And International

$$\lambda = \beta A(2n + 1) - Bn \tag{E14}$$

Taking m = n - 1 yields $a_{n-1} = 0$, which further implies $a_{n-3} = a_{n-5} = \dots = 0$, so ψ must be either even or odd in x.

For $0 \le m \le n - 2$, Eqs.(E12) and (E14) give the recursion relation for the coefficients a_m

$$\frac{a_{m}}{a_{m+2}} = \frac{A(m+2)(m+1)}{-Bm + \beta A(2m+1) + \lambda}$$
$$= \frac{A(m+2)(m+1)}{B(n-m) - 2\beta A(n-m)}$$
$$= \frac{(m+2)(m+1)}{(n-m)(B/A - 2\beta)}$$
$$a_{m} \qquad (m+2)(m+1)$$

$$\frac{m}{a_{m+2}} = \frac{(m+2)(m+1)}{2(m-n)\eta}$$

(E15)

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where $\eta = \beta - B/2A$

$$= [(B/2A)^{2} - C/A]^{1/2}$$
(E16)

Eq. (E15) is the recursion relation for the Hermite polynomial

$$\sum_{m} a_{m} x^{m} = H_{n}(\eta^{1/2} x)$$
(E17)

So the solutions are

$$ψ_n(x) = H_n(η^{1/2}x) \exp(-\beta x^2/2)$$

 $λ_n = βA(2n + 1) - Bn = (2nη + β)A$

with β and η defined by Eqs. (E13) and (E16).

Q

(E18)

Appendix F

Sketch of Proof that Local Method is Valid When Inequality (21) is Satisfied:

We wish to show that inequality (21)

$$L(k_{L}^{2} + k_{y}^{2})^{1/2} >> 1$$

is the criterion for being able to neglect the terms in Eq.(17) involving higher derivatives of D, so that Eq.(20) is a good approximation to at least some of the solutions of Eq.(15). We will consider a general dispersion function D(x, k_x, ω), not just the D appropriate to the model considered in Sec.III. However, we will not attempt to make our results as general as possible, but will make various physically reasonable assumptions about D(x, k_x, ω) when convenient. Sometimes we will make mathematical assumptions about D(x, k_x, ω) and the physical content of these assumptions will not be obvious; in such cases we cannot prove that the assumptions will be valid for all models of physical interest (or even for the model considered in Sec.III), but we conjecture that the assumptions will be valid for most models of physical interest.

To simplify notation, we will use $\psi(x) \equiv \exp(-ik_L x) \phi(x + x_L)$ and $Q(x,k_x, \omega) \equiv D(x + x_L, k_x + k_L, \omega)$. In this notation, Eq.(15) is

$$Q(x, -i \frac{\partial}{\partial x}, \omega) \psi(x) = 0$$
 (F1)

and Eq.(16) is

$$Q(0,0,\omega_{\rm L}) = \frac{\partial Q}{\partial x} (0,0,\omega_{\rm L}) = \frac{\partial Q}{\partial k_{\rm x}} (0,0,\omega_{\rm L}) = 0$$

Anticipating our result that Eq.(17) is nearly valid without the higher derivatives of D, we expand Q and ψ in perturbation series, and write

$$Q(x, k_{x}, \omega_{n}) = Q^{(0)}(x, k_{x}, \omega_{L}) + \lambda_{n} + Q^{(1)}(x, k_{x}, \omega_{n})$$

[The subscript n indicates we are looking at a normal mode whose frequency ω_n is given to lowest order by Eq.(20b)]

$$Q^{(0)}(x,k_{x},\omega_{L}) \equiv -1/2 \frac{\partial^{2}Q}{\partial k_{x}^{2}} (0,0,\omega_{L}) k_{x}^{2} -i \frac{\partial^{2}Q}{\partial k_{x}^{2}} (0,0,\omega_{L}) xk_{x} + 1/2 \frac{\partial^{2}Q}{\partial x^{2}} (0,0,\omega_{L}) x^{2}$$

$$\lambda_{n} = \lambda_{n}^{(0)} + \lambda_{n}^{(1)} + \dots = (\omega_{n} - \omega_{L}) \frac{\partial Q^{(0)}}{\partial \omega} (0,0,\omega_{L})$$

$$\lambda_{n}^{(m)} \equiv \Delta \omega_{n}^{(m)} \frac{\partial Q^{(0)}}{\partial \omega} (0,0,\omega_{L})$$

$$\omega_{n} = \omega_{L} + \Delta \omega_{n}^{(0)} + \Delta \omega_{n}^{(1)} + \dots$$

$$\psi_{n}(x) = \psi_{n}^{(0)}(x) + \psi_{n}^{(1)}(x) + \dots$$

Note that $Q^{(1)}$ contains all of the terms neglected in Eq.(17). As a result of these terms, ω_n and ψ_n have small corrections $\Delta \omega_n^{(1)} + \Delta \omega_n^{(2)} + \dots$ and $\psi_n^{(1)} + \psi_n^{(2)} + \dots$ Eq.(F1) may be written

$$(Q^{(0)} + Q^{(1)} + \lambda_n^{(0)} + \lambda_n^{(1)} + \dots)(\psi_n^{(0)} + \psi_n^{(1)} + \dots) = 0$$
(F2)

The zero-order terms of Eq.(F2) give Eq.(18), or in our notation

$$[Q^{(0)}(x, -i\frac{\partial}{\partial x}, \omega_{L}) + \lambda_{n}^{(0)}] \psi_{n}^{(0)}(x) = 0$$
 (F3)

where $\psi_n^{(0)}$ and $\lambda_n^{(0)}$ are given by Eq.(19). The first-order terms of Eq.(F2) give

$$(Q^{(0)} + \lambda_n^{(0)}) \psi_n^{(1)} + (Q^{(1)} + \lambda_n^{(1)}) \psi_n^{(0)} = 0$$
 (F4)

Following the usual procedure of quantum mechanical perturbation theory, we expand $\psi_n^{(1)}(x)$ in the eigenfunctions $\psi_m^{(0)}(x)$ of $Q^{(0)}$ [given by Eq.(19a)]:

$$\psi_n^{(1)}(x) = \sum_{m=0}^{\infty} c_{mn} \psi_m^{(0)}(x)$$
 (F5)

where $c_{mn} \equiv \int \psi_{m}^{(0)*}(x) \psi_{n}^{(1)}(x) dx / \int |\psi_{m}^{(0)}(x)|^{2} dx$ Then, using $Q^{(0)} \psi_{m}^{(0)} = -\lambda_{m}^{(0)} \psi_{m}^{(0)}$,

Eq.(F4) becomes

$$\sum_{m=0}^{\infty} (\lambda_n^{(0)} - \lambda_m^{(0)}) c_{mn} \psi_m^{(0)} + (Q^{(1)} + \lambda_n^{(1)}) \psi_n^{(0)} = 0$$
 (F6)

Eq.(F6) can be used to find $\lambda_n^{(1)}$ and c_{mn} for $m \neq n$. If we left-multiply Eq.(F6) by $\psi_n^{(0)*}$ and integrate over x, we obtain

$$\int \psi_{n}^{(0)*} Q^{(1)} \psi_{n}^{(0)} dx + \lambda_{n}^{(1)} \int |\psi_{n}^{(0)}|^{2} dx = 0$$

or $\lambda_{n}^{(1)} = -\int \psi_{n}^{(0)*} Q^{(1)} \psi_{n}^{(0)} dx / \int |\psi_{n}^{(0)}|^{2} dx$ (F7)

Left-multiplying Eq.(F6) by $\psi_m^{(0)*}$ for m≠n and integrating over x yields

$$c_{mn}(\lambda_n^{(0)} - \lambda_m^{(0)}) = -\int \psi_m^{(0)*} Q^{(1)} \psi_n^{(0)} dx / \int |\psi_m^{(0)}|^2 dx$$
 (F8)

Using Eq.(F7) we will show that (with certain assumptions) inequality (21) implies $\lambda_n^{(1)} << \lambda_n^{(0)}$ for small n. Similar arguments (which we will not go through), using Eq.(F8), would show that inequality (21) also implies $c_{mn}^{(1)} << 1$ for m≠n, for small n. Thus we will show that if inequality (21) is satisfied, Eqs.(20a) and (20b) are good approximations to the exact normal mode potentials and frequencies, for small n.

We wish to estimate an upper bound for the expression $\int \psi_n^{(0)*} Q^{(1)} \psi_n^{(0)} dx$ appearing in Eq.(F7). We will consider n = 0; the same arguments apply for any n~ 1.

From Eq.(19a)

$$\psi_0^{(0)}(x) = \exp(-\beta x^2/2)$$

We note that $\psi_0^{(0)}(x)$ is negligibly small for $x \gg (\operatorname{Re}\beta)^{-1/2}$, and that its fourier transform is negligibly small for $k_x \gg \beta^{1/2}$. This suggests that $\int \psi_0^{(0)*} Q^{(1)} \psi_0^{(0)} dx$ will depend on $Q^{(1)}(x, k_x, \omega)$ only for $|x| \leq (\operatorname{Re}\beta)^{-1/2}$ and $|k_x| \leq \beta^{1/2}$, and not for much larger values of x and k_x . Also, if $\lambda_0^{(1)} << \lambda_0^{(0)}$, then $\Delta \omega_0^{(1)} << \Delta \omega_0^{(0)}$, and we need only consider values of ω satisfying $|\omega - \omega_L| \leq \Delta \omega_0^{(0)}$. Then we expect

$$\int \psi_0^{(0)*} Q^{(1)} \psi_0^{(0)} dx \leq Q_{\max}^{(1)} \int \psi_0^{(0)*} \psi_0^{(0)} dx$$
(F9)

where $Q_{\max}^{(1)}$ is the maximum absolute value attained by $Q^{(1)}(x, k_x, \omega)$ for any $|x| \leq (\operatorname{Re}\beta)^{-1/2}$, $|k_x| \leq \beta^{1/2}$ and $|\omega - \omega_L| \leq \Delta \omega_0^{(0)}$. It is possible to imagine operators $Q^{(1)}$ for which Eq.(F9) would not be satisfied, e.g. $Q^{(1)}$ could be exponentially large at $|x| >> (\operatorname{Re}\beta)^{-1/2}$ for $|k_x| >> \beta^{1/2}$. However, we conjecture that Eq.(F9) is true for any physically reasonable choice of $Q^{(1)}$.

Then Eqs.(F7) and (F9) imply $\lambda_0^{(1)} \leq Q_{\max}^{(1)}$. Using the definition following Eq.(F2),

$$Q^{(1)}(\mathbf{x},\mathbf{k}_{\mathbf{x}},\omega) = \sum_{\substack{\mathbf{p},\mathbf{q},\mathbf{r} \geq 0\\\mathbf{p}+\mathbf{q}+2\mathbf{r} \geq 3}} \frac{\partial^{\mathbf{p}+\mathbf{q}+\mathbf{r}} Q(0,0,\omega_{\mathbf{L}})}{\partial x^{\mathbf{p}} \partial \mathbf{k}_{\mathbf{x}}^{\mathbf{q}} \partial \omega^{\mathbf{r}}} \frac{x^{\mathbf{p}}\mathbf{k}_{\mathbf{x}}^{\mathbf{q}}(\omega-\omega_{\mathbf{L}})^{\mathbf{r}}}{\mathbf{p}! \mathbf{q}! \mathbf{r}!}$$

Thus

$$\lambda_{0}^{(1)} \leq Q_{\max}^{(1)} \leq \sum_{\substack{p,q,r \geq 0 \\ p+q+2r \geq 3}} \left| \frac{\partial^{p+q+r} Q(0,0,\omega_{I})}{\partial x^{p} \partial k_{x}^{q} \partial \omega^{r}} \right| \frac{|\operatorname{Re\beta}|^{-p/2} |\beta|^{-q/2} |\Delta \omega_{0}^{(0)}|^{r}}{p! q! r!}$$
(F10)

We need an upper bound on

$$\frac{\partial^{p+q+r} Q(0,0,\omega_L)}{\partial x^p \partial k_x^q \partial \omega^r}$$
. We can estimate this by noting that Q is the sum of

several terms

$$q = \sum_{j} q_{j}$$

where each term Q_j is due to a different physical effect, e.g. the vacuum term, the electron convection term, the ion Landau damping term, etc. Each Q_j is assumed to be nearly independent of x for $|x| \leq L_j$; thus L_j is an effective scale length for the plasma properties associated with Q_j (e.g.

density or density gradient of one species). $Q(0,0,\omega_L) = 0$ because all of the Q_j cancel out at x = 0, $k_x = 0$, $\omega = \omega_L$. In the vicinity of this point, at least two of the Q_j must be much greater than Q. Similarly, for the dominant terms Q_j, $\partial Q_j / \partial x$ and $\partial Q_j / \partial k_x$ will be much greater than $\partial Q / \partial x$ and $\partial Q / \partial k_x$. We will assume, however, that $\partial Q_j / \partial \omega \sim \partial Q / \partial \omega$, $\partial^2 Q_j / \partial k_x^2 \sim \partial^2 Q / \partial k_x^2$ and $\partial^2 Q_j / \partial x^2 \sim \partial^2 Q / \partial x^2$ for at least some of the dominant terms Q_j.

As may be seen from Eq.(12), k_x appears in D only as part of the expression $(k_x^2 + k_y^2)^{1/2}$, hence k_x appears in Q only as part of $[(k_x + k_L)^2 + k_y^2]^{1/2}$. So we expect each term Q_j to be nearly independent of k_x for $k_x << k \equiv (k_L^2 + k_y^2)^{1/2}$. Finally, we assume that all dominant terms Q_j are nearly independent of ω for $|\omega - \omega_L| \leq \Omega$, and that $\partial Q/\partial \omega \sim Q_j/\Omega$. [If, on the contrary, $\partial Q/\partial \omega << Q_j/\Omega$, we can still use the local method if we re-define $\lambda_n^{(0)}$ as $\Delta \omega_n^{(0)} \partial Q^{(0)}/\partial \omega + (1/2) (\Delta \omega_n^{(0)})^2 \partial^2 Q^{(0)}/\partial \omega^2$, and $\lambda_n^{(1)}$ as $\Delta \omega_n^{(1)} \partial Q^{(0)}/\partial \omega + \Delta \omega_n^{(1)} \Delta \omega_n^{(0)} \partial^2 Q^{(0)}/\partial \omega^2$. Then Eqs.(19a), (19b) and (20a) will still be valid if $\lambda_n^{(1)} << \lambda_n^{(0)}$, but Eq.(20b) must be replaced by

$$\omega_{n} = \omega_{L} - (\partial D/\partial \omega) (\partial^{2} D/\partial \omega^{2})^{-1} \pm [(\partial D/\partial \omega)^{2} (\partial^{2} D/\partial \omega^{2})^{-2} - (2nn+\beta) (\partial^{2} D/\partial k_{x}^{2}) (\partial^{2} D/\partial \omega^{2})^{-1}]^{1/2}$$
(F11)

In other words, for each integer n there will be two normal modes, with nearly identical potentials given by Eq.(20a), and with frequencies given by the + and - choices of Eq.(F11). This situation occurs near the values of k and ω where two modes interact, e.g. a drift wave and a Bernstein wave.]

If $Q_j(x, k_x, \omega)$ is analytic in the range $|x| << L_j$, $k_x << k$, $|\omega - \omega_L| << \Omega$, we can write

$$\sum_{\substack{p,q,r \geq 0\\p+q+r \geq 1}} \frac{\partial^{p+q+r} Q_j(0,0,\omega_L)}{\partial x^p \partial k_x^q \omega \partial^r} \frac{x^p k_x^q (\omega - \omega_L)^r}{p! q! r!}$$
(F12)

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Since

$$|Q_{j}(x, k_{x}, \omega) - Q_{j}(0, 0, \omega_{L})| \leq Q_{j}(0, 0, \omega_{L})$$

for all $|x| \leq L_j$, $|k_x| \leq k$, $|\omega - \omega_L| \leq \Omega$, it follows from Eq.(F12) that

$$\frac{\partial^{p+q+r}Q_{j}(0,0,\omega_{L})}{\partial x^{p} \partial k_{x}^{q} \partial \omega^{r}} \lesssim p! q! r! L_{j}^{-p} k^{-q} \Omega^{-r} |Q_{j}(0,0,\omega_{L})|$$

If there are only a few dominant terms Q_{i} ,

$$\frac{\partial^{\mathbf{p}+\mathbf{q}+\mathbf{r}}Q(0,0,\omega_{\mathbf{L}})}{\partial \mathbf{x}^{\mathbf{p}} \partial \mathbf{k}_{\mathbf{x}}^{\mathbf{q}} \partial \omega^{\mathbf{r}}} \lesssim \frac{\partial^{\mathbf{p}+\mathbf{q}+\mathbf{r}}Q_{\mathbf{j}}(0,0,\omega_{\mathbf{L}})}{\partial \mathbf{x}^{\mathbf{p}} \partial \mathbf{k}_{\mathbf{x}}^{\mathbf{q}} \partial \omega^{\mathbf{r}}}$$

it follows that

$$\left|\frac{\partial^{\mathbf{p}+\mathbf{q}+\mathbf{r}}Q(0,0,\omega_{\mathrm{L}})}{\partial x^{\mathrm{p}} \partial k_{\mathrm{x}}^{\mathrm{q}} \partial \omega^{\mathrm{r}}}\right| \leq p! q! r! \mathrm{L}^{-\mathrm{p}} \mathrm{k}^{-\mathrm{q}} \Omega^{-\mathrm{r}} |Q_{\mathrm{j}}(0,0,\omega_{\mathrm{L}})| \qquad (F13)$$

where L is the smallest L for any of the dominant terms Q_j , and Q_j is a typical dominant term. Putting Eq.(F13) into Eq.(F10) yields

$$\lambda_{0}^{(1)} \leq \sum_{\substack{p,q,r \geq 0 \\ p+q+2r \geq 3}}^{(1)} (L | Re\beta|^{1/2})^{-p} (k |\beta|^{-1/2})^{-q} (\Omega | \Delta \omega_{0}^{(0)}|^{-1})^{-r} Q_{j}^{(0,0,\omega_{L})}$$

÷

We want to compare $\lambda_0^{(1)}$ to $\lambda_0^{(0)}$ where

$$\lambda_{0}^{(0)} = \Delta \omega_{0}^{(0)} \partial Q / \partial \omega \sim \Delta \omega_{0}^{(0)} Q_{j} \Omega^{-1}$$

So $\lambda_{0}^{(1)} << \lambda_{0}^{(0)}$ if
 $L << |Re\beta|^{-1/2}$ (F14)

and

 $k << |\beta|^{1/2}$ (F15)

and

$$\Delta \omega_0^{(0)} << \Omega$$
 (F16)

We will assume $\text{Re}\beta \sim \beta$ (this may not be true for plasmas in which magnetic shear is important) so Eq.(F14) becomes

$$L << |\beta|^{-1/2}$$
 (F17)

If B^2 - 4AC ~ -4AC in Eq.(18) (This is certainly true for all of the modes found in Sec. V) then, from the definition of β after Eq.(19),

$$\beta \sim \left(\frac{\partial^2 Q}{\partial x^2} / \frac{\partial^2 Q}{\partial k_x^2}\right)^{1/2} \sim kL^{-1}$$

$$(F18)$$

From Eq.(20b),

$$\Delta \omega_{0}^{(0)} = -(\beta/2) (\partial Q/\partial \omega)^{-1} (\partial^{2}Q/\partial k_{x}^{2}) -(k/2L) (Q_{j}/\Omega)^{-1} (2Q_{j}/k^{2}) = \Omega (kL)^{-1}$$
(F19)

Using Eq.(F18) and F(18), Eq.(F15), (F16), and (F17) all reduce to

Lk >> 1
i.e.
$$L(k_L^2 + k_y^2)^{1/2} >> 1$$
 (F20)

which is inequality (21).

Appendix G

Sketch of Proof that Normal Modes are in the Form of Eq. (32):

Floquet's theorem²⁴ tells us that a differential equation of order m

$$D(x, -i\frac{\partial}{\partial x}, \omega) \phi (x) = 0$$
 (G1)

whore

$$D(x, k_{x}, \omega) = P_{m}(x, \omega) k_{x}^{m} + P_{m-1}(x, \omega) k_{x}^{m-1} + \dots + P_{o}(x, \omega)$$
(G2)

with coefficients $P_i(x,\omega)$ periodic in x with period $2\pi k \frac{-1}{0}$

$$P_{i}(x, \omega) = P_{i}(x + 2\pi k_{o}^{-1}, \omega)$$
 for $i = 1, 2, ... m$

has m independent solutions $\phi_i(x)$ which may be chosen so that

$$\phi_{j}(x + 2\pi k_{o}^{-1}) = \exp \left[2\pi i K_{j}(\omega) k_{o}^{-1}\right] \phi_{j}(x)$$
for j = 1,2,...m (G3)

where each $K_{j}(\omega)$ depends on the functions $P_{o}(x,\omega)$, $P_{i}(x,\omega)$,... $P_{m}(x,\omega)$. Then the most general solution to Eq. (G1) is

$$\phi(\mathbf{x}) = \sum_{j=1}^{m} \mathbf{A}_{j} \phi_{j}(\mathbf{x}) = \sum_{j=1}^{m} \mathbf{A}_{j} \exp(\mathbf{i}\mathbf{K}_{j}\mathbf{x}) \sum_{p=-\infty}^{\infty} \phi_{p,j} \exp(\mathbf{i}p\mathbf{k}_{o}\mathbf{x})$$

where $\sum_{p} |\phi_{p,j}|^{2} = 1$

If we take as boundary conditions the requirement that $\phi(x)$ remain finite as $x \rightarrow \pm \infty$, then

$$A_j = 0$$
 or $ImK_j(\omega) = 0$

for every j, $1 \le j \le m$. If $ImK_j(\omega) = 0$ for only one value of j, then the only $\phi(x)$ which satisfies the boundary conditions is

$$\phi(\mathbf{x}) = \exp(\mathbf{i}\mathbf{K}_{j}\mathbf{x}) \sum_{p=-\infty}^{\infty} \phi_{p,j} \exp(\mathbf{i}p\kappa_{o}\mathbf{x})$$
(G4)

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If $ImK_j(\omega) = 0$ for more than one value of j, then there is a degeneracy at this value of ω , but we are still free to choose the normal modes to be of the form given by Eq.(G4).

In fact $D(x, k_x, \omega)$ is not a polynomial function of k_x , as in Eq.(G2), but is a more complicated analytic function of k_x , given by Eq.(12), so Eq.(G1) is an integral equation rather than a differential equation. Is it possible that there is a normal mode solution $\phi(x)$ of Eq.(G1) [with $D(x, k_x, \omega)$ an arbitrary analytic function of k_x , still periodic in x] which could not be written in the form of Eq.(G4)? Suppose that such a mode $\phi(x)$ does exist. If $\phi(x)$ is differentiable, then there is some wave number k such that $\tilde{\phi}(k_x)$ is negligibly small for $|k_x| \ge k$. There is also some integer m such that $D(x, k_x, \omega)$ can be approximated very well by Eq.(G2) for $|k_x| \le k$. Then $\phi(x)$ would be very close to a normal mode of Eq.(G1) with $D(x, k_x, \omega)$ given by Eq.(G2), which would contradict our supposition that $\phi(x)$ cannot be written in the form of Eq.(G4).

APPENDIX H

Derivation of Eq. (35):

We start with the definitions of $G_{0,s}$ and $G_{\pm 1,s}$ after Eq. (25), and use

$$\exp(-k^{2}a_{s}^{2}) I_{l}(k^{2}a_{s}^{2}) \approx (2\pi)^{-1/2}(ka_{s})^{-1}\exp(-l^{2}/2k^{2}a_{s}^{2})$$

(H1)

(valid when $k_{s}^{2} \stackrel{2}{_{s}} > \ell > 1$)¹⁹ and replace the sum over ℓ by an integral [valid when Im $\omega/\omega_{ci} \geq 1$ so that $(\omega/\omega_{cs}-\ell)^{-1}$ varies slowly with ℓ , and $k_{0} \leq k$ so that $\exp(i\ell\alpha_{\pm 1})$ varies slowly with ℓ ; these conditions also ensure that $\ell > 1$ for most terms that contribute significantly, since typically $|\ell| \approx |\omega|/\omega_{cs}$].

We first consider $G_{0,s}$. With the above approximations,

$$G_{0,s}(k_{x},\omega) = (2\pi)^{-1/2} (\omega_{ps}^{2}/v_{s}^{2}) (ka_{s})^{-1} \int_{-\infty}^{\infty} d\ell \ \ell \ \exp(-\ell^{2}/2k^{2}a_{s}^{2}) (\omega/\omega_{cs}^{-\ell})^{-1}$$
(H2)

We do the integration in Eq. (H2):

$$\int_{-\infty}^{\infty} d\ell \,\ell \,\exp(-\ell^{2}/2k^{2}a_{s}^{2})(\omega/\omega_{cs}-\ell)^{-1} = \int_{-\infty}^{\infty} d\ell \,(\ell-\omega/\omega_{cs}) \,\exp(-\ell^{2}/2k^{2}a_{s}^{2})(\omega/\omega_{cs}-\ell)^{-1} + (\omega/\omega_{cs}) \,\int_{-\infty}^{\infty} d\ell \,\exp(-\ell^{2}/2k^{2}a_{s}^{2})(\omega/\omega_{cs}-\ell)^{-1} = -(2\pi)^{1/2} \,ka_{s}^{2} - (\omega/\omega_{cs})\pi^{1/2} \,Z(\omega/\sqrt{2}k \,v_{s})$$
(H3)

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Eqs. (H2) and (H3) yield the first part of Eq. (35):

$$G_{0,s}(k_{s},\omega) = -(\omega_{ps}^{2}/v_{s}^{2})[1 + \zeta_{0}^{Z}(\zeta_{0})]$$
(H4)

where $\zeta_0 \equiv \omega/\sqrt{2} k v_s$

To calculate $G_{\pm 1,s}$, we note that

$$\exp[-(k^{2} + k_{\pm 1}^{2})a_{s}^{2}/2] = \exp[-(k - k_{\pm 1})^{2}a_{s}^{2}/2]$$
$$\exp(-k k_{\pm 1}a_{s}^{2})$$
(H5)

$$(k - k_{\pm 1})^2 \approx k_x^2 k_0^2 / k_{\pm 1} + \theta'(k_0^3 / k)$$
 (H6)

$$\alpha_{\pm 1} \approx \bar{+} k_{y} k_{0} / k_{\pm 1} + \theta' (k_{0}^{3} / k^{3})$$
 (H7)

where Eqs. (H6) and (H7) are valid when $k_0^{<<} k$, and may be derived from the definitions of $k_{\pm 1}$ and $\alpha_{\pm 1}$ after Eq. (25). Starting with the definition of $G_{\pm 1,s}$ after Eq. (25), and using Eqs. (H1), (H5), (H6) and (H7), we obtain

$$G_{\pm 1,s}(k_{x},\omega) = (\omega_{ps}^{2} \Delta_{s}^{2} v_{s}^{2}) (2\pi k k_{\pm 1}^{a} s_{s}^{2})^{-1/2}$$

$$\int_{-\infty}^{\infty} d\ell \exp(\bar{+}i\ell k_{y}^{k} k_{0}^{2} / k k_{\pm 1}) \exp(-k_{x}^{2} k_{0}^{2} a_{s}^{2} / 2k k_{\pm 1})$$

$$\exp(-\ell^{2} / 2k k_{\pm 1}^{a} a_{s}^{2}) (\ell \pm ik_{y}^{k} k_{0}^{a} s_{s}^{2}) (\omega / \omega_{cs}^{2} - \ell)^{-1} \qquad (H8)$$

Rearranging terms, Eq. (H8) becomes

$$G_{\pm 1,s} = (\omega_{ps}^{2} \Delta_{s}^{2} / 2v_{s}^{2}) (2\pi k k_{\pm 1} a_{s}^{2})^{-1/2} \exp(-k_{0}^{2} a_{s}^{2} k / 2k_{\pm 1})$$

$$\int_{-\infty}^{\infty} d\ell \exp(-\ell^{2} / 2k k_{\pm 1} a_{s}^{2} + i\ell k_{y} k_{0}^{2} / k_{\pm 1} + k_{0}^{2} k_{y}^{2} a_{s}^{2} / 2k k_{\pm 1})$$

$$(\ell \pm ik_{y} k_{0} a_{s}^{2}) (\omega / \omega_{cs}^{2} - \ell)^{-1}$$
(H9)

Changing the variable of integration from ℓ to $\ell' \equiv \ell \pm ik k_0 a_s^2$, the integral in Eq. (H9) becomes

$$\int_{-\infty}^{\infty} d\ell' \exp(-\ell'^2/2k k t^a s^2) \ell' (\omega/\omega_{cs} - \ell' \pm ik k_0 a^2)^{-1}$$

This integral is of the same form as the integral in Eq. (H2), but with ω/ω_{cs} in Eq. (H2) replaced by $\omega/\omega_{cs} \pm ik_y k_0 a_s^2$, and it can be solved in the same way. To obtain the second part of Eq. (35), we must further replace $k_{\pm 1}$ everywhere by k, which is justified if $k_0 << k$. Note that we could not have replaced $k_{\pm 1}$ by k before writing down Eq. (H8), or we would have lost the factor $\exp(-k_x^2 k_0^2 a_s^2/2k_{\pm 1})$ appearing in Eq. (H8).

It might be argued that it was not justified to neglect the higher order terms in Eqs. (H6) and (H7), since these terms, although small compared to the leading terms, may not be small compared to unity when they appear in the exponentials in Eq. (H8); hence they might significantly change Eq. (H8). However, a careful analysis including these higher order terms shows that they are only important when $G_{\pm 1,s}$ is exponentially small, in which case $G_{\pm 1,s}$ would not be important in the dispersion relation anyway.

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APPENDIX I

Outlines of the main program ROOTS [which finds the normal modes $\phi(x)$ and frequencies ω , using the nonlocal method described in Sec. IIIB] and of the three large subroutines FOLLOW [which follows a given branch of $\omega(k_y)$ out from small k_y to large k_y], DISP (which calculates the dispersion function det $A_{p,p}$, for a given ω and k_y) and MRAF [which finds the zeroes of det $A_{p,p}$, (ω) for a given k_y] are shown in the flow charts. These charts are intended only as rough guides to the programs. Details of the I/O and details of how various decisions are made (e.g. how FOLLOW decides when two roots are a complex pair or a fork; how FOLLOW decides when to increase or decrease the increment in k_{yi} ; how MRAF decides when $\phi_0 = 0$) are not shown. Also not shown are certain modes of operation which were not found to be very useful, e.g. a mode of operation where the frequency, rate of spreading and acceleration were found for wave packets localized at different values of x. For such details the listing may be consulted.

The subroutine MRAF evolved from a root-finding routine published by Rodman¹⁶ and revised by Whitley¹⁷, using a method developed by Muller and Traub¹⁸, a second order iterative method analogous to the secant method. The published version was further revised by one of us (A.B.L.) and used as a subroutine in the first version of ROOTS (written by C.K.B., D.F., and A.B.L.). Several additions and changes to ROOTS and its subroutines were made by M.J.G., including:

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15,7

- 1) Extension of the dispersion function routine DISP and its subroutines to include Maxwellian and loss cone distributions [only ring distributions, $g_{g}(v_{1},v_{\parallel},x) \propto \delta(v_{1} - v_{0})$, were used in the first version], and to include many species (e.g. deuterium ring and tritium Maxwellian).
- 2) Use of possible symmetry of dispersion function around imaginary axis [i.e. detA_{p,p}, (ω) = detA_{p,p}, $(-\omega^*)$] in finding roots (symmetry around real axis was used in published version of MRAF¹⁶).
- 4) Optional boundary to region of complex plane in which MRAF can search for roots. When the root-search goes beyond this boundary, the search ends (say at point x_i) and the function is divided by (x-x_i). Eventually the boundary is surrounded by poles, and the root-search tends to stay away from the boundary and find all the roots inside.
- 5) Dividing up of complex plane into several regions, each searched separately by MRAF. This greatly reduces the number of roots missed by MRAF, if a large area of the complex plane, with a large number of roots, is to be searched.
- 6) Modifications in criteria for convergence (i.e. for having found a root) in MRAF, necessary when there are many roots close together.
- 7) Use of eigenfunctions (ϕ_p) as a criterion for convergence in MRAF.

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- Option for using straight line ion orbits in calculating the dispersion function.
- 9) Implementation of subroutine FOLLOW to follow each branch $\omega(k_y)$ through increasing k. Use of eigenfunctions (ϕ_p) to sort out different branches.
- 10) Option which allowed ROOTS to find $\omega(k_y)$ for an infinite medium with constant (1/n) dn/dx, using the local approximation [i.e., $D(x=0,\omega,k_x=0) = 0$] rather than finding normal modes.



-65-

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XO is the k_{y_1} of the starting point YO is the ω/ω_{ci} of the starting point PHIZP = .TRUE. iff $\phi_0 = 0$ for this branch (useful, since this is a property which does not change for a given branch as k_{y_1} changes)



(see on next page)

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(see previous page)



-68-

FLOW CHART OF DISP (OMEGA)



-69-

FLOW CHART OF MRAF (F, etc.)

F is the function which MRAF is finding the zeroes of


(see previous page)



Listing of ROOTS and Subroutines

The version of ROOTS listed here runs on the A-machine (a CDC 7600) of the National Magnetic Fusion Energy Computer Center at Lawrence Livermore Laboratory, using the NEWl compiler (which differs from the better known CHATR compiler in that it does complex division correctly), and the libraries ORDERLIB, CF76LIB, and TV80LIB. Features differing from standard Fortran include I/O (especially graphics routines), tests to see if the time limit has nearly been reached (OOTIM) and treatment of logical variables.

Although the code originated at Livermore, it spent a number of its formative years at Lawrence Berkeley Laboratory (growing from 600 to nearly 3000 cards) and it still contains a number of archaic features from its Berkeley days (e.g. the routines SGNL and WARN).

It is hoped that this code, which should be useful for a wide range of problems, will be made available to the users of the A-machine in the not too distant future, after it has been cleaned up and documented.

15 13.28.56 06/17A 1976 PROGRAM ROOTS(INPUT,HSP,OUTPUT,TAPE10=OUTPUT) C-170A 000001 000002 000003 DUD 10,20,00
 PROGRAM ROOTS(INPUT, HSP, GUTPUT, TAPE10=GUTPUT)
 THE FOLLOWING CHANGES SHOULD BE MADE IN *ROOTS* AND THE SUBROUTINES EVENTUALLY, BUT I HAVE NOT HAD TIME TO MAKE THEM YET, AND PROBABLY WILL NOT FOR GUITE A WHILE -MJG.
 11 *RING2*, THE SUBROUTINE USED FOR RING DISTRIBUTIONS WITH THE LGCAL APPROXIMATION, IS NOT WORKING CORRECITY, AND SHOULD BE CORRECTED
 21 DIFFUSION HAS NOT YET BEEN IMPLEMENTED WITH THE NON-JOCAL (PMAX.GT.O)
 23 DISPERSION RELATIONS, ONLY WITH THE LOCAL (PMAX=0) DISPERSION RELATIONS
 33 THE UNMAGNETIZED PARTICLES OFTION (UNMAG=.TRUE.) HAS NOT YET BEEN IMPLEMENTED FOR RING DISTRIBUTIONS, ONLY FOR MAXWELLIAN AND LOSS COME DISTRIBUTIONS.
 34 SO FAR *MRAF* HAS WORKED GUITE WELL IN FINDING THE STARTING POINTS BUT IT MIGHT NOT BE SO EFFICIENT (I E. IT MIGHT MISS ROOTS) IN FARTHER REACHES OF PARAMETER SPACE. IN THIS CASE IT MIGHT BE NECESSARY TO USE A NYQUIST TECHNIQUE, SUCH AS CALLEN USED IN HIS IN OBJECT ON THE HAW MANY ROOTS THERE ARE AND THEIR APPROXIMATE LOCATIONS, RATHER THAN THE BLIND HUNTING OF *MRAF*.
 MRAF IS PRESENTLY NOT VERY GOOD AT FINDING ROOTS FAR FROM THE REAL AXIS, HENCE SKYAI MUST BE SET LOW, TO BE SURE OF NOT MISSING ANY ROOTS IF HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CONSIDERABLE TIME SAVED ABOVE, SKYAI COULD BE SET HIGHER, AND CORSIDERABLE TIME SAVED THE WAREFUL BLOWS UP WHEN NATS IS SET TOO HIGH. THIS HAS BEEN ABOVE, SKYAI COULD BE SET HIGHER, AND COCKS GECASIONALLY GET LOST
 MAKE *ZEEE* MORE EFFICIENT AND MORE ACCURATE CRIGHT NOW IT IS NEITHER F BOX BO5 ŏŏŏŏ11 000014 000015 000016 000017 000028 000029 000030 000031 000033 000034 000035 WHEN THEY REALLY SHOULD BE ARGUMENTS OF SUBROUTINES, AND VICE VEL EXTERNAL DISP LGGICAL C(200) LGGICAL PHIZ(200), ZEH(100) LGGICAL WARN, AGGT LGGICAL WARN, AGGT LGGICAL WARN, AGGT, SW71, SW31, GPT(10) LGGICAL UNMAG, SINGLX, GGGD, PRDAMP, SNGLX CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMMPLEX (MEGA, DISP, ZERG, ZERG, A(20, 20), PHI(18, 22) CGMPLEX (MASS, KOAI, KOSG, KYAI, KYSG, KOSG, KYAI, KOAI, MASG, PHAX, START, RHIGH, NFL, CYAI CGMMGN/CCRAF/SW, RTC, ITC, EP3, FRMAX, XRMAX, H, RGGT, FUNC, FRGGT, EIGV, SW7, START, RHIGH, NFL, CYAI CGMMGN/CRAF/SV, RSP, NSP1 CGMMGN/CRAF/SV, RSP, NSP1 CGMMGN/CRAF/SV, SNSP1 CGMMGN/CNX/PNP CGMMGN/CCARAM/ RMAX, AIL, EPSI, EPSE, NSPECP, DFUI, DFUE, EP, SINGLX, XL(10), NXL CGMMGN/CRGAM/R, GAM, RE, GAME, R1, RE1, AEAL1 VMW 000043 000044 000045 000045),NLL COMMON/CRGAM/R,GAM,RE,GAME,R1,RE1,AEAI1 COMMON/CFAST/ DK,FKYAI, NPTC,YI,YR,XC, KYD, YM YZ,XZ,Y,X,NPT,NPTZ,ZEH,IGV,AGOT,NFOL,FKYAI2,RMIN COMMON/CFOL/YSTART(100),XSTART(100),NSTART,YEND(100),XEND(100), YMX. 000074 000075 000076 000077 000078 000079 000080

DATA DFU/4*0./ DATA TEXTURE/0./ DATA PRDAMP/.TRUE./ CALL_DROPFILE 000094 000095 ELAPSED = 0. CALL TIMECHK(ACOT,0.,ELAPSED) CALL KEEP80(10HROOTPLOTAA) KXSQ = 0. ŎŎŎŎ**Ŏ**Ŧ CALL KEEPBULIDHROOIPLOIAA) KXSQ = 0. CALL FRAME(1) NRCUT(1) = 0 NZCUT(1) = 0 NCCUT(1) = 0 NCUT(1) = 0 000100 000101 000102 000103 000104 000105 000105 000105 000107 000108 000109 000110 000111 000112 000113 (I.E. ALL WARM PARTICLES ARE UNMAGNETIZED, NOT YET IMPLEMENTED FOR RING DISTRIBUTIONS.) IF PMAX.GT.O, WE USE THE NON-LOCAL DISP.REL. (DEVELOPED BY LANGDON) A SINUSCIDAL DENSITY PROFILE, AND SINGLX,NXL,XL,AND EP ARE IGNORED. OPT(1),OPT(2),ETC. ARE VARIOUS OPTIONS WHICH MAY BE TRUE(1) OR FALSI OPT(1) MEANS THAT ROOTS WITH PHI(0) = 0. WILL NOT BE FOLLOWED. OPT(2) MEANS THAT, IF UNMAGNETIZED IONS ARE USED, ONLY ROOTS WITH PHASE VELOCITY LESS THAN ION THERMAL VELOCITY WILL BE FOLLOWED. READ(ICR, 121)SINGLX, UNMAG,NXL, (OPT(1),I=1,10) 121 FORMAT(2A1,12,1011) IF(SINGLX.EQ.1HT) SINGLX = .TRUE. IF(UNMAG.EQ.1HT) UNMAG = .TRUE. IF(UNMAG.EQ.1HF) UNMAG = .FALSE. IF(SINGLX.EQ.1HF) SINGLX = .FALSE. IF(.NOT.SINGLX) READ(ICR,122) (XL(I),I=1,NXL) 122 FORMAT(10F5.5) SNGLX = SINGLX = 1 UNDER COLOR OF THE HADDONICS WANTED PMAY*KO IS ESSENTIALLY IF(SINGLX) NXL = 1 UNDER COLOR OF THE HADDONICS WANTED PMAY*KO IS ESSENTIALLY UNDER COLOR OF THE UNDER 000117 000118 000119 OR FALSE(0). 000120 000121 000123 000123 000124 000125 000124 000125 000124 000125 000124 000123 000131 000133 000134 000134 000135 000136 000137 000138 000139 000140 000141 000142 000143 000144 000145 000145 000145 000147 000148 000149 000151 000152 000152 000154 RMIN RMIN 2 FORMAT(315,3F5.3,515) NPP = 2*PMAX + 1 IF WE ARE EXECUTING SEVERAL RUNS ON ONE JOB, WE MAY BE OUT OF TIME AT THE BEGINNING OF THIS RUN, SO CHECK TO SEE IF THIS IS THE CASE. CALL TIMECHK(ACOT, 0., ELAPSED) IF(ACOT) CALL EXIT NSPECP = NSPEC DECL DE EN ALL AFAL OMPSO FEST MASS FE DELL DEVE BETA 000158 000159 000161 000162 000163 000164 000165 000165 000167 NSPECP = NSPEC READ(ICR, 5) AIL, AEAI, OMPSQ, EPSI, MASS, EP, DFUI, DFUE, BETA 5 FORMAT(9F5.0) RE AND GAME REFER TO ELECTRONS. SEE RHCALC. READ(ICR, 12) R, GAM, RE, GAME 12 FORMAT(4F5.3) IF WE ARE USING MANY SPECIES OF PARTICLES (NSPEC.GT.0), READ IN THE LARMOR RADIUS, MASS, CHARGE, PARTICLE DENSITY AND PARAMETER J FOR EACH SPECIES, WHERE F(VPERP) = CONST.*(VPERP**J)*MAXWELLIAN ALSO, THE DENSITY GRADIENT EPS, AND DIFFUSION COEFFICIENT DFU. FOR NONLOCAL (PMAX.GT.0) CASE, EPS IS THE DENSITY VARIATION FOR EACH ION SPECIES. IF EPS IS LEFT BLANK, IT IS SET EQUAL TO EPSI. IF(NSPEC.LE.0) GO TO 82 IF(NSPEC.GT.4) NSPEC = 4 DO 83 I = 1, NSPEC READ(ICR, 84) RLARM(I), AMASS(I), CHARG(I), DENSE(I), JAY(I), EPS(I), DFU (1) С 000168 000169 000170 000172 000172 000173 000175 000175 000175 000177 000178 000179 000180 READ(TCR, 64) REARM(T), AMASS(T), CHARG(T), DEROE(T), DEROE(T 000183 000184 CONTINUE GO TO 89 TO 89

000187		82	NSPEC = 4
000188			NSP = 2 DFU(1) = DFUI
000190			DFU(2) = DFUI DFU(3) = DFUE
000192			DFU(4) = DFUE
000194			IF(PMAX.GT.O.AND.I.LE.NSP) EPS(I) = EPSI
000195 000196		108	AMASS(3) = 1./MASS
000197			AMASQ(3) = AMASS(3)**2 1F(R.LE.O.) G0 T0 98
000199			DENSE(1) = R/(R - GAM) $DENSE(2) = -GAM/(R - GAM)$
000201			RLARM(2) = 1./SQRT(R)
000202		98	GG 10 90 JAY(1) = -1
000204 000205		90	IF(AEAI.EQ.O.) GO TO 89 RLARM(3) = AEAI
000206			IF(RE, LE, 0.) GO TO 99 DENSE(3) = RE/(RE - GAME)
000208			DENSE(4) = -GAME/(RE - GAME)
000210			AMASS(4) = AMASS(3)
000211			AMASQ(4) = AMASQ(3) GO TO 89
000213		- 99 - 89	JAY(3) = -1 Fkyal2 = Fkyal
000215			$\begin{array}{l} \text{IF} (\text{PMAX}, \text{EQ}, 0) \text{EIGV} = 0 \\ \text{IF}(\text{FIGV}, (T, 2)) \text{EIGV} = 0 \end{array}$
000217			IGV = EIGV
000219			F(R, GT, O,) R1 = 1./SQRT(R)
000220 000221			IF(RE,GI,O,J) REI = I./SGRI(RE) $KOAI = 2.0*PI*AIL$
000222			KOSQ=KOAI*KOAI NSP1 = NSP + 1
000224	ç		LCULATE THE ELECTRON GUIDING CENTER DENSITY GRADIENT EPSE NEEDED TO
000226	č	F	RST FIND RATIO OF OSCILLATING PART OF PARTICLE DENSITY TO OSCILLATING
000227	C	101	DÖ 93 I = 1, NSPEC
000229			KOAE = KOAI*RLARM(I) RM = 1
000231		00	$\frac{1}{1} \frac{1}{1} \frac{1}$
000233	С	TH	IEN FIND THIS RATIO (TIMES TOTAL ION CHARGE DENSITY) FOR IONS AS A WHOLE
000234 000235			RHEI = 0. D0 94 I = 1, NSP
000236	С	94 Th	RHEI = RHEI + DENSE(I)*CHARG(I)*RH(I)*EPS(I) IFN FIND FOR ELECTRONS AS A WHOLE.
000238	-		RHEE = 0, DC 95 L = NSP1 NSPEC
000240		95	RHEE = RHEE - DENSE(I)*CHARG(I)*RH(I)
000242			DO 96 I = 1, NSP
000243 000244		96	EPSF(I) = ,5*EPS(I)*0MPSQ*CHARG(I)**2*DENSE(I)/AMASS(I) DØ 97 I = NSP1, NSPEC
000245 000246	с	97 Pf	EPSF(I) = .5*EPSE*ØMPSQ*CHARG(I)**2*DENSE(I)/AMASS(I) NINT OUT PARAMETERS
000247	-	103	CALL PARAM(99) PRINT 7 IRMIN RMAX
000249		7	FORMAT (40H ROOTS FOUND BETWEEN CYCLOTRON HARMONICS, 13, 4H AND, 13)
000251			$\frac{1}{NPT} = 0$
000252 000253			NPTC = 0 NPTZ = 0
000254			NFOL = 0 $YMX = FLOAT(RMAX) + .5$
000256			DØ 114 IXL = 1, NXL 58804X - 1 5420
000258			RMX =FLCAT(RMAX) + .5
000259			RHIGH = RMX RMIN = IRMIN
000261 000262	c.		IF(NSW1.NE.O) SW(1) = .TRUE. COMPLEX CONJUGATES ARE ALSO ROOTS. IF THERE IS NO DIFFUSION
000263			SW(3) = .TRUE.
000265		107	IF(DFU(I).GT.O.) SW(3) = .FALSE.
000266 000267			IF(UNMAG) SW(3) = .FALSE. SW(4)=,TRUE.
000268			SW(5) = .TRUE. IE(UNMAG) = .SW(5) = .EALSE
000270	ç	P	IAX = 0 MEANS WE ARE USING A LOCAL APPROXIMATION, IF NOT USING A
000272	U	L(IF (PMAX, GT.O) SW7 = , TRUE,
000273			KYAI = SKYAI KYSQ=KYAI * KYAI
000275			CYAI = KYAI CHI = KYSQ + PMAX*PMAX*KQSQ
000277			MINQ = 2*RMAX
000279	ç	16	THE RANGE OF FREQ. WE ARE INTERESTED IN (RMX - RMIN) IS TOO LARGE
000280	CC	FC Al	JR *MRAF* ID HANDLE, WE BREAK IT INTO REGIONS OF CH CYCLUTRON HARMONICS ID LOOK AT ONE REGION AT A TIME.
000282 000283			CH = 5, RLOW = RMIN
000284		105	RHIGH = RMX

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IF(RHIGH - RLOW.GT.CH+.5) RHIGH = RLOW + CH + .5 CALCULATE QF, KMAX, AND KMIN (USED IN *DISP*) CALL QFKCAL(QF, KMAX, KMIN, RHIGH, RLOW) HOW MANY POLES ARE THERE. NPOLES = 0 HOW CALL = 0 HOW MANY POLES ARE THERE. 000287 000288 000288 С С HOW MANY POLES ARE THERE. NPOLES = 0 DO 92 I = 1, NSPEC 92 NPOLES = NPOLES + KMAX(I) - KMIN(I) + 1 IF THE NUMBER OF ROOTS(CONVERGED OR NOT) EXCEEDS NRTS, OR THE NUMBER OF UNCONVERGED ROOTS(NOT COUNTING NEGATIVES AND CONJUGATES) EXCEEDS N IT IS NOT WORTHWHILE TO LOOK FOR MORE ROOTS (DETERMINED EMPIRICALLY) NRTS = 2*NPP*(NPOLES + 3*NSPEC - 3) NFL = NPP*NPOLES + NPOLES IF(UNMAG) NRTS = 10*NPP IF(UNMAG) NRTS = 10*NPP IF(NRTS.GT.200) NRTS = 200 IF(NRTS.GT.200) NRTS = 200 IF(NRTS.GT.200) NRTS = 100 IF(PMAX.GT.0) GO TO 117 IF(SINGLX) GO TO 116 DREL(I) IS THE RELATIVE DENSITY, N(X)/NO FOR SPECIES I EPS(I) IS THE LOCAL DENSITY GRADIENT OF SPECIES I CALL XLSET(XL(IXL),XKO,EPSI,EPSE,NSP,NSP1,NSPEC,EPS,DREL,RH,KOAI) GO TO 117 16 DO 116 I = 1, NSPEC 18 DREL(I) = 1. 17 CONTINUE STARTING VALUES FOR ROOT SEARCH. IF THESE ARE TOO CLOSE TOGETHER, IT TAKES TOO LONG TO CONVERGE TO A ROOT. START = 0. IF(UNMAG) GO TO 112 BIOMAGN SECTOR TO CONVERGE TO A ROOT. START = 0. IF(UNMAG) GO TO 112 BIOMAGN SECTOR TO CONVERGE TO A ROOT. START = 0. IF(UNMAG) GO TO 112 BIOMAGN SECTOR BOD TO A ROOT. START = 0. IF(UNMAGN BOD TO 112 BIOMAGN SECTOR BOD TO 112 BIOMAG c NFL, Ċ 000298 C C ŎŎŎ<u></u>ŎŎ 000309 000310 117 IT TAKES TOO LONG TO CONVERGE TO A ROOT. START = 0. IF(UNMAG) GO TO 112 P1 = -1.952 + RLOW P2 = -0.951 + RLOW GG TO 113 112 P1 = .001 P2 = EPSI*RH(1)*KOAI*KYAI + .0019 P3 = SQRT(EPSI*RH(1)*KOAI)*KYAI + .0027 113 SW(6) = SW(3) 14 CALL MRAF(DISP,RT,C,NRTS,P1,P2,P3,1.E-06,1.E-12,25) IF(SW(1)) PRINT 10, (C(I),I=1,NRTS) 10 FORMAT(1X,2015) IF(SKYAI.GE.FKYAI) GO TO 114 IF(SKYAI.GE.FKYAI) GO TO 114 IF(SKYAI.GE.FKYAI) GO TO 21 IF(REAL(RT(I)).LT.-1.E-06,AND.SW7) GO TO 21 IF(REAL(RT(I)).LT.-1.E-06,AND.SW7) GO TO 21 IF(REAL(RT(I)).LT.-1.E-06,AND.SW(3)) GO TO 21 IF(CABS(RT(I)).LT.RLOW - 1.) GO TO 21 IF(RHIGH.LT.RMX.AND.CABS(RT(I)).GT.RHIGH - 1.5) GO TO 21 IF(CABS(RT(I)).GT.RMX) GO TO 21 NSTART = NSTART + 1 USE THE GOOD ROOTS AS STARTING POINTS FOR *FOLLOW*. XSTART(NSTART) = KYAI 000313 000320 000321 000323 000324 000325 000326 000327 000331 NSTART = NSTART + 1 USE THE GOOD ROOTS AS STARTING POINTS FOR *FOLLOW*. XSTART(NSTART) = KYAI YSTART(NSTART) = RT(I) ZEH(NSTART) = PHIZ(I) IF(NSTART, GE. 100) GO TO 106 21 CONTINUE RLOW = PLOUD CU C żEH(NSTART) = PHI4LL IF(NSTART, GE, 100) GG TG 106 21 CONTINUE RLOW = RLOW + CH IF(RHIGH.IT.RMX) GG TG 105 NSTRTI = NSTART WE HAVE FOUND ALL THE STARTING PGINTS. NOW PREPARE TG FGLLOW EACH RGGT. 106 RHIGH = 1.E+06 SW31 = sW43 SW71 = SW7 SW(3) = .FALSE. SW(4) = .FALSE. SW(4) = .FALSE. SW(5) = .FALSE. SW(2) = .FAL 000343 000344 000345 000346 000348 000349 c 000359 C 000371 000372

9 FRINT 57, NEGL YSTARTINEGL) XSTARTINEGL) = LEPSED 9 FRINT 716 M TOTES 9 FRINT 716 M PRINT 57, NFØL, YSTART(NFØL), XSTART(NFØL), ELAPSED 57 FØRMAT (5H RØØT, 14, 19H STARTS AT ØMEGA = ,2E22.13,8H KYAI = ,E22.1 3,5X,8H TIME IS,F8.3) IF(CABS2(YSTART(NFØL)).GT.1.E-12) GØ TØ 58 Ĉ 000411 000412 000415 000416 C 000419 000420 000422 000423 С 000429 000430 000433 000434 000436 000437 000438 000439 000440 000441 000442 000443 000443 000444 0004445 C 000448 000449 000451 000452 000453 000453 000456 000457 000459 000460 000463 000464 000466 000467 000469 000470 000471 000472 000473 000473 000474

YINC = (RMX - RMIN)/50. XMIN = IFIX(SKYAI) RANGE = FKYAI - XMIN XRANGE = 500. IF(RANGE,LE.250.) XRANGE = 250. IF(RANGE,LE.100.) XRANGE = 100. IF(RANGE,LE.75.) XRANGE = 50. IF(RANGE,LE.20.) XRANGE = 50. IF(RANGE,LE.20.) XRANGE = 10. IF(RANGE,LE.10.) XRANGE = 10. IF(RANGE,LE.5.) XRANGE = 5. XMAX = XMIN + XRANGE XINC = XRANGE/100. IF(PRDAMP) GG TG 154 IF(UNMAG) GG TG 80 154 IF(.NGT.SW(1)) CALL PAGE IF(NPT.EQ.0) GG TG 79 CALL PRNPLT(X,Y,XMAX,XINC,RMX,YINC,0,0,NPT) PRINT 39 CALL PAGE 79 IF(NPTZ.EQ.0) GG TG 80 CALL PRNPLT(XZ,YZ,XMAX,XINC,RMX,YINC,0,0,NPTZ) PRINT 39 39 FORMAT(11H REAL RGGTS) CALL PARAM(99) 80 IF(.NGT.SW(1)) CALL PAGE IF(.NGT.SW(1)) CALL PAGE CALL PARAM(99) 90 IF(NPTC.EQ.0) GG TG 48 IF(.NGT.SW(1)) CALL PAGE CALL PARAM(99) 90 IF(NPTC.EQ.0) GG TG 48 IF(.NGT.SW(1)) CALL PAGE CALL PARAM(99) PRINT 40 40 FORMAT(28H REAL PART OF COMPLEX RGGTS,) CALL PAGE YMAX = IFIX(YI(MAXGRG) + 1.) 000483 000484 000489 000498 fill PRIME 1/1/C VALLM KAXX (INC, RMX, YINC, O, O, NPTC) CALL PARAM(09) PRINT 40 Of CORMAT(20H REAL PART OF COMPLEX ROOTS,) CALL PARAM(09) TYNC: YMAX/06(0) + 1.) YINC: YMAX/06(0) + 1.) CALL PRAM(09) PRINT 41 PRINT 42 PRINT 42 PRINT 42 PRINT 41 PRINT 42 PRINT 000509 000510 000511 000517 000518 000522 000523 000527 000529 000530 000533 000539 000541 000542 000543 000544 000544 000548 000552 000553 000557 000558 000562 000563

WRITE(IGRAPH, 148) I, OMEGA, KYAI, XX 148 FORMAT(31H FASTEST-OROWING WAVE OF BRANCH, I3, 9H OMEGA = ,2E20.12 ./8H KYAI = ,E20.12,5H X/L=,F6.3) IF(.NOT.UNMAG) GO TO 146 CALL XLSET(XX,XKO,EPSI,EPSE,NSP,NSP1,NSPEC,EPS,DREL,RH,KOAI) ZERO = DISP(CMEGA) DFDW = 1.E+04*DISP(OMEGA + .0001) KXSQ = .002 D2FDK2 = 1.E+03*DISP(OMEGA) 000571 000572 000573 D2FDK2 = 1.E+03*DISP(GMEGA) KXSQ = 0. XX = XX + .0001 CALL XLSET(XX,XK0,EPSI,EPSE,NSP,NSP1,NSPEC,EPS,DREL,RH,KOA1) DFX = 1.E+04*DISP(OMEGA)*AIL TRZER0 = 1.E+12*CABS2(ZER0) IF(CABS2(DFDW).LT.TRZER0.OR.CABS2(D2FDK2).LT.TRZER0.OR.CABS2(DFDX) .LT.TRZER0) G0 T0 151 D2WDK2 = D2FDK2/DFDW DWDX = DFDX/DFDW SPREAD = SQRT(CABS(D2WDK2)/AIMAG(OMEGA)) CHANGE = CABS(OMEGA/DWDX) ERROR = AIMAG(OMEGA) *SPREAD/CHANGE GAMMA = AIMAG(OMEGA) - ERROR WRITE(IGRAPH,150) SPREAD,CHANGE,GAMMA PRINT 150, SPREAD,CHANGE,GAMMA 150 FORMAT(SPREADS A DISTANCE OF",E12.3," ION LARMOR RADII"/" THE FREQUENCY FOR A GIVEN KYAI CHANGES OVER A DISTANCE OF",E12.3," ION LARMOR RA DII"/" GAMMA PROBABLY CLOSER TO ",E12.3/) GO TO 153 451 OUNTAL COMPACE OF DEDX 1.E+03*DISP(OMEGA) D2FDK2 =GO TO 153 151 PRINT 152 152 FORMAT(" 2, ZERO,DFDW,D2FDK2,DFDX SOMETHING WRONG,DISP = ",2E22.13/"DFDW,D2FDK2,DFDX = ",6E 18.11)
153 CONTINUE
16.11)
153 CONTINUE
146 CONTINUE
146 CONTINUE
147 CALL FRAME(1)
C PLOT THE EIGENFUNCTIONS PHI(X) IF DESIRED.
48 CALL EIGPLT
CALL TIMECHK(ACOT, 0., ELAPSED)
PRINT 78, ELAPSED
IF(NOCRT.NE.O) CALL EXIT
C DO THE CRT PLOT OF THE REAL PART OF ROOTS.
XMIN = IFIX(SKYAI)
XMAX = IFIX(FKYAI+1.)
IF(XMAX = 0, FKYAI + 1.) XMAX = FKYAI
IF(XMAX = 0, FKYAI + 1.) XMAX = FKYAI
IF(XMAX = 0, FKYAI + 1.) XMAX = FKYAI
IF(XMAX = 0, FKYAI + 1.) XMAX = FKYAI
IF(XMAX = 0, FKYAI + 1.) XMAX = FKYAI
YMIN = IFIX(RMX - .5) + 1.5
IF(YMAX = 0, FKYAI + 1.) YMAX = RMX
YMIN = IFIX(RMX + 1.) YMAX = RMX
YMIN = IFIX(RMX - .5) + 1.5
IF(YMAX = 0, RMX + 1.) YMAX = RMX
YMIN = IFIX(RMX) - .1.E-06
NX2 = XMAX - XMIN
IF(NX2.GT.100) NX2 = NX2/10
NX3 = 1
IF(NX2.LT.10) NX3 = 2
IF(NX2.LT.5) NX3 = 5
NY2 = 2.*(YMAX - YMIN)
CALL MAPS (XMIN, XMAX, YMIN, YMAX, 0.11328, 1.0, 0.30, 1.0)
CALL PAGE
IF(.NOT.SW31) TEXTURE = 3.
IF(PRDAMP) GO TO 155
IF(UNMAG) GO TO 64
NR = NR - 1
DO 65 J = 1, NR
IMIN = NRCUT(J) + 1
IMAX = NRCUT(J) 18.11) CONTINUE 000613 000614 77 FORMAT(14H PLOTTING ROOT,14,120,7H POINTS) K = 0 DO 72 I = IMIN,IMAX K = K + 1 YPLOT(K) = Y(I) IF(YPLOT(K).LT.YMIN) YPLOT(K) = YMIN XPLOT(K) = X(I) 72 IF(K.GT.500) GO TO 65 65 CALL TRACET(TEXTURE,XPLOT,YPLOT,NUM) 64 IF(NZ.LE.1) GO TO 66 NZ = NZ - 1 DO 67 J = 1,NZ IMIN = NZCUT(J) + 1 IMAX = NZCUT(J) + 1 IMA FORMAT(14H PLOTTING ROOT, 14, 120, 7H POINTS) J, NUM, XLR(J+1) 5, 6H X/L =,F5.2 =,F5.2) DO 73 I = IMIN, K = K + 1 YPLOT(K) = YZ(I) IF(YPLOT(K).LT.YMIN) YPLOT(K) = YMIN XPLOT(K) = XZ(I) IF(K.GT.500) GO TO 67 CALL TRACET(TEXTURE,XPLOT,YPLOT,NUM) IF(NC.LE.1) GO TO 68 NC = NC - 1 NC = NC - 1 USE DOTTED LINE FOR COMPLEX ROOTS. C

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TEXTURE = 2, IF(.NOT.SW31) TEXTURE = 0, DG 69 J = 1 NC IMIN = NCCUT(J+1) NUM = MCIOUT(J+1) NUM = MCIOUT(J+1) NUM = MCIOUT(J+1) IF(SINGLX.GR.PMAX.GT.0) PRINT 77, J.NUM IF(.NOT.SINGLX.AND.PMAX.EG.0) PRINT 126, J.NUM, XLC(J+1) K = 0 DG 74 I = IMIN, IMAX K = K + 1 IMIN, IMAX K = K + 1 IMIN, IMAX K = K + 1 IMIN, YPLOT(K) = YMIN YPLOT(K) = YR(1) IF(YPLOT(K).LT(YMIN) YPLOT(K) = YMIN 74 IF(X.GT.SOU) GG TO 69 69 CALL TRACET(TEXTURE,XPLOT,YPLOT,NUM) 64 TEXTURE = 0, TVKNGE = TVXMAX - TVXMIN FACTY = TVRMGE/(YMAX - YMIN) TVRNGE = TVYMAX - TVYMIN FACTY = TVRNGE/(YMAX - YMIN) LIST THE PARAMETERS UNDER THE CRT PLOT. CALL PARAM(98) CALL SETCH (1, 7, 1, 0, 1, 0, 0) WRITE(IGRAPH, 46) CALL FRAME(1) DG THE CRT PLOT OF THE IMAGINARY PART OF ROOTS. IF(NPTC.E0.0) GG TG 47 NY2 = IFIX(YI(MAXGRG) + 1.) YMAX = NY2 YMIN = 0, NY2 = 2.XNY2 CALL MAPS (XMIN,XMAX,YMIN,YMAX,0.11328,1.0,0.30,1.0) IF (NC.LT.1) GG TO MG 71 J = 1 NG TO IMIN = NCCUT(J), + 1 IMAX = NINO(IMAX - IMIN + 1, 500) K = 0 DG 75 I = IMIN, IMAX K = K + 1 YPLOT(K) = YI(I) VALUE = YI(I) YPLOT(K) = YI(I) YPLOT(K) = YI(I) YPLOT(K) = YI(I) 000670 000671 000672 000673 000673 000675 000675 000678 000678 000681 000682 000686 000687 С Ć 000701 000702 000703 000703 000705 000705 000705 000708 000709 000719 000711 000712 000714 000714 NUM = MINO(IMAX - IMIN + 1, 500) K = 0 D0 75 I = IMIN, IMAX K = K + 1 YPL0T(K) = YI(I) IF(YPL0T(K).LT.YMIN) YPL0T(K) = YMIN XPL0T(K) = XC(I) IF(K.GT.500)G0 T0 71 CALL TRACET(TEXTURE, XPL0T, YPL0T, NUM) FACTY = TVRNGE/(YMAX - YMIN) CALL TRACET(TEXTURE, XPL0T, YPL0T, NUM) FACTY = TVRNGE/(YMAX - YMIN) CALL SETCH(1.,9.,1,0,1,0,0) WRITE(IGRAPH,44) YR(MAXGR0),YI(MAXGR0),XC(MAXGR0),XL0MX WRITE(IGRAPH,158) GAMK2,XC(MGAMK2) WRITE(IGRAPH,41) CALL FRAME(1) CALL FRAME(1) CALL FRAME(1) CALL FRAME(1) CALL TIMECHK(A00T,0.,ELAPSED) PRINT 78, ELAPSED CALL EXIT END 70 000716 000717 000718 000719 ĚND

SUBROUTINE XLSET(XL,XKO,EPSI,EPSE,NSP,NSP1,NSPEC,EPS,DREL,RH,KOAI)
DGES EVERYTHING THAT HAS TO BE DONE WHEN A NEW VALUE OF X/L IS USED.
REAL KOAI
DIMENSION EPS(NSPEC),DREL(NSPEC),RH(NSPEC)
DATA PI/3.141592653589/
XKO = 2.*PI*XL
DO 115 I = 1, NSP
DREL(I) = 1. + EPSI*RH(I)*COS(XKO)
115 EPS(I) = -EPSI*RH(I)*KOAI*SIN(XKO)/DREL(I)
DO 100 I = NSP1,NSPEC
DREL(I) = 1. + EPSE*RH(I)*COS(XKO)
100 EPS(I) = -EPSE*RH(I)*KOAI*SIN(XKO)/DREL(I)
RETURN 000726 000727 000728 000729 000730 000731 000731 000733 000734 000735 С RETURN

SUBRGUTINE DETAIL(N) CALCULATES DETAILED INFORMATION ON CNE RUGT,PH1, RHOE, AND RHOI. CALCULATES DETAILED INFORMATION ON CNE RUGT,PH1, RHOE, AND RAD SAVPHI BEFORE CALLING THIS ROUTINE, IT IS NECESSARY TO CALL DISP AND SAVPHI WHICH CALCULATE ARH AND PH1 (USED IN THIS SUBROUTINE. LOGICAL DICALC COMPLEX A(20,20), PH1(18, 22),PHNORM, RHNORM COMPLEX A(20,20), PH1(18, 22),PHNORM, RHOEM COMPLEX ARH(20,20,41), CMEGA, RHOE(20), RHOI(20) COMPLEX ARH(20,20,41), CMEGA, RHOE(20), RHOI(20) COMPLEX ARH(20,20,41), CMEGA, RHOE(20), RHOI(20) COMMON/CID/LACLALC, ARH COMMON/CID/LACLALC, ARH COMMON/CID/LACLALC, ARH COMMON/CETAL/ DICALC, ARH COMMON/CETAL/ DICALC, ARH COMMON/CID/LARP,NSP1 COMMON/CTVY/TVY,TVY,NSP,NSP1 COMMON/CTVY/TVY,NSP,NSP1 COMMON/CTVY/TVY,TVY,NSP,NSP1 COMMON/CTVY/TVY,NSP,NSP1 COMMON/TVY/NSP,NSP1 COMMON/TTYTUNE/LPENON,LPENOFF,ITALICS,IWINK,INTENSE,IRIGHT,IUP DATA P13.1415926535/ PRINT 19, A(1,1),A(1,2) NP1 = NPP - NPP 19 FORMAT(6E13.4) DO 22 J = 2,NP1 19 FORMAT(6E13.4) DO 22 J = 1,NSPEC PRINT 19, ARH(1,1,J),ARH(1,2,J) 20 FORMAT(4MATRIX FOR SPECIES",13/) PRINT 19, ARH(1,1,J),ARH(1,2,J) 21 PRINT 19, ARH(1,1,J),ARH(1,2,J) 22 PRINT 19, ARH(1NPF,NP1,J), ARH(NPF,NPF,J) CALCULATE RHOE(P), RHOI(P) DO 51 I = 1, NPP RHOE(1) = 0. NP1 = NPP - 1 DO 14 J = NSP1, NSPEC RHOE(1) = 0. NP1 = NPP - 1 DO 14 J = NSP1, NSPEC RHOE(1) = 0. NP1 = NPP - 1 DO 14 J = NSP1, NSPEC RHOE(1) = 0. NP1 = NPP - 1 DO 14 J = NSP1, NSPEC RHOE(1) = ARH(1, 1,J)*PHI(N, 1) + ARH(1 1, 2,J)*PHI(N, 2) 000741 000742 C MUST 000746 000747 000748 000749 000750 000751 000753 000754 000756 000757 000757 000761 000762 000763 000765 000766 000767 000768 000769 000770 000772 000772 000773 000774 000775 000776 000777 000778 C NP1 = NPP - 1 DØ 14 J = NSP1, NSPEC RHØE(1) = ARH(1, 1) + ARH(1, 2, J) * PH1 (N, 2) 1, J) * PHI (N, 000782 000783 000783 000785 000785 +RHOE(1) = ARH(NPP,NP1,J)*PHI(N,NP1) + ARH(NPP,NPP,J)*PHI(N,NPP) 14 RHOE (NPP)) = 1, NSP 1) = ARH(1, 1, J)*PHI(N, 1) + ARH(1, 2, J) * PHI (N, 2) 16 RH01(NPP) +RH01(NPP) D0 1 I = 2 D0 17 J = 1 = ARH(NPP,NP1,J)*PHI(N,NP1) + ARH(NPP,NPP,J)*PHI(N,NPP) 000789 000790 .+RHOI(NPP) D0 1 I = 2, NP1 D0 17 J = NSP1, NSPEC 17 RHOE(I) = ARH(I,I-1,J)*PHI(N,I-1) + ARH(I,I,J)*PHI(N,I) + ARH(I,I+ .1,J)*PHI(N,I+1) + RHOE(I) D0 18 J = 1, NSP 18 RHOI(I) = ARH(I,I-1,J)*PHI(N,I-1) + ARH(I,I,J)*PHI(N,I) + ARH(I,I+ .1,D)*PHI(N,I) + RHOI(I) 000793 000793 000794 000795 Nol(1) = ARR(1,1=1,3)*FHI(N,1=1) + ARR(1,1,3)*FHI(N,1) + ARR(1,1,3)*FHI(N,1) + ARR(1,1,3)*FHI(N,1) + ARR(1,1,3)*FHI(N,1) + ARR(1,1,3)*FHI(N,1) + ARR(1,1,3) + ARR(1,3,3) 1, J)*PHI(N, I+1) + RHOI(I) CONTINUE . 1 000798 C С WRITE(IGRAPH,9) PHI(N,NPP+1),KYAI CALL PARAM(98) CALL PLOT(RHOI,NPP,RHIMAX) NAME = 9H ICN CALL SETCH(1.,10.,1,0,1,0,0) WRITE(IGRAPH,12) NAME,RHIMAX WRITE(IGRAPH,9) PHI(N,NPP+1),KYAI CALL PARAM(98) CALL FRAME(1) CALL SETCH(1.,42.,1,0,1,0,0) RETURN RETURN END

SUBRGUTINE PLOT(F,NP,FMAX) EXTERNAL PLEUNC COMPLEX F(20), FTWIDL(20) DIMENSION C(11) COMMON/CONPLOT/ NPP,FTWIDL,FMAX1 DATA K1,K2,NX,NY /-11,6,100,50/ NPP = NP CALL FRAME(1)	C(2) = 1.0 DG 1 = 1.NP 1 FTU DL(1) = F(1) CALL MAPS(0.11.0.11.0.11328,1.0.0.30,1.0) FMAX1 = 0. DG 2 E(Y=1.50 DG 2 E(Y=1.50 Y = FL0AT(1))50.	CO 3 IX = 1,100 X = FLGAT(IX)/100. GENERIC PLFUNC FXY = PLFUNC(X,Y) 3 IF(FXY.GT.FMAX) FMAX = FXY 2 CONTINUE FMAX1 = FMAX FMAX1 = FMAX CALL LCURVS1(K1,C,K2,NX,NY) RETURN END	SUBROUTINE LCURVS1(K1,C,K2,NX,NY) EXTERNAL PLFUNC DIMENSION C(1) CALL LCURVS(K1,C,K2,PLFUNC,NX,NY) RETURN END
0000844 00008445 00008445 00008448 0000850 000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000844 0000884 00000884 00000884 00000884 00000884 00000884 00000884 00000884 00000884 00000884 0000884 0000884 00000884 00000884 00000884 0000884 0000884 0000884 0000884 0000884 0000884 0000884 0000884 0000884 0000884 00000884 00000884 00000885 0000884 0000885 0000884 00000885 00000885 00000885 00000885 00000885 0000085 0000885 0000085 000085 000085 000085 000085 0000085 0000085 0000085 00000000	000855 000855 000855 0008558 0008558 0008558 0008558 00008555 0000855 0000855 00000055 0000005 0000005 0000005 000000	00000865 00000865 00000865 0000865 0000865 000865 000865 000865 000865 000865 000865 000855 000855 000855 000855 000855 000855 000855 00005 00005 0005 0005 000000	000872 000873 000874 000875 000875 000875 000877

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۹ ۲ GENERIC PLFUNC FUNCTION PLFUNC(X,Y) INTEGEX F(20), EIKY, EIPKO, F INTEGER PMAS, EIKY, EIPKO, F COMMONCONPLOT NP,F,FMAX DATA P1/3,141592653569/ PMAX = (NP-1)/2 FP = 0. DOG 1 J = 1,NP P = J = PMAX - 1 PKO = 2.4XP1xP P = 0. -

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000899	•	SUBROUTINE FOLLOW(XO,YO,PHIZP)
000900	ç	THE STARTING POINTS FOR THE SEARCH FOR THE ROOT AT EACH NEW VALUE OF
000902 000903	С С	LAST 3 VALUES OF KYAI AND EXTENDING THIS PARABOLA TO THE NEW KYAI,
000904 000905	С С	DELT2 IS CUT IN HALF AND WE TRY AGAIN. THIS IS ALSO DONE IF THE ONLY
000906 000907	C C	ROOT FOUND IS IMPLAUSIBLE (BECAUSE THE EIGENFUNCTION CHANGES TO DRIVE TO THE ROOT IS OR OMEGA CHANGES TOO MUCH), THE INCREMENT IS INCREASED IF THE ROOT IS
000908	Č	FOUND VERY CLOSE TO THE STARTING POINT (LESS THAN ERRHIN) AND THERE IS VERY LITTLE CHANGE IN THE EIGENFUNCTION, BUT THE RELATIVE INCREMENT DK (=DELT2/
000910	Č	RYAI) IS NEVER MORE THAN DKMAX, IF DK FALLS BELOW DKMIN, WE GIVE UP
000912	Ŭ	EXTERNAL DISP LEGICAL SW(6), SW7, CV(2), WARN, DEJAVU, PHIZ(200), PHIZP, ZEH(100)
000914		LOGICAL REALZ, FLIP, GOOD(18), PAIR, AOOT, SW71, SW31
000916		LOGICAL UNMAG
000917		COMPLEX YY(2), YINIT
000919		COMPLEX PHIOED(20), PHINEW(20), PHIVED(20)
000921 000922		COMPLEX YREV INTEGER EIGV, PMAX
000923 000924		REAL KYAI, KYSU, KUSU DIMENSION X(2000),Y(2000),XC(1000),YR(1000),YI(1000),XZ(4000),
000925 000926		YZ(4000) COMMON/CIO/ ICR, IHSP, IGRAPH
000927		COMMON/CFAST/DK1,FKYAI1, NPTC,YI,YR,XC, KYD, YMAA, YZ,XZ,Y,X,NPT,NPTZ,ZEH,IGV,AOOT,NFOL,FKYAI2,YMIN
000929		COMMON/CMRAF/ŚW, RTC, ITC, EP3, FRMAX, XRMAX, H, ROOT, FUNC, FROOT, EIGV, SW7. START, RMX, NFL, CYAI
000931		COMMON/DIS/QF, KYSQ, KOSQ, EPSF(4), MASS, KYAI, KOAI, MASQ, PMAX, OMPSQ
000933		COMMON/CPHCK/ PHINEW, PHZP
000934		COMMON/FMX/NF COMMON/FOLEIG/ GOOD, IBRANCH(18)
000936		COMMON/CFOL/YSTART(100), XSTART(100), NSTART, YEND(100), XEND(100),
000938 000939		NSTRT1 COMMON/CCHECK/XCRIT(100),YCRIT(100),NCRIT,DEJAVU,Y2,N,PZCRIT(100)
000940		COMMON/CCMP/ CMPPR, CMCRIT(100) COMMON/CPHZ/ PHIZ
000942		COMMON/CQN/MINQ COMMON/CMAG/ UNMAG
000944		DATA GOOD/10*.FALSE./ DATA FRAMES/3./
000946		DATA AUGT/.FALSE./
000948		DATA NCRIT/O/
000949		DATA FKYAI2/25./
000951		CALL TIMECHK (ACOT, FRAMES, ELAPSED)
000953	ç	FKYALL IS THE VALUE OF FKYAL READ IN. FKYALZ IS A SMALLER VALUE USED
000955 000955	С	FOR REAL ROOTS WHICH WE ARE NOT INTERESTED IN FOLLOWING SO FAR.
000957 000958		IF(XO.GE.FKYAI2) FKYAI = FKYAI1 IF(FKYAI2.GT.FKYAI1) FKYAI = FKYAI1
000959		BASE = .5*X0 BASE2 = .25*FKYAI + X0
000961		PHZP = SGNL(1., PHIZP) FLIP = FALSE
000963		FLIPC = , FALSE, PALSE = FALSE
000965		
000966		ERRMIN = ,0005 ERRMIN = ,0005
000968		SUM2MX = NP*.01 SPREAD =001
000970 000971		YINIT = YO DKMAX = DK1
000972 000973		DKMIN = 1.E-10 IF(SW31) G0 T0 110
000974		IF (AIMAG(YO).LE.O.) GO TO 15 FKYAL = FKYAL1
000976		NGRÓW = NGRÓW + 1 PRINT 55 NGRÓW
000978		
000980		FKYAI = FKYAI1
000981		PRINT 55, NCRIT PRINT 55, NCRIT
000983	ç	IF THE ROOT IS ALREADY COMPLEX AT INITIAL KYAI, GIVE HARMLESS NONSENSE
000985 000986	C	PZCRIT(NCRIT) = 0.
000987 000988		YCRIT(NCRIT) = -1. $XCRIT(NCRIT) = -1.$
000989 000990		CMCRIT(NCRIT) = .TRUE. 15 IF(.NOT.SW71) GO TO 134
000991		YREV = CMPLX(AIMAG(YO),REAL(YO)) IF(REALZ(YREV)) G0 T0 134
000993		NCRIT = NCRIT + 1 PRINT 135 NCRIT

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135 FORMAT(13H MOVING ROOT,14,7H BEGINS) PZCRIT(NCRIT) = 0, YCRIT(NCRIT) = -1, CMCRIT(NCRIT) = -1, CMCRIT(NCRIT) = .FALSE. FIND FIRST AND SECOND DERIVATIVES OF ROOT WITH RESPECT TO KYAI 134 DELT1 = 1.E-05*X0 DELT2 = 1.E-04*X0 IF(SW31.OR.NEGL_LE.NSTRT1) 00 TO 116 000995 000996 000997 000999 000999 001000 001001 001002 001003 001004 001005 001006 001007 001008 001009 001010 IF(\$W31.06.NF0LLE.NSTRT1) GO TO 116 DELT1 = 1.E-10*X0 DELT2 = 1.E-09*X0 DEL11 = 1.E-10*X0 DELT2 = 1.E-09*X0 DEL = 1.E-09 : 116 P3 = Y0 P2 = Y0 - (0, 001) P1 = Y0 - (.001,0.) 16 KYAI = X0 + DELT1 IF(WARN(0)) RETURN KYSQ = KYAI*KYAI CYAI = KYAI CHI = KYSQ + PMAX*PMAX*KOSQ MINQ = 2.*REAL(P3) CALL QINT(CHI) CALL MRAF(DISP,YY,CY,2,P1,P2,P3,1.E-06,1.E-12,25) IF NO PLAUSIBLE ROOT IS FOUND, REDUCE DELT1 AND TRY AGAIN. IF(.NOT.CV(1)) GO TO 29 Y1 = YY(1) IF(.NOT.CV(2)) GO TO 59 IF(PHZP.EQ.SONL(1.,PHIZ(1)))GO TO 59 IF(PHZP.NE.SONL(1.,PHIZ(2)))GO TO 291 Y1 = YY(2) 59 IF(AIMAG(Y1).LT.-1.E-06.AND.SW31) GO TO 292 001011 001012 001013 001014 001015 001016 001017 001018 001021 001022 001022 001023 001022 001025 001025 001025 001028 001028 001029 C iF(PHZP.Ed.SoNL(1, PHIZ(1))GG TG 59
iF(PHZP.KE.SONL(1, PHIZ(2))GG TG 291
y1 = yY(2)
iF(REAL(Y1).LT.-1.E-06.AND.SW31) GG TG 292
iF(REAL(Y1).LT.-1.E-06.AND.SW31) GG TG 294
gG TG 48
gG TG 48
gG TG 48
gG TG 48
gG TG 29
291 WHY =11 \$ GG TG 29
292 WHY =12 \$ GG TG 29
294 WHY =12 \$ GG TG 58
DELT1 = DELT1/10
FF(SW(1) AND WHY E0.1) PRINT 57, PHIZP, PHIZ(1)
57 FGRMAT(9H PHIZP IS,L5,11H PHIZ(1) IS,L6)
GG TG 16
48 IF(19V.EQ.0) GG TG 68
DG 31 I = 1.NP
9 PHIVLD(1) = PHINEW(1)
64 A = (0, 0, 0)
65 A = (0, 0, 0)
65 A = (0, 0, 0)
65 A = (0, 0, 0)
66 A = (0, 0, 0)
67 G 3
1 = 1.NP
9 PHIVLD(1) = PHINEW(1)
11 GE (100 WAS ORIGINALLY 3 OR 4, WE ONLY LET IT BE EQUAL TH
07 GG TG 27
8 IF(WARN(0)) RETURN
EVEN IF EIGV WAS ORIGINALLY 3 OR 4, WE ONLY LET IT BE EQUAL TH
07 GG TG 27
8 IF(WARN(0)) RETURN
EVEN IF EIGV WAS ORIGINALLY 3 OR 4, WE ONLY LET IT BE EQUAL TH
07 GG TG 27
8 IF(WARN(0)) RETURN
EVEN VERVERY SG GFTEN, GTHERWISE WE WILL BE PLOTTING A
RIDICULCUS NUMBER OF EIGENFUNCTIONS.
IF(10V.EQ.0) GG TG 53
IF(10V.EQ.0) GG TG 53
IF(10V.EQ.0) GG TG 53
1F(10V.EQ.0) GG TG 54
153 IF(XYA1.LT.BASE2) GG TG 54
154 FGRMAT(5X, 6H TIME VERY SG OFTEN, TG MAKE SURE WE ARE NGT ALMOST
CALL GFKCALCGF, KMAX, KMIN,P3+3, P3)
CALL GFKCALCGF, KMAX,KMIN,P3+3, P3)
CALL GFKCALCGF, KMAX,KMIN,P3+3, P3)
CALL GFKCALCGF, YAMAX,KMIN,P3+3, P3)
CALL GFKCALCGF, YAMAX,KMIN,P3+3, P3)
CALL MRAF(DISP,YY, CV.2,P1,P2,P3,1.E-06,1.E-12,25)
Y2 = YY(1)
Y3 = YY(2)
IF((1)
Y3 = YY(2)
IF((1)
Y3 = YY(2)
IF((1)
Y4 = YY(2)
IF((1)
Y4 = Y10)
Y4 = Y10
Y4
Y4 = Y4 Y4
Y4 Y4 Y4 001033 001032 001033 001033 001033 001035 001035 001035 001038 001039 001040 001041 001044 001044 001045 001045 001049 001049 00105 001051 001052 TO ITS 001052 001053 001054 001055 001055 001057 001058 001058 001058 001060 TO MAKE SURE WE ARE NOT ALMOST OUT OF TIME 001061 С 001063 001064 001065 001065 001066 001067 001068 001070 001071 001072 001073 001074 001075 001076 001077 001078 001078 CALL MKAP(DISP,TT, CV,Z,FT,FZ,FG,TTL CO,TTL T2,LC, Y2 = YY(1) Y3 = YY(2) IF(.NOT.CV(1)) GO TO 90 E CHANGE LOGICAL VARIABLES TO REAL VARIABLES, SO WE CAN USE THEM V IF(A.EQ.B) STATEMENTS. PHZ1 = SGNL(1.,PHIZ(1)) PHZ2 = SGNL(1.,PHIZ(2)) IF(.NOT.SW31) GO TO 121 RL3 = SGNL(1.,REALZ(Y2)) RL3 = SGNL(1.,REALZ(Y2)) RL4 = SGNL(1.,REALZ(Y3)) RL1 = SGNL(1.,REALZ(Y3)) RL7 = SGNL(1.,REALZ(YREV)) YREV = CMPLX(AIMAG(Y2),REAL(Y2)) AIM2 = SGNL(1.,REALZ(YREV)) YREV = CMPLX(AIMAG(Y3),REAL(Y3)) AIM3 = SGNL(1.,REALZ(YREV)) YREV = CMPLX(AIMAG(Y1NIT),REAL(YINIT)) AIM1 = SGNL(1.,REALZ(YREV)) IF(.NOT.CV(2)) GO TO 14 F TWO PLAUSIBLE ROOTS, GO TO 20 IF(PHZ1.EQ.PHZ2) GO TO 91 YY(1 WE C ÎÑ 001079 001080 001081 001082 001083 001083 001084 121 001086 001087 001088 120 ĬF С C NØ 001089 001090

001091 С 001091 001092 001093 001094 001095 001095 001096 001097 001098 001099 001100 C 001101 001102 001103 001104 001105 001106 001106 001107 001108 001109 001110 001111 С ŏŏiii3 001114 C 001115 001116 001117 č 001118
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 C č С 0000 C 128 FORMAT(29H WE HAVE KERCHED THOUSING WEET) FORMAT(29H WE HAVE KERCHED THOUSING WEET) FORMAT(29H WE HAVE KERCHED THOUSING) FORMATION FORMATION FORMATION FORMATION FORMATION FOR THE SEGURATION FORMATION FOR THE SEGURATION FORMATION FOR THE SEGURATION FOR С 001151 001152 001153 001154 001155 001156 001157 С 001158 001159 001160 001161 001162 001163 001164 P2 = 10 - (0.,.1)*AIMAG(Y0) X0 = KYAI CALL TIMECHK(A00T,FRAMES,ELAPSED) IF(CMPPR) PRINT 21, NCRIT,YCRIT(NCRIT),XCRIT(NCRIT),ELAPSED 21 FORMAT(13H GROWING R00T,14,19H BEGINS AT OMEGA = ,E22.13, 8H KYAI = ,E22.13,5X,8H TIME IS,F8.3) IF(.N0T.CMPPR) PRINT 129,NCRIT,YCRIT(NCRIT),XCRIT(NCRIT),ELAPSED 129 FORMAT(13H MOVING R00T,14,19H BEGINS AT GAMMA = ,E22.13, 8H KYAI = ,E22.13,5X,8H TIME IS,F8.3) IF(A00T) RETURN G0 TO 16 13 NSTART = NSTART + 1 RECORD THE END OF A GROWING(OR MOVING) R00T, AND ADD THE TWO NEW REAL (OR PURELY IMAGINARY) R00TS TO THE LIST OF R00TS TO BE FOLLOWED LATER. (OR, IF THIS IS THE END OF A GROWING R00T AND WE ARE USING AN IRREVERSIBLE DISPERSION RELATION, ADD THE DAMPED R00T TO THE LIST. YSTART(NSTART) = Y2 001166 001 167 001167 001168 001169 001170 001171 001172 001173 001174 001175 001176 001177 001177 0000 001179 001180 001181

XSTART(NSTART) = KYAI ZEH(NSTART) = PHIZ(1) IF(.NGT.SW31.AND.AIMAG(Y1).GT.O..AND.CMPPR) GG TG 111 NSTART = NSTART + 1 YSTART(NSTART) = Y3 XSTART(NSTART) = KYAI ZEH(NSTART) = PHIZ(2) IF(.NGT.SW31.AND.CMPPR) GG TG 115 49 YEND(NCRIT) = REAL(Y2) IF(.NGT.CMPPR) YEND(NCRIT) = AIMAG(Y2) XEND(NCRIT) = KYAI IF(.NGT.CMPPR) PRINT 22, NCRIT,YEND(NCRIT),XEND(NCRIT) ZE FORMAT (13H GROWING ROOT,14,17H ENDS AT OMEGA = ,E22.13, 0.8H KYAI = ,E22.13) IF(.NGT.CMPPR) PRINT 130,NCRIT,YEND(NCRIT),XEND(NCRIT) 30 FORMAT (13H GROWING ROOT,14,17H ENDS AT GAMMA = ,E22.13, 0.8H KYAI = ,E22.13) IF(.NGT.CMPPR) PRINT 130,NCRIT,YEND(NCRIT),XEND(NCRIT) PRINT 22, NGROW, RY2, KYAI 115 Y2 = Y21 RETURN IF WE HAVE FOUND ONLY ONE ROOT, SEE IF HAS CHANGED TOO MUCH IN SOME WAY (E.G. FROM REAL TO COMPLEX OR VICE VERSA, OR IN ITS EIGENFUNCTION), AND IF SC, REDUCE DK. 14 IF(AIM2.LT.O. AND.RL2,LT.O.) GG TG 132 IF(.NGT.SW31) GG TG 131 IF(RL2.NE.ALI) GG TG 133 132 IF(REAL(Y2).LT.-1.E-O6.AND.SW31) GG TG 95 IF(REAL(Y2).LT.-1.E-O6.AND.SW31) GG TG 96 IF(PH21.NE.PHZP) GTG 97 FOR PHI(0) = 0 ROOTS, PHINEW IS NORMALIZED IN SUCH A WAY THAT PHI(-NP) (=PHINEW(1)) ALWAYS HAS POSITIVE REAL PART (SEE DEFINITION OF ARMM IN WARAF&).IF PHI(-NP) HAS SWITCHED FROM -I TO -... (=PHINEW(1)) ALWAYS HAS POSITIVE REAL PART (SEE DEFINITION OF ARMM IN WARAF&).IF PHIO(NERDED TOG TG 14 IF(MOT, SW31) IF DHIO YA SUMTALIZED IN SUCH A WAY THAT PHI(-NP) (=PHINEW(1)) ALWAYS HAS POSITIVE REAL PART (SEE DEFINITION OF ARMM IN WARAF&).IF PHIOLONE OF TA 26 001182
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 0012228 C Ĉ BY -1 SO IT IS NORMALIZED IN THE SAME WAY AN ANALYZED THEM. THEM. IF(IGV,EQ.O) GO TO 11 IF(.NOT.PHIZP) GO TO 36 IF(AIMAG(PHINEW(1))*AIMAG(PHIOLD(1)).LT.O.) GO TO 35 GO TO 36 35 IF(ABS(AIMAG(PHINEW(1))).GT.ABS(REAL(PHINEW(1)))) GO TO 37 GO TO 36 37 DO 38 I = 1, NP PHIVLD(I) = -PHIVLD(I) 38 PHIOLD(I) = -PHIVLD(I) 38 PHIOLD(I) = -PHIOLD(I) FLIP = .TRUE. 36 SUM2 = 0. SEE HOW MUCH THE EIGENFUNCTION HAS CHANGED, AND, IF DESIRED, 36 SUM2 = 0. SEE HOW MUCH THE EIGENFUNCTION HAS CHANGED, AND, IF DESIRED, PRINT OUT THE COMPONENTS WHECH HAVE CHANGED DRASTICALLY. DO 30 I = 1, NP DIFF2 = CABS2(PHINEW(I)-PHIOLD(I)-(PHIOLD(I)-PHIVLD(I))*(DELT2/ .DELT1)) SUM2 = SUM2 + DIFF2 DO 30 I = 1, NP DIFF2 = CABS2(PHINEW(I)-PHIOLD(I)-(PHIOLD(I)-PHIVLD(I))*(DELT2/ DIFF2.E. CABS2(PHINEW(I)-PHIOLD(I)-(PHIOLD(I))-PHIVLD(I))*(DELT2/ DIFF2.E. O1) GO TO 30 33 IF(SW(1)) PRINT 34 I, PHINEW(I), PHIOLD(I), PHIVLD(I) 34 FORMAT(3H PHINEW(I)2, SH) =, 2214.6, 10H PHIOLD = , 4E14.6) 35 CONTINUE 16 F(SW(1)) PRINT 62 SUM2 27 FTHE EIGEMENNCTION HAS CHANGED TOO MUCH(I.E. IF SUM2 IS TOO LARGE) REDUCE DK AND TRY AGAIN. 16 F(SUM2.0T.SUM2M2) GO TO 63 17 FTHE EIGEMENNCTION HAS CHANGED TOO MUCH(I.E. IF SUM2 IS TOO LARGE) 70 G TO 10 63 IF(FAIR) GO TO 61 17 FLIPC = , FALSE. 17 0 GO TO 10 63 IF(FAIR) GO TO 103 SCMCTIME CASE PHINEW AND PHIOLD WILL BE NORMALIZED DIFFERENTLY. SEE 17 FLIPC 30 TO 72 18 (MT.FLIP.AND..NOT.PHIZP) GO TO 37 17 (FLIPC) GO TO 72 18 (MT.FLIP.AND..NOT.PHIZP) GO TO 37 17 (FLIPC) GO TO 73 SCMCTIMES TWO ROOTS PASS THROUGH EACH OTHER AT OMEGA = 0 WITHOUT INTERRACTING, IN WHICH CASE THE PHI(I) OF ONE ROOT IS THE CASE, SWITCH TO FOLLOWING THE OTHER ROOT. 74 IF(SW31.AND.SW71.AND.CABS2(Y2).LT.1.E-06)GO TO 70 96 TO 98 17 FLIP = .FALSE. 74 IF(SW31.AND.SW71.AND.CABS2(Y2).LT.1.E-06)GO TO 70 96 TO 98 17 FLIP = .FALSE. 74 IF(SW31.AND.SW71.AND.CABS2(Y2).LT.1.E-06)GO TO 70 96 TO 98 97 PHIVLD(I) = -PHIVLD(I) 97 PHIVLD(I) = CONJ0(PHIVLD(I)) 71 PHIOLD(I) = CONJ0(PHIVLD(I)) 72 PHICE = .TRUE. 90 TO 36 . C C č 001262 001262 001262 001263 001264 0012664 0012666 0012666 0012666 001267 0012669 001270 0012772 0012773 0012773 0012756 С Č č C 001276

001277 001278 001279 001280 001281 001282 001282 001283 001283 001285 001285 001286 001286 001288 001289 001299 001292 001293 001293 001293 TECROPY OF CARACTERIANY SO TO 12 DEC 10 TF (SUMP2, 01) OT (CK/DKMAX)***0 KG = DK*SORT(.01/SUM2) DC = DKMAX TF (CRMCP/ERRMIN).OT (DK/DKMAX)***0 KG = DK*SORT(.01/SUM2) DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY, OR WE MAY GET ON THE WRONG ROOT. DC = DKMAX SORT TO RAPIDLY OF THE DEC TO DC = DKMAX SORT TO RAPIDLY OF THE THE DEC TO DC = DKMAX SORT THE NUMBER OF FRAMES TO BE PLOTTED, SINCE THIS IS USED DF PANDED THE NUMBER OF FRAMES TO BE PLOTTED, SINCE THIS IS USED DF PANDED THE NUMBER OF FRAMES TO BE PLOTTED, SINCE THIS IS USED DF PANDED THE NUMBER OF FRAMES TO BE PLOTTED, SINCE THIS IS USED DF PANDED THE NUMBER OF FRAMES TO BE PLOTTED, SINCE THIS IS USED DF PANDED THE SORT THE NUMBER OF FRAMES TO BE PLOTTED, SINCE THIS IS USED DF PANDED THE SORT THE SOR CO12976 CO12978 CO13001 CO13002 CO13003 CO13005 CO13007 CO1300 С C C С С 001348 001349 001350 001351 001352 001353 001354 001355 C C 001356 001357 001358 001359 C 001360 001361 001362 001363 001364 001365 001366 С 001368 001369 001370 001371

17 IF(NPTC.EG.O) GG TG 50 IF(ABS(AIMAG(Y2) - Y(INPTC)).GT.OI) GG TG 50 DGRPH=((KYAI-XC(NPTC))/FKYAI)**2+((REAL(Y2)-YR(NPTC))/YMAX)**2 IF(DGRPH.LT.4.E-O6) GG TG 19 C RECORD CGMPLEX RGGT 50 NPTC = NPTC + 1 YR(NPTC) = REAL(Y2) YI(NPTC) = REAL(Y2) YI(NPTC) = REAL(Y2) YI(NPTC) = KYAI IF(NPTC.GE.IOOO) RETURN 19 Y0 = Y1 Y1 = Y2 X0 = X0 + DELT1 DELT1 = DELT2 DELT2 = DK*KYAI C INCREASE KYAI AND PREDICT WHERE RGGT WILL BE (P3). 27 KYAI = KYAI + DELT2 FG(XYAI.AND.SW3I.AND.CABS2(P3).LT.1.E-O6) P3 = .0005 C IF THE PREDICTED NEW RGGT IS TGG FAR FRGM THE GLD RGGT, EITHER RELATIVELY C GR ABSGLUTELY, THE PREDICTION IS UNRELIABLE, SG REDUCE DK AND TRY AGAIN. IF(CABS2(Y2-P3).GT.1.) GG TG 45 IF(CABS2(Y2-P3).GT.1.) GG TG 45 IF(CABS2(Y2-P3).CT.1.E-O6) GG TG 26 IF(REAL(P3).EG.0.,AND.AIMAG(P3).EG.0.) GG TG 45 IF(CABS2(Y2-P3).CT.1.O.)).LT.04)GG TG 26 IF(REAL(P3).EG.0.,AND.AIMAG(P3).EG.0.) GG TG 45 IF(CABS2(Y2-P3).CT.1.) GJ TG 45 IF(CABS2(Y2-P3).CT.1.) GG TG 45 IF(CABS2(Y2-P3).CT.1.) SFREAD C'XI = KYAI KYSG = CXAIXKYAI CHI = KYSG + FMAX*FKASG CHI = 001372 001373 001374 001375 001375 001376 001377 001378 001379 001380 001382 001383 001383 001384 001385 001385 001385 001387 001388 001389 001390 001390 001391 001393 001393 001394 001395 001396 001396 001396 001396 001400 001401 001402 001403 EITHER RELATIVELY 001402 001403 001404 001405 001406 001407 001408 001408 L2 = L3 = (1.,0.)*SPREAD CYAI = KYAI KYSQ = KYAI*KYAI CHI = KYSQ + PMAX*PMAX*KOSQ MINQ = 2.*REAL(P3) C SUBROUTINE *QINT* PRINTS OUT THE NEW KYAI IF DESIRED AND CALCULATES C THE PARAMETERS GMAX AND NMAX USED IN THE DISPERSION FUNCTION CALCULATION CALL QINT(CHI) IF(SW(1).AND..NOT.SW(2)) PRINT 65, P3 65 FORMAT(5H P3 =,2E22.13) GO TO 8 C IF DK HAS BEEN REDUCED, PRINT THIS FACT(IF DESIRED) AND TELL WHY. THIS C INFORMATION IS INVALUABLE FOR DEBUGGING. 90 IWHY = 0 \$ GO TO 9 91 IWHY = 1 \$ GO TO 9 93 IWHY = 2 \$ GO TO 9 93 IWHY = 3 \$ GO TO 9 94 IWHY = 4 \$ GO TO 9 94 IWHY = 4 \$ GO TO 9 001409 001410 001411 001412 001413 001413 001415 001415 001416 001416 001417 001418 001419 001419 001420 001421 001422 001423 001423 001425 001425 001425 001428 001428 001428 001428 TO 3 5 60 4 5 60 5 5 60 6 5 60 7 5 60 8 60 9 5 60 94 95 = 9 = TO 9 **9**6 = 97 98 = ŤŎ ğ İWHY 001430 ŤØ 9 TØ 9 = 001431 001432 001433 99 100\$ GC TC 9 122 \$ GC TC 9 124 \$ GC TC 124 \$ GC TC 126 \$ GC TC 133 \$ GC TC İŴHŶ = 100 122 124 126 133 IWHY = 001433 001434 001435 001436 001437 I WHY = I WHY = I WHY = I WHY = g 9 33 IWHY = 133 \$ GO TO 9
136 IWHY = 136 \$ GO TO 9
136 IWHY = 136 \$ GO TO 9
103 PAIR = .FALSE.
 IWHY = 13 \$ GO TO 9
9 KYAI = KYAI - DELT2
 DK = .5*DK
 IF(DK.LT.DKMIN) GO TO 46
 DELT2 = DK*KYAI
 IF(SW(1)) PRINT 10, DK, IWHY
10 FORMAT(15H DK REDUCED TO , E12.6,11H FOR REASON,14)
 IF(.NOT.SW(1)) GO TO 27
 IF(IWHY.EQ.1.OR.IWHY.EQ.7) PRINT 43,PHZP,PHIZP,PHZ1,
 .PHIZ(1),PHZ2,PHIZ(2)
43 FORMAT(8H PHZP = ,E22.13,L5/8H PHZ1 = ,E22.13,L5/8H PHZ2 = ,E22.13
 .L5) 001438 103 001438 001439 001440 001441 001442 001443 001444 001445 001446 001447 001448 001448 001449 001450 001451 001452 001453 001453 001455 001455 15) 00 TO 27 Y2 = Y1 GG 10 27
58 Y2 = Y1
IF DK OR DELT1 IS LESS THAN DKMIN, GIVE UP ON THIS ROOT.
46 PRINT 47, IWHY
47 FORMAT(" DK IS LESS THAN DKMIN, FOR REASON ",14/" EITHER .QMAX AND NMAX (IN SUBROUTINES QINT AND QCALC) ARE TOO SMALL," .Y" OR ZEE IS NOT ACCURATE ENOUGH, OR WE HAVE ENCOUNTERED A TOPO-" ./"LOGICAL SITUATION WHICH THE PROGRAM CANNOT HANDLE") IF(IWHY.NE.100) RETURN
WE HAVE REACHED A CRITICAL POINT, IN THE IRREVERSIBLE CASE. CHECK IF WE'VE BEEN HERE BEFORE. IF NOT, GO JUST PAST IT AND USE THE TWO NEW ROOTS AS STARTING POINTS. CALL CHECK IF(.NOT.DEJAVU) GO TO 112 PRINT 113, N
113 FORMAT(31H WE HAVE REACHED CRITICAL POINT, IA) Υī 58 C 001456 001457 001458 001459 001460 IF 001461 Č 001462 001463 001464 001465 001466

<pre>II2 KFUR KYSG = KYAI + 10.*DELT2/((DELT1/DELT2)*CABS((Y1-Y2)/(Y0-Y1))-1.) KYSG = KYAI CYAI = KYAI CYAI = KYAI CYAI = KYAI CYAI = KYAI CYAI = KYAI CYAI = KYAI CYAI = KYAI CYAI = KYAI F(SONL(1, PHIZ(1)).NE.PHZP.OR.SONL(11, PHIZ(2)).NE.PHZP) RETURN F(SONL(1, PHIZ(1)).NE.PHZP.OR.SONL(11, PHIZ(2)).NE.PHZP) RETURN F(SONL(1, PHIZ(1)).NE.PHZP.OR.SONL(11, PHIZ(2)).NE.PHZP) RETURN Y2 = YY(2) Y2 = YY(2) Y2 = YY(2) NCRIT = NCRIT + 1 F(NCRIT OF 100) F(NCRIT = NCRIT + 1 F(NCRIT OF 100) F(NCRIT OF 100) F(NCRIT) = FHZP Y2 = YY(1) Y2 /pre>	SUBROUTINE TIMECHK (AGOT, FRAMES, ELAFSD) SUBROUTINE TIMECHK (AGOT, FRAMES, ELAFSD) CHECKS TO SEE TO VER AFFER EXECUTOR TOL MAKE STREAMES IN THIS FRAMES ENDORS OF FRAMES TO BE PLOTIED EMPIRICALLY WE ARE ANNOSI OUT OF TIME, TAKE ABOUT 1 CU., AGOT 1 FWE ARMOS IN THE CDC7600. TAKE ABOUT 1 CU., AGOT 1 FWE ARMOS TOTT OF TIME, TAKE ABOUT 1 CU., AGOT 1 FWE ARMOS TOTT OF TIME, TAKE CU (AS USED AT LUL) IS ABOUT 0.25 SECONDS ON THE CDC7600. COMMON/ONNOS COMMON/	C CHECKS TG SEE IF A COMPLEX ROOT HAS ALREADY BEEN REACHED BY ANOTHER ROUTE. COMPLEX PHINEW(20) COMPLEX PHINEW(20) COMPLEX PHINEW(20) COMPLEX PHINEW(20) COMPLEX PHINEW(20) COMPLEX V COMPLEX V	-89-
0001475 0001475 0001475 00014775 00014775 00014775 00014775 00014775 00014775 00014775 00014775 00014775 0001482 0001482 0001482 0001482 0001482 0001482 0001482 0001482 0001482 0001482 0001475 00014775 00000000000000000000000000000000000	00000000000000000000000000000000000000		196100

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001542 001543 001544 001545 001546 001546 001546 001548 001549 001550 LOGICAL REALZ FUNCTION REALZ(Z) COMPLEX Z REALZ = .TRUE, PREC = 1.E-06 IF(ABS(AIMAG(Z)).GT.PREC) REALZ = .FALSE. RETURN END

 $\begin{array}{c} 001551\\ 001552\\ 001552\\ 001555\\ 001555\\ 001555\\ 001555\\ 001555\\ 001555\\ 001555\\ 001555\\ 001556\\ 001562\\ 001562\\ 001565\\ 001565\\ 001565\\ 001565\\ 001565\\ 001565\\ 001575\\ 0005\\ 00055\\ 0005\\ 0005\\ 0005\\ 0005\\ 0005\\ 0005\\ 0005\\ 0005\\ 0005\\ 0005$ SUBROUTINE QINT(CHI) INTERFACE BETWEEN *QCALC* AND *ROOTS* OR *FOLLOW*) INTEGER QMAX LOGICAL SW(6) REAL KYAI COMMON/CMRAF/SW,RTC,ITC,EP3,FRMAX,XRMAX,H,ROOT,FUNC,FROOT,EIGV, SW7,START,RMX,NFL,KYAI COMMON/CRBESJ/ QMAX(4),NMAX(4) COMMON/CSPEC/ RLARM(4),AMASS(4),CHARG(4),JAY(4),DENSE(4),NSPEC DO 1 I = 1 NSPEC IF(RLARM(I).EQ.O.) GO TO 1 CHI2 = CHI*RLARM(I)*RLARM(I) IF(JAY(I).EQ.O) GO TO 2 CALL QCALC(CHI2,QMAX(I),NMAX(I),1) GO TO 3 2 CALL QCALC(CHI2,QMAX(I),NMAX(I),0) С GU TU 3 2 CALL QCALC(CHI2,QMAX(I),NMAX(I),0) NMAX(I) = NMAX(I) + JAY(I) 3 IF(SW(1)) PRINT 7,KYAI,QMAX(I),NMAX(I),I 7 FORMAT(7H KYAI =,E22.13,8H QMAX = ,I4,8H NMAX = ,I4,12H FOR SPECIE CONTINUE RETURN END 1

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001577 001578 001579 001580 001581 001582 001583 001588 SUBROUTINE PARAM(N1) THIS ROUTINE WRITES A LIST OF THE PARAMETERS ON PRINTER(N1=99) OR CRT(N1=98) INTEGER PMAX, RMAX, TYPE REAL LINE LOGICAL SINGLX, UNMAG COMMON/CIO/ ICR, IHSP, IGRAPH COMMON/CBETA/ BETA COMMON/CPARAM/ RMAX, AIL, EPSI, EPSE, NSPEC, DFUI, DFUE, EP, SINGLX, XL(10) NXI С COMMON/CBETA/ BÉTA COMMON/CBETA/ BÉTA COMMON/CPARAN/ RMAX,AIL,EPSI,EPSE,NSPEC ,DFUI,DFUE,EP,SINGLX,XL(10) NXL CCMMON/CRGAM/R,GAM,RE,GAME,R1,RE1,AEAI1 COMMON/CDFU/ DFU(4),DRÉL(4) COMMON/CDFU/ DFU(4),DRÉL(4) COMMON/TVTUNE/LPENON,LPENOFF,ITALICS,IWINK,INTENSE,IRIGHT,IUP COMMON/DIS/GF,KYSG,KOSG,EPSF(4), MASS,KYAI,KOAI,MASG,PMAX,GMPSG ,AEAI,KMAX(4),AMASG(4),ÉPS(4),KMIN(4) COMMON/CMG/UNMAG IF(N1,EG.98) N = IGRAPH IF(UNMAG) TYPE = 6H IF(UNMAG) TYPE = 6H IF(UNMAG) TYPE = 6H LOCAL IF(NSPEC.GT.O) GG TG 36 IF(N.EG.IGRAPH) CALL SETCH(1,6.1,0,1,0,0) PMAX = 0 MEANS WE ARE USING THE LOCAL APPROXIMATION IF(FMAX.EG.O) GG TG 4 WRITE(N.3) TYPE,PMAX,AIL,OMPSG,EPSI,EPSE,MASS,AEAI,DFUI,DFUE 3 FORMAT(A6, 6H PMAX=, 12, 5H AIL=, F6.3, 7H OMPSG=,F8.1, 6H EPSI=, F6.3, 6H EPSE=, F6.3/ 6H MASS=, F9.2, 7H AE/AI=, F7.4, 6H DFUI=, F6.4,6H DFUE=, F6.4) GG TG 5 G TG TG 5 G TG TG 7 WRITE(N.6) OMPSG,EP,MASS,AEAI,DFUI,DFUE,BETA 6 FORMAT(14H LOCAL APPROX,9H CMPSG=,F9.3,8H EPS = ,F7.3, .8H MASS = ,F9.3/9H AE/AI = ,F8.4,7H DFUI =,F6.4,7H DFUE =,F6.4, .6H DFUE=,F6.4) GG TG 5 7 WRITE(N,6) TYPE,AIL,GMPSG,MASS,AEAI,DFUI,DFUE,SETA 6 FORMAT(14H LOCAL APPROX,9H CMPSG = ,F9.3,8H EPS = ,F7.3, .8H MASS = ,F9.3/9H AE/AI = ,F8.4,7H DFUI =,F6.4,7H DFUE =,F6.4, .6H DFTA=,F5.2) .7WITE(N,8) TYPE,AIL,GMPSG,MASS,AEAI,DFUI,DFUE,(XL(1),1=1,NXL) 8 FORMAT(5H SINE,A6,5H AIL=,F7.4,7H OMPSG=,F7.1,6H MASS=,F7.1,7H AE/ .AI=,F5.3,6H DFUI=;F6.4/6H DFUE=,F6.4,5H X/L=,10F4.2) 001585 001586 001588 001588 001589 001599 001599 001593 001593 001593 001593 001594 001595 001596 001597 001598 001597 001598 001602 001602 001604 001605 001605 001605 С 001608 001609 001610 001611 001612 001613 001614 001615 001616 ų

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001617	5	CONTINUE
001610		
001620	20	WRITE(N, SU) Eadmat/seu ding distribution de tong
001620	30	Contra (20 RING DISTRIBUTION OF TONS)
001622	10	
001622	19	
001623	22	WRITE(N, 32) Erdmat(184 Mayustitan 1865)
001625	56	CONTATION NAWELLIAN TONS)
001626	21	
001627	24	WRITE(N, 24) R, GAN Frankt(184 Middad Batta - 57 a su camma- 56 a cu 560 tano)
001626	23	CONTAILING MIRROR RATIO = ,F7.2, ON GAMMA= ,F6.3, 9H FOR IONS)
001620	29	
001630		
001631	26	WATTERN (20) Eadmattist cald Electrones
001632	20	CA TA 27
001633	25	
001634	20	
001635	29	TATILLY (29) FARMAT(3)H RING DISTRIBUTION OF ELECTRONS)
001636	23	A TA 27
001637	28	IF (BAME GT 0) GO TO 33
001638		WRITE(N 34)
001639	34	FORMAT (21H MAXWELLIAN ELECTRONS)
001640	• •	GC TC 27
001641	33	WRITE(N.35) RE.GAME
001642	35	FORMAT(16H MIRROR RATIO = .F7.2.9H GAMMA = .F6.3.14H FOR ELECTRONS
001643		·····
001644	27	CONTINUE
001645		RETURN
001646	36	IF(N.EQ.IGRAPH) CALL SETCH(30.,7.,1,0,1,0,0)
001647		IF(PMAX.EQ.O.AND.SINGLX) WRITE (N,37) BETA
001648	37	FORMAT(22H LOCAL APPROX., BETA =, F5.2)
001649		IF(N.EQ.IGRAPH) CALL SETCH(1.,6.,1,0,1,0,0)
001650		IF (PMAX.EQ.O.ANDNOT.SINGLX) WRITE (N,41) TYPE, AIL, EPSI, EPSE,
001651		(XL(I), I=1, NXL)
001652	41	FORMAT(5H SINE, A6, 5H AIL=, F7, 4, 6H EPSI=, F6, 3, 6H EPSE=, F6, 3/5H X/L=
001653		,10F4.2)
001654		IF(PMAX.GT.O) WRITE (N, 40) PMAX, AIL,EPSI, EPSE
001655	· 40	FORMAT(6H_PMAX=,13,5H AIL=,F7.4,6H EPSI=,F6.3,6H EPSE=,F6.3)
001656		WRITE(N, 42)
001657	42	FORMAT ("SPECIES GYRORADIUS MASS DIFFUSION J DENSITY DENS. GRA
001658	•	D. OMPSQ GYROFREQ.")
001659		NSPECI_= MINO(NSPEC, 4)
001660		IF(N1.EQ.99) NSPEC1 = NSPEC
001661		100 38 1 = 1, NSPEC1
001662		WP2 = OMPSQ*DENSE(I)*CHARG(I)**2/AMASS(I)
001663		WC = CHARG(I)/AMASS(I)
001664		WKILE(N, 39) I, KLAKM(I), AMASS(I), DFU(I), JAY(I), DENSE(I), EPS(I),
001665	· - '	WP2,WC
001666	39	FORMAI(15, 3E10, 2, 15, 4E10, 2)
001667	38	
001668		
001669		END

001670		
001671		SUBROUTINE RHCALC(R.GAM.KOA.RHF)
001672	С	CALCULATES THE RATIO OF THE ACTUAL PARTICLE DENSITY TO THE OUTDING
001673	Č	CENTER DENSITY AND STOPES IT IN BUE DIE THE MIDDOW DATE OUDING
001674	č	IS USED TO INDICATE A BING DISTRIBUTION IS THE MIRROR RATIO (R.L.)
001675	č	WHICH THE LASS CANE IS END TO TOT TO YOUR TO THE DEGREE TO
001676	Ŭ	THE LUSS CONCERS EMPTY, AND KUA IS KU*LARMOR RADIUS.
001677		
001676		$r_{\rm NOSM} = r_{\rm NOA} + r_{\rm NOA}$
001070	•	
001079	C	RING DISTRIBUTION
001680		CALL QCALC(KOSQ, N, N, 1)
001681		CALL BESSJ(KOA, N, J2)
001682		RHE = J2(1)
001683		RETURN
001684	С	MAXWELLIAN OR LOSS-CONE DISTRIBUTION
001685		1 RHE = EXP $(-5*KOSQ)$
001686		IF (GAM. GT. O.) GO TO 2
001687		RETURN
001688	Ċ	LOSS-CONE DISTRIBUTION
001689	-	2 KOSOR = KOSO/R
001690		RHF = (R*RHF - GAM*FYR(- 5*KOSOR))/(R - GAM)
001691		RETURN
001692		FND

COMPLEX DISP FUNCTION DISP (OMEGA) THIS ROUTINE CALCUL ATES THE DISPERSION FUNCTION COMPLEX GREGA.CMSG.DETC.DC.DI.AC20.20.AP.BP.APP.BPP. 1 DEFOLE.SUN.SUM2.DIV.DENOM.DN.PH1(18,22).CMEGA1.W.OMEGD COMPLEX MAXU2.RING2.TERM(4) INTEGER P.PMAX.GMAX. REAL MASS.COAD.KOSG.KYAI,KYSG.MASG.KPSG.KXSG LCAL UNMAG COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CBETA/ BETA COMMON/CAT / KYSG.KOSG.EPSF(4) CALL,KMAX(4), AMASG(4),EP(4), MASS.KYAI,KOAI,MASG.PMAX,GMPSG AEAI,KMAX(4), AMASG(4),EP(4),BP(4), GMEGA1,APP(4),BPP(4) COMMON/CATPL/A. NETSV.PHI COMMON/CATPL/A. NETSV.PHI COMMON/CATPL/A. NETSV.PHI COMMON/CATPL/A. NETSV.PHI COMMON/CATPL/A. NETSV.PHI COMMON/CASEA/ MARSS(4),CHARG(4),JAY(4),DENSE(4),NSPEC COMMON/CASS/VINMAG COMSS = DMEGA/MESA COMMON/CASS/VINM 001693 001694 001696 001697 001698 001698 001698 001700 001701 001702 001703 001704 001705 001706 001707 001708 001709 001711 001712 C 001713 001714 001715 001715 001717 001718 001717 001729 001722 001722 001723 001725 001726 001725 001726 001727 001728 001729 С 001730 001731 001732 001733 001734 001735 001736 001737 001738 001738 001740 С DC 2 LE - (1., NSPEC DC 2 I = 1, NSPEC CMEGD = CMÉGA + (0, 1.)*KYSQ*DFU(I) IF(KOAI.EQ.O. AND.PMAX.GT.O) GC TC 23 IF(EP(I).EQ.O.AND.PMAX.EQ.O) GC TC 23 IF(I.EQ.1) GC TC 22 IF(PMAX.EQ.O.AND.EP(I-1).EQ.O.) GC TC 23 IF(DFU(I).EQ.DFU(I-1)) GC TC 23 22 DEPOLE = CMEGD*DEPOLE 23 IF(KMAX(I).EQ.O) GC TC 2 KMX = KMAX(I) JI = KMIN(I) IF(PMAX.GT.O) GC TC 11 DC 1 J = JI,KMX 001740 001741 001742 001743 001743 001743 001746 001746 001747 001748 001749 001750 001751 001752 001752 DEPENDENT OF THE SECONDENSION OF THE SECONDENSIAN OF THE SECONDES 1 001754 Ĵ = JI, KMX _= depôle*(0msq - J*J/Amasq(I)) 11 DEPOLE 001757 001757 001759 001759 001760 001760 001762 001763 001763 001766 001766 001766 001766 001766 001776 001777 001777 001777 001777 001777 001777 001778 001780 001788 001788 001788 001788 001788 001788 001788 001788 001788 001788 001788 001788 001788 001788 001788 001768 001768 001767 001778 001777 001777 001777 001777 001777 001777 0017777 001777 001777 001777 001777 001777 001777 001777 001776 001776 001776 001776 001776 001776 001776 001777 001778 0001778 0001780 12 CONTINUE THE FACTOR QF IS PUT IN TO MAKE DEPOLE ON THE ORDER OF 1 DEPOLE = DEPOLE*QF IF (PMAX.GT.O) DEPOLE = DEPOLE*QF 2 С IF(PMAX.EQ.0) GO TO 10 CALCULATE DISP USING SINUSCIDAL DENSITY PROFILE IF PMAX.GT.O P=-PMAX-1 M=1 NP = PMAX+1 THE LOOP DENS IN THE DESCRIPTION OF THE DESCRIPTION С NP = PMAX+1 THE LOOP OVER ROWS IN THE DETERMINANT STARTS HERE N IS THE INDEX OF THE ROWS, RUNNING FROM 1 TO 2PMAX +1 M IS THE COLUMN INDEX, WITH VALUES N-1,N,N+1 Ç Ĉ M IS THE COLUMN INDEX, WITH VALUES N-1,N,N+1 DO 9 N=1,NP P=P+1 THIS SKIPS THE 1,0 ELEMENT WHICH IS NOT NEEDED IF(N. EQ. 1) GO TO 40 THIS IS THE CALCULATION OF THE M=N-1 OR LEFT ELEMENTS A(N,M) = 0. DO 3 I = 1, NSPEC ARH(N,M,I) = EPSF(I)*(APP(I) + BPP(I)) A(N,M) = A(N,M) + ARH(N,M,I) A(N,M) = A(N,M)*DEPOLE M=M+1 THIS IS THE CALCULATION OF THE DIAGONAL ELEMENTS CONTINUE С C 3 M=M+1 THIS IS THE CALCULATION OF THE DIAGONAL ELEMENTS CONTINUE CALL ABCALC A(N,M) = KYSQ + P*P*KOSQ DO B [= 1, NSPEC ARH(N,M,I) = OMPSQ*AP(I)*CHARG(I)**2*DENSE(I)/AMASS(I) A(N,M) = A(N,M) + ARH(N,M,I) A(N,M) = A(N,M) + ARH(N,M,I) A(N,M) = A(N,M)*DEPOLE M=M+1 THIS IS THE CALCULATION OF THE M=N+1 OR RIGHT ELEMENTS IF(N, EQ, NP) GO TO 60 CALL ABCALC Ŧ С 40 8 001788 001789 001790 С IF(N .EQ. NP) GO TO 60 CALL ABCALC 001791 001792

001793 001794 001795 001796 001797 001798 001799		5	A(N,M) = 0. D0 5 I = 1, NSPEC ARH(N,M,I) = EPSF(I)*(AP(I) - BP(I)) A(N,M) = A(N,M) + ARH(N,M,I) A(N,M) = A(N,M)*DEP0LE M=M-1 O(M) = A(N,M)*DEP0LE
001799 001800 001801 001802 001803 001803 001805	С	9	CONTINUE CONTINUE WE MAKE USE OF THE FACT THAT A(P,P') = A(-P',-P) N1 = NP NN1 = N1 NP = 2*PMAX IF(NP.LT.NN1) GO TO 104 DO 20 N = NN1 NP
001808 001807 001808 001809 001810 001811 001812		21	DO 20 n = NN1, NFA(N, N+1) = A(N1-1, N1)A(N+1, N) = A(N1-1, N1-1)A(N+1, N+1) = A(N1-1, N1-1)DO 21 I = 1, NSPECARH(N, N+1, I) = ARH(N1-1, N1, I)ARH(N+1, N+1, I) = ARH(N1-1, I)ARH(N+1, N+1, I) = ARH(N1-1, I)
001814 001815 001816 001817 001818 001819 001820	С	20 104	NI = N1-1 THE DETERMINANT OF THE COEFFICIENTS, DETC, IS DONE HERE IF(.NOT.DTCALC) RETURN DETC=A(1,1) DN = 1, IF(PMAX.EQ. 0) GO TO 70 DO 7 N=2,NPP
001821 001822 001823 001824 001825 001825	С	7 70	DO=DN DN=DETC DETC=A(N,N)*DN-A(N,N-1)*A(N-1,N)*DO DISP=DETC/OMEGA RETURN
001827 001828 001829 001830 001831 001832	č	10 10	ALCULATE DISP USING THE LOCAL APPROXIMATION IF PMAX,EQ.O KPSQ = KYSQ + KXSQ DO 15 I = 1, NSPEC OMEGD = OMEGA + (0,,1.)*KYSQ*DFU(I) IF(DENSE(I),EQ.O.) GO TO 110 FPS = FP(I)
001833 001834 001835	С	SI	IF(RLARM(I).GT.0.) GØ TØ 16 PECIES I IS COLD TERM(I) = -KPSQ - (1.+.5*BETA)*EPS*KYAI*CHARG(I)/(AMASS(I)*OMEGD) OF TERM(I) = -KPSQ - (1.+.5*BETA)*EPS*KYAI*CHARG(I)/(AMASS(I)*OMEGD)
001836 001837 001838 001839		16	IF(JAY(I),GE.O) GO TO 17 TERM(I) = RING2(OMEGA,KYAI,EPS,RLARM(I),AMASS(I),CHARG(I),NMAX(I) ,QMAX(I),OMEGD)
001840 001841 001842 001843 001844 001845 001845	CCC	17 EI EI I	GO 10 10 EPSI = 2.*EPSF(I)*AMASS(I)/(CMPSQ*CHARG(I)**2*DENSE(I)*DREL(I)) PSI IS THE EPSI OR EPSE (DEPENDING ON WHETHER SPECIES I IS IONS OR LECTRONS) USED IN *ROOTS*, DIVIDED BY DREL (=N(X)/NO). IT IS NEEDED N *MAXW2* WHEN WE ARE USING THE HIGH GROWTH RATE SINUSCIDAL DISP. REL. TERM(I) = MAXW2(CMEGA,KYAI,EPSI,KPSQ,DREL(I),CHARG(I),JAY(I), NMAX(I),GMAX(I),CMEGD,KOAI,EPSI,KPSQ,DREL(I))
001847 001848 001849 001850		18 13 110	TERM([]) = TERM([])/RLARM([])**2 TERM([]) = TERM([])*AMASS([])*DENSE([])*DREL([]) GÖ TÖ 15 TERM([]) = 0. CONTINUE
001852 001853 001853 001854 001855 001856 001857		19	DISP = KPSQ/OMPSQ + ,5*BETA DO 19 I = 1, NSPEC DISP = DISP - TERM(I) DISP = DISP*DEPOLE RETURN END

001861 CGMMGN/BSTGRE/ BJ(2000) 001862 CGMPLEX GMEGA, W, SUM1, SUM2, WD, GMEGD 001863 INTEGER GMAX 001863 INTEGER GMAX 001864 REAL KYAI, KY, KO, KOAI, KOSQ 001865 GMEGA = W*AMASS/CHARG 001866 GMEGD = WD*AMASS/CHARG 001866 GMEGD = WD*AMASS/CHARG 001866 GMEGD = WD*AMASS/CHARG 001867 KYAI = KY*RLARM 001868 EPS = E*RLARM 001869 CALL BESSJ(KYAI, NMAX, BJ) 001870 SUM1 = D, 001871 IF(EPS.EQ.O.) GØ TØ 3 001872 SUM1 = BJ(1)*BJ(1)/ØMEGD 001873 3 SUM2 = O. 001874 DØ 1 I = 1, GMAX 001875 AI = I 001876 SUM1 + BJ(I+1)*BJ(I+1)*(1./(ØMEGD-AI) + 1./(ØMEGD+AI)) 001876 JUM2 = SUM2 + BJ(I+1)*(BJ(I) - BJ(I+2))*AI*(1./(ØMEGD-AI) - 1./ 001878 .(ØMEGD+AI)) 001879 SUM2 = SUM2*(1. + ØMEGA*EPS/KYAI) 001881 RING2 = SUM1 + SUM2 001882 RETURN

COMPLEX MAXW2 FUNCTION MAXW2(W,KY,E,RLARM,AMASS,CHARG,JAY,NMAX,QMAX,WD, KO,EPSI,KP,DREL) COMPLEX OMEGA,W,SUM1,SUM2,TP,TN,AI,WD,OMEGD COMPLEX IXKO,VP,VPPVD,VPMVD,ZEE COMPLEX CMIX,CPIX,ZVP,ZVPP,ZVPM INTEGER QMAX REAL KYAI,KY,KYSQ,KO,KOAI,KOSQ,KP,KPSQ,KPAI LOGICAL UNMAG,SNGLX COMMON/CXK/XKO COMMON/CXK/XKO COMMON/CSNGLX/ SNGLX OMEGA = W*AMASS/CHARG OMEGA = WTAMASS/CHARG CMEGA = WTAMASS/CHARG CMEGA = WTAMASS/CHARG CMEGA = KYAI*KYAI EPS = E*RLARM 001888 001889 001890 001891 001892 001892 001893 001894 001895 001896 CO1897 CO1898 CO1898 CO1898 CO1901 CO1902 CO1902 CO1902 CO1903 CO1903 CO1903 CO1903 CO1904 CO1905 CO1906 CO1907 CO1908 CO1907 CO1908 CO1907 CO1911 CO1911 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1907 CO1908 CO1908 CO1907 CO1908 CO1917 CO1922 CO IF (UNMAG) GO TO 4 CALL BESSI (KYSQ, NMAX, BI) SUM1 = 0. IF (EPS.EQ.0.) GO TO 3 SUM1 = BI(1)/OMEGD 3 SUM2 = 0. 3 SUM2 = 0. D0 2 I = 1, QMAX AI = CMPLX(FLGAT(I),0.) TP = BI(I+1)/(CMEGD - AI) TN = BI(I+1)/(CMEGD + AI) SUM2 = SUM2 + AI*(TP - TN) 2 SUM1 = SUM1 + TP + TN SUM2 = SUM2*(1. + CMEGA*EPS/KYAI) SUM1 =-SUM1*EPS*KYAI MAXW2 = SUM1 + SUM2 RETURN SUM1 =-SUM1*EPS*KYAI MAXW2 = SUM1 + SUM2 RETURN KOAI = KO*RLARM KOSQ = KOAI*KOAI KPSQ = KP*RLARM*RLARM KPAI = SQRT(.5)*0MEGD/KPAI IF(.NOT.SNGLX) GO TO 5 MAXW2 = -1. - OMEGA*EPS/KYAI - SQRT(.5)*(-EPS+ OMEGD*(1.+OMEGA*EPS /KYAI)/KYAI)*ZEE(VP) RETURN IXKO = CMPLX(0.,XKO) VDDOTK = KOAI*KYAI/KPAI VPPVD = SQRT(.5)*(OMEGD/KPAI + CMPLX(0.,VDDOTK)) VPMVD = SQRT(.5)*(OMEGD/KPAI - CMPLX(0.,VDDOTK)) CPIX= CEXP(1XKO) CVP = ZEE(VP) ZVPP = ZEE(VP) ZVPP = ZEE(VPVD) ZVPP = ZEE(VPVD) ZVPM = 4 001924 001925 001925 001927 001928 001929 001931 001933 001933 001933 001935 001935 001935 001935 001938 001938 001938 5 EXPL*(,5*CMIX*VPPVD*ZVPP + ,5*CPIX*VPMVD*ZVPM) -1, - MAXW2*EPSI - VP*ZVP/DREL 001940 001941 MAXW2 = RETURN ËND 001942 001943 001944 001945 001946 001947 001948 001949 001959 001952 001955 001955 001955 001955 001955 001955 001955 001959 001959 SUBROUTINE QFKCAL(QF,KMAX,KMIN,RMX,RMIN) LOGICAL UNMAG DIMENSION KMAX(4), KMIN(4) COMMON/CSPEC/ RLARM(4),AMASS(4),CHARG(4),JAY(4),DENSE(4),NSPEC COMMON/CMAG/UNMAG Q = 1. DØ 11 I = 1, NSPEC IF(UNMAG) GØ TØ 87 IF(DENSE(I).EQ.O.) GØ TØ 87 IF(I.EQ.1) GØ TØ 91 DØ 85 J = 2, I 85 IF(AMASS(I).EQ.AMASS(J - 1).AND.DENSE(J-1).NE.O.) GØ TØ 87 91 KMAX(I) = RMX*AMASS(I) KMIN(I) = (RMIN - 2.)*AMASS(I) IF(KMIN(I).LT.1) KMIN(I) = 1 GØ TØ 88 87 KMAX(I) = 0 88 IF(KMAX(I).EQ.0) GØ TØ 11 KMX = (KMAX(I) - KMIN(I))/2 + 1 DØ 86 K = 1, KMX 86 Q = FLØAT(K)*Q/AMASS(I) 11 CØNTINUE QF = 1./(Q*Q) 001960 001961 001962 001963 001964 001965 001965 QF = 1.7(Q*Q) RETURN I 001967 END 001968

001884 001885 001887

001969 001970 001972 001973 001974 001975 001976 001976 001976 001980 001980 001982 001983 001984 001985 001985 001988 001989 001999 001993 001993 001995 001995 001995 001995 001995 001999 001999 001999 001999 001999 001995	<pre>SUBROUTINE QCALC(CHI, QMAX, NMAX, N) INTEGER QMAX COMMON/CQN/MINQ F(N, GT, O) GO TO 1 IF(N, GT, O) GO TO 1 C THIS IS USED FOR A MAXWELLIAN DISTRIBUTION OF IONS. C QMAX IS MADE LARGE ENOUGH SO THAT MODIFIED BESSEL FUNCTIONS OF INDEX G GREATER THAN QMAX ARE LESS THAN 1,E-O3 C THE PARAMETER NMAX USED IN THE SUBROUTINE BESSI IS CHOSEN SO THAT IF(CHI.E.64.) AMAX = 28. IF(CHI.GT.64.) AMAX = 28. IF(CHI.GT.64.) AMAX = 3.5*SQRT(CHI) NMAX = AMAX. C QMAX AND NMAX MUST BE SIGNIFICANTLY MORE THAN THE HIGHEST CYCLOTRON C MAX AND NMAX MUST BE SIGNIFICANTLY MORE THAN THE HIGHEST CYCLOTRON C HARMONIC SOUGHT. IF(NMAX,LT.MINQ) NMAX = MINQ C IF NMAX IS TOO GREAT BESSI WILL OVERFLOW. IF(NMAX,LT.MINQ) NMAX = 1988 MAX = NMAX RETURN C THIS IS USED FOR RING DISTRIBUTIONS OF IONS. C QMAX AND E LARGE ENOUGH SO THAT BESSEL FUNCTIONS OF INDEX G QMAX = 10 IF(CHI.GT.4.) QMAX = SQRT(CHI) + 6.5*CHI**.166667 I QMAX = 10 IF(CHI.GT.4.) QMAX = SQRT(CHI) + 11,*CHI**.166667 IF(NMAX.GT. 398) NMAX = 398 IFF(CHI.GT.9.) NMAX = 398 IFF(CHI.GT.9.) NMAX = 398 IFF(CHI.GT.9.) NMAX = SQRT(CHI) + 11,*CHI**.166667 IF(NMAX.GT. 398) NMAX = 398 IFF(CHI.GT.9.) NMAX = SQRT(CHI) + 11,*CHI**.16667 IF(NMAX.GT. 398) NMAX = 398 IFF(CHI.GT.9.) NMAX = SQRT(CHI) + 11,*CHI**.16667 IF(NMAX.GT.MAX - 2) QMAX = NMAX - 2 RETURN END</pre>	
002001 002002 002003 002005 002006 002009 002009 002010 002011 002012 002013 002013 002015 002016 002016 002017 002018 002019 002020 002023	SUBROUTINE BESSI(X,N,BI) DIMENSION BI(2000) N1 = N+1 IF (X.LT.1.E-14) GC TO 4 BI(N+2) = 0. BI(N+1) = 1. DC 1 I = 1, N M = N - I + 1 AM1 = M 1 BI(M) = (2.*AM1/X)*BI(M+1) + BI(M+2) SUM = BI(1)/2. DC 2 M = 2, N1 2 SUM = SUM + BI(M) SUM = SUM + BI(M) SUM = SUM + SUM DC 3 I = 1, N1 3 BI(I) = BI(1)/SUM GC TO 6 4 BI(1) = 1. DC 5 I = 2, N1 5 BI(I) = 0. 6 RETURN END	
002024 002025 002027 002028 002029 002030 002031 002032 002033 002033 002035 002035 002036 002038 002038 002038 002040 002041 002041 002044 002044 002044 002044 002044 002045 002047	SUBROUTINE BESSJ (X,N,BJ) DIMENSION BJ(400) N1 = N+1 IF (X.LT.1.E-14) GO TO 4 BJ(N+2) = 0. BJ(N+1) = 1. DO 1 I = 1, N M = N - I + 1 AM1 = M 1 BJ(M) = (2.*AM1/X)*BJ(M+1) - BJ(M+2) SUM = BJ(1)/2. N2 = N/2 DO 2 M = 1, N2 2 SUM = SUM + BJ(M+M+1) SUM = SUM + SUM DO 3 I = 1, N1 3 BJ(I) = BJ(I)/SUM GO TO 6 4 BJ(I) = 1. DO 5 I = 2, N1 5 BJ(I) = 0. 6 RETURN END	

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002054 002055 002055 002057 002059 002060 002061 002062 002063 002063 002063 002064 002065 002066 002068 002068 002068 002068	c cc	COMPLEX ZEE FUNCTION ZEE(X) COMPLEX X, TERM DOUBLE PRECISION RZED, IZED, RTERMD, ITERMD, RTNEW, RX2, IX2 DATA PI/3.14159265359/ ABX2 = CABS2(X) IF ABS(X) IS GREATER THAN 5.7, USE THE ASYMPTOTIC FORM. IF(ABX2.GT.33.) GO TO 4 ZEE = (0.,1.)*SQRT(PI)*CEXP(-X*X) IF(ABX2.GT.1.) NMAX = 2.8*ABX2 + 19. IF(ABX2.LE.1.) NMAX = 11.*ABX2 + 10. TERM = 2.*X IF ABS(X) IS GREATER THAN 2.6, WE MUST USE DOUBLE PRECISION IF ZEE IS TO BE ACCURATE TO 1.E-12. IF(ABX2.GT.7.) GO TO 2 DO 1 N = 1.NMAX
002071 002072 002073 002075 002075 002076 002078 002078 002081 002081 002083 002083 002085 002085 002085 002085 002087 002089 002091 002093 002093 002093 002095 002095 002095		<pre>ZEE = ZEETERM 1 TERM = -TERM*2.*X*X/(2*N + 1) RETURN 2 RX2 = REAL(X*X) IX2 = AIMAG(X*X) RZED = REAL(ZEE) IZED = REAL(ZEE) IZED = REAL(TERM) DO 3 N = 1, NMAX RZED = RZED - RTERMD IZED = IZED - ITERMD IZED = IZED - ITERMD RTNEW = -2.*(RX2*ITERMD + IX2*ITERMD)/(2*N +1) ITERMD = -2.*(RX2*ITERMD + IX2*RTERMD)/(2*N + 1) 3 RTERMD = RTNEW RZEE = RZED AZEE = IZED ZEE = CMPLX(RZEE, AZEE) RETURN 4 NMAX = 20 TERM = 1./X ZEE = 0. D0 5 N = 1, NMAX ZEE = ZEE_ TERM 5 TERM = FLOAT(2*N-1)*TERM/(2.*X*X) RETURN END</pre>
002098 002100 002101 002102 002103 002105 002105 002105 002106 0021105 0021106 0021112 00021112 00021114 00021114 00021115 00021121 00021121 000212123 00021223 00021226 00021230 0002131 0002133 0002133	CC	<pre>SUBROUTINE ABCALC THIS IS AN INTERFACE BETWEEN DISP AND THE ROUTINES MAXW, RING, AND COLD WHICH CALCULATE THE ALPHAS AND BETAS. DIMENSION CHIP1(4) COMMEA. (ACLOUDER THE ALPHAS AND BETAS. DIMENSION CHIP1(4), AP(4), BPP REAL J1(400,4) COMMEA. (ACLOUDER ALMAND, APP, BPP REAL J1(400,4), AP(4), BP(4), CMEGA, APP(4), BPP(4) COMMEAN/CCHIP/CHIP(4), AP(4), BP(4), CMEGA, APP(4), BPP(4), COMMEAN/CCHIP/CHIP(4), AP(4), BP(4), CMEGA, APP(4), BPP(4), COMMEAN/CCHIP/CHIP(4), AP(5), BP(6), APP(1), DENSE(4), NSPEC DE 5 1 = 1, NSPEC IF(DENSE(1), EQ.C.) GE TO 6 W = OMEGA*AMASS(1)/CHARG(1) F(RLARM(1), EQ.C.) GE TO 7 IF(LARM(1), EQ.C.) GE TO 7 IF(JAY(1), GE.O) CALL MAXW(W, RLARM(1), AP(1), BP(1), APP(1), BPP(1), CHIP(1), CHIP1(1), J1(1, 1), J2(1, 1), 1) FACTOR = (AMASS(1)/CHARG(1)*RLARM(1)))**2 AP(1) = AP(1)*FACTOR BP(1) = APP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = D(1)*FACTOR BPP(1) = D(1)*FACTOR GE TO 5 GATE COLD(W, AP(1), BP(1), APP(1), BPP(1), CHIP(1), CHIP1(1)) FACTOR = (AMASS(1)/CHARG(1))**2 AP(1) = APP(1)*FACTOR BPP(1) = D. BPP(1) = D. BPP(1) = D. BPP(1) = D. BPP(1) = BP(1)*FACTOR BPP(1) = D. BPP(1) = BP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = BP(1)*FACTOR BPP(1) = BP(1)*FACTOR BPP(1) = BP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = BP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = BP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = BP(1)*FACTOR BPP(1) = APP(1)*FACTOR BPP(1) = APP(1)*FACTOR BP(1) = APP(1)*FACTOR BP(1) = APP(1)*FACTOR BP(1) = APP(1)*FACTOR BP(1) = APP(1)*FACTOR APP(1) = APP(1)*FACTOR BP(1) = APP(1)*FACTOR APP(1) = APP(1)*FACTOR APP(</pre>

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SUBRGUTINE MAXW(GMEGA, RLARM, AP, BP, APP, BPP, LAMDP, LAMDP1, K)
THIS ROUTINE CALCULATES ALPHAS AND BETAS FOR A MAXWELLIAN DISTRIBUTI
DIMENSION IPP(2000)
CGMPLEX GMEGA, UMEGA, OMSG, AP, BP, DIV, DENOM, SUM1, SUM2, VP, ZEE
CGMPLEX APP, BPP, EIAP, EIAP, EIGAP, SUM3, SUM4, VPM, VPP, ZVPP, ZVPM
REAL MASS, KOAI, KOSG, KYAI, KYSG, MASG, LAMDP, LAMDP1, LAM, IPP
INTEGER PMAX, P, GMAX
LOGICAL UNMAG
CCMMON/DIS/GF, CYSG, COSG, EPSF(4), MASS, CYAI, COAI, MASG, PMAX, OMPSG
, AEAI, KMAX(4), AMASG(4), EPS(4), KMIN(4)
CCMMON/CRBSJ/GMAX(4), NMAX1(4)
COMMON/CMAG/ UNMAG
NG = GMAX(K) + 1
NMAX = NMAX1(K)
KYAI = CYAI*RLARM
KYSG = KYAI*RLARM
KOSG = COAI*RLARM
KOSG = COAI*RLARM
KOSG = COAI*RLARM
KOSG = GMEGA*OMEGA
IF(M, NE.N) GG TG 2
IF(N, NE.N) GG TG 2
IF(N, NE.N) GG TG 2
IF(N, NE.N) GG TG 7
CALL BESSI(LAM, NMAX, IPP)
SUM1 = 0.
DO 5 K1=2, NG
GENGM = (1., 0.)/(CMSG - G*G)
5 SUM1 = SUM1+ IPP(K))*DENOM*G*G
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 D0 5 K1=2,NQ

Q=K1-1

DENOM = (1.,0.)/(CMSQ - Q*Q)

5 SUM1= SUM1+ IPP(K1)*DENOM*Q*Q

AP = -2.0*SUM1

BP= 0.

APP = 0.

RETURN

7 VP = CMEGA/SQRT(2.*LAM)

AP = 1. + VP*ZEE(VP)

BP = 0.

APP = 0.

BPP = 0.

RETURN

2 IF(M.NE.N+1) GO TO 3

LAM = LAMDP1 * LAMDP

EXPL = EXP(-.5*(LAMDP1 - LAMDP)**2)

IF(UNMAG) GO TO 8

EIAP = (KYSQ + P*(P+1)*KOSQ -(0.,1.)*KOAI*KYAI)/LAM

CALL BESSI(LAM,NMAX,IPP)

SUM1 = 0.

SUM2 = IPP(1)/CMEGA

SUM3 = 0.
                                                                                              ς.
                                                                                                                                   SUM2 = IPP(1
SUM3 = 0,
SUM4 = SUM2
EIQAP = 1.
DO 6 K1=2,NQ
Q=K1-1
EIQAP = EIQA
SUM1 = SUM1
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                                                                                                                                  Q=K1-1
EIQAP = EIQAP*EIAP
SUM1 = SUM1 + Q*IPP(K1)*(EIQAP/(OMEGA-Q) - 1./(EIQAP*(CMEGA+Q)))
SUM2 = SUM2 + IPP(K1)*(EIQAP/(OMEGA-Q) + 1./(EIQAP*(CMEGA+Q)))
SUM3 = SUM3 + Q*IPP(K1)*(1./(EIQAP*(CMEGA-Q)) - EIQAP/(CMEGA+Q))
SUM4 = SUM4 + IPP(K1)*(1./(EIQAP*(CMEGA-Q)) + EIQAP/(CMEGA+Q))
AP = -EXPL*SUM1
BP = (0.,1.)*KOAI*KYAI*EXPL*SUM2
APP = -EXPL*SUM3
BPP = (0.,1.)*KOAI*KYAI*EXPL*SUM4
RETURN
                                                                                                                     6
                                                                                                                          AP = (0,1)*KOAI*KTAI*L

APP = -EXPL*SUM3

BPP = (0,1)*KOAI*KTAI*L

SQ2LM = SQRT(2.*LAM)

VP = CMEGA/SQ2LM

VPM = (CMEGA + (0,1.)*KYAI*KOAI)/SQ2LM

VPP = (CMEGA + (0,1.)*KYAI*KOAI)/SQ2LM

VPP = ZEE(VPP)

ZVPM = ZEE(VPP)

ZVPM = ZEE(VPM)

EXPL = EXP(-.5*KOSQ)

AP = EXPL*(1. + VP*ZVPP)

BP = -(0.,1.)*KOAI*KYAI*EXPL*ZVPP /SQ2LM

APP= EXPL*(1. + VP*ZVPM)

BPP= -(0.,1.)*KOAI*KYAI*EXPL*ZVPM /SQ2LM

RETURN
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                                                                                                                                 \begin{array}{l} APF = EA\\ BPP = -(\\ RETURN\\ AP = 0\\ BP = 0\\ \end{array}
                                                                                                                     3
                                                                                                                                     APP
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SUBROUTINE RING (OMEGA, RLARM, AP, BP, APP, BPP, CHIP, CHIP1, J1, J2, K)
CALCULATES THE ALPHAS AND BETAS FOR A RING DISTRIBUTION.
DIMENSION J1(1), J2(1)
COMPLEX CMEGA, OMSG, AP, BP, DIV, DENOM, SUM1, SUM2
COMPLEX SUM3, SUM4, ELAP, ELGAP, APP, BPP
REAL MASS, KOAI, KOSQ, KYAI, KYSQ, MASQ, J1, J2, JP1, JP2
INTEGER PMAX, P, QMAX
COMMON/DIS/OF, CYSQ, COSQ, EPSF(4), MASS, CYAI, COAI, MASQ, PMAX, CN
, AEAI, KMAX(4), AMASQ(4), EPS(4), KMIN(4)
COMMON/APBP/N, M, P
COMMON/CRBESJ/GMAX(4), NMAX1(4)
NMAX = NMAX1(K)
NQ = OMAX(K) + 1
NQ1 = CYAI*RLARM
KYSQ = KYAI*KYAI
KOAI = COAI*RLARM
KOSQ = KOAI*RLARM
KOSQ = KOAI*RLARM
KOSQ = COAI*RLARM
KOSQ = COAI*RLARM
COMMON/CHEASJ(CHIP1, NMAX, J2)
CHIP = CHIP1
D0 7 K1 = 1, NO1
7 J1(K1) = J2(K1)
CALL BESSJ(CHIP1, NMAX, J2)
SUM1 = 0.
D0 6 5 K1=2, NQ
Q=K1-1
DENOM = (1, 0, )/(CMSQ-Q*Q)
JP1 = .5*(J1(K1-1) - J1(K1+1))
5 SUM1 = SUM1 + Q*Q*JP1*J1(K1)*DENOM
AP = -4.0*SUM1*CHIP
BP = 0.
002220
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                                                     C
002223
002224
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002227
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                                                                                                                                                                                                                                                                                        MASS, CYAI, COAI, MASQ, PMAX, CMPSQ
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 002253
                                                                                     SUM1 = SUM1 + G*G*J
AP = -4.0*SUM1*CHIP
BP = 0.
APP = 0.
BPP = 0.
002253
002254
002255
002256
002257
                                                                          AFF = 0.

BFP = 0.

RETURN

2 IF(M.NE.N+1) GO TO 3

EIAP = (KYSQ + P*(P+1)*KOSQ - (0.,1.)*KOAI*KYAI)/(CHIP*CHIP1)

SUM1 = 0.

SUM2 = J1(1)*J2(1)/OMEGA

SUM3 = 0.

SUM4 = SUM2

EIQAP = 1.

DO 6 K1=2,NQ

Q=K1-1

EIQAP = EIQAP*EIAP

JP1 = .5*(J1(K1-1) - J1(K1+1))

JP2 = .5*(J2(K1-1) - J2(K1+1))

SUM1 = SUM1 + Q*(CHIP*JP1*J2(K1) + CHIP1*JP2*J1(K1))*(EIQAP/

.(GMEGA-Q) - 1./(EIQAP*(OMEGA+Q)))

SUM2 = SUM2 + J1(K1)*J2(K1)*(EIQAP/(CMEGA-Q) + 1./(EIQAP*(CMEGA+Q)))

SUM2 = SUM2 + J1(K1)*J2(K1)*(EIQAP/(CMEGA-Q) + 1./(EIQAP*(CMEGA+Q)))
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                                                                                   .))
SUM3
                                                                            \dot{S}úM3 = SUM3 + Q*(CHIP*JP1*J2(K1) + CHIP1*JP2*J1(K1))*(1./(ElQAP* .(OMEGA-Q)) - EIQAP/(OMEGA+Q))
6 SUM4 = SUM4 + J1(K1)*J2(K1)*(1./(EIQAP*(OMEGA-Q))+ EIQAP/(OMEGA+Q))
                                                                                      )
AP = -SUM1
BP = (0.,1.)*KOAI*KYAI*SUM2
APP = -SUM3
BPP= (0.,1.)*KOAI*KYAI*SUM4
  002280
  002281
002282
002283
                                                                             RETURN
3 AP = 0,
BP = 0,
  002284
  002285
                                                                                        \begin{array}{rcl} \mathsf{APP} &= & \mathsf{O} \, . \\ \mathsf{APP} &= & \mathsf{O} \, . \\ \mathsf{BPP} &= & \mathsf{O} \, . \end{array}
  002287
                                                                                        RETURN
  002289
                                                                                         END
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002290		
002291		SUBROUTINE COLD COMEGA, AP, BP, APP, BPP, CHIP, CHIPT
002292	С	CALCULATES ALPHAS AND BETAS FOR COLD_ELECTRONS
002293	-	COMPLEX OMEGA. OMSQ. DIV. AP. BP, APP, BPP
002204		RFAL MASS KOAL KOSO KYAL KYSO MASO
002205		INTEGER PMAX P
002290		COMMONICIES COL KUSO KOSO FESE(4) MASS.KYAL.KOAL.MASQ.PMAX.OMPSQ
002290		$A = A + M A \times (A) + A M A \subseteq (A) + E \subseteq (A) + K = M A \subseteq (A) + K = M A \times (A) + A M A \subseteq (A) + E \subseteq (A) + K = M A \times (A) + K = M A $
002297		
002298		
002299		OMSU = OMEGA*OMEGA
002300		
002301		IF(N,EQ.1)_CHIP1 = SQRT(KYSQ + PMAX*PMAX*KUSQ)
002302		CHIP = CHIP1
002303		CHIP1 = SQRT(KYSQ + (P+1)*(P+1)*KOSQ)
002304		AP = -CHIP*CHIP/(GMSQ - 1,)
002305		BP = 0.
002306		APP = 0.
002307		BPP = 0
002308		RETURN
002000		2 COSAP = (KYSO + P*(P+1)*KOSO)/(CHIP*CHIP1)
002309		2 SUMAP = - (NOAL + (NAL) + (CHIP + CHIP))
002310		$\Delta P = -(1 - 1 + 1 + 1 + 1 + 1 + 1 + 1 + 1 + 1 + $
002311		AP = -CHIF*CHIF*(CUSAF + (0, 1,), "CMECA*SINAP)/(CMSQ + 1,)
002312		
002313		IF (RUALINE.U.) GO TO T
002314		BP_= 0.
002315		BPP = O,
002316		RETURN
002317		1 BP = KOAI*KYAI/OMEGA
002318		BP =CMPLX(-AIMAG(BP), REAL(BP))
002319		BPP = BP
002320		RETURN
002321		FND
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002322		
002323		SUBROUTINE SAVPHI
002324	C FI	NDS PHI(P)
002325		COMPLEX PHI(18,22) A(20,20) PHI1(19) PHI2(19)
002326		COMMON /PHIC/ A NRTSV PHI
002327		
002328		DATA NETSV/1/
002020		
002023		
002330		
002001	•	
002332	1	$F_{11}(1) = (1, 0, 0, 0)$
002333		IFU(NF-1)/2 .EQ. 0) GO (0 5300
002334		PHII(2) = -A(1,1)/A(1,2)
002335		UO 5301 K2 = 3,NP
002336		K1 # K2-1
002337		PHI1(K2) = -(A(K1,K1-1)*PHI1(K1-1)+A(K1,K1)*PHI1(K1))/A(K1,K2)
002338	5301	CONTINUE
002339	•	PHI2(NP) = PHI1(NP)
002340		PHI2(NP-1) = -A(NP, NP)*PHI2(NP)/A(NP, NP-1)
002341		DO 5302 K2 = 3.NP
002342		K1 = NP - K2 + 1
002343		PHI2(K1) = -(A(K1+1,K1+1)*PHI2(K1+1)+A(K1+1,K1+2)*PHI2(K1+2))
002344		/A(K1+1.K1)
002345	5302	CONTINUE
002346		DC = 5303  K1 = 1.NP
002347		PHI(NRTSV, K1) = 0.5*(PHI1(K1) + PHI2(K1))
002348	5303	CONTINUE
002349	5300	CONTINUE
002350		RETURN
002351		FND

002352 002353 002354	1	SUBROUTINE PAGE COMMON/CIO/ICR, IHSP, IGRAPH WRITE(IHSP, 1)
002356 002357	•	RETURN END

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SUBROUTINE MRAF(F,RT,CONV,NRTS,PP1,PP2,PP3,EP1,EP2,MXM) DIMENSION ARPHI (00), ABPHI(00) LOGICAL PHIZ(200), WARN LOGICAL CONV(200) LGGICAL PHIZ(200), WARN LGGICAL CONV(200) LGGICAL CONV(200) LGGICAL SW1,SW2,SW3,SW4,SW5,SWR,SW7,SWR1,NREDH,SW31,SW71 CGMPLEX PHI(18,22), A(20,20), PHJ CGMPLEX F,RT(200), Ph P1,PP2,P3,CHANGE CGMPLEX H,RGGT,FUNC,FRGGT CGMPLEX H,RGGT,FUNC,FRGGT CGMPLEX H,RGGT,FUNC,FRGGT CGMPLEX PHINEW(20) REAL EP3,FRMAX,XRMAX REAL EP3,FRMAX,XRMAX REAL EP3,FRMAX,XRMAX REAL EP1,EP2 INTEGER NRTS,MXM INTEGER RTC,ITC INTEGER RTC,ITC INTEGER RTC,ITC INTEGER RTC,ITC INTEGER RTC,ITC INTEGER RTC,ITC CGMMGN/CIG/ ICR,IHSP,IGRAPH CGMMGN/CMRAF2/ ERRMAX,NRTSU,SW31,SW71 CGMMGN/CMRAF2/ ERRMAX,NRTSU,SW31,SW71 CGMMGN/CMRAF2/ ERRMAX,NRTSU,SW3,SWR,RTC,ITC,EP3,FRMAX,XRMAX , H,RGGT,FUNC,FRGGT,EIGV, SW7,START,RMX,NFL,KYA1 CGMMGN/CPH2/ PHIZ CGMMGN/CPHCK/ PHINEW,PH2P DATA SW1,SW2,SW3,SW4,SW5,SWR,6*(.FALSE.)/ DATA SW7/.FALSE.; DATA SW7/.FALSE.; MULLER'S (METHOD FGR) RGGTS (GF A LGCALLY) ANALYTIC FUNCTION'. NRTS RGGTS GF F(Z) WILL BE FGUND (HOPEFULLY) AND PLACED IN ARRAY č č NRTS ROOTS OF F(Z) WILL BE FOUND (HOPEFULLY) AND PLACED IN ARRAY NRTS ROOTS OF F(2) WILL BE FOUND (HOPEFULLY) AND PLACED IN ARRAY RT(NRTS). P1, P2, P3 ARE STARTING VALUES FOR THE SEARCH (EXCEPT SEE SW5 BELOW). ROOTS NEAREST THE STARTING VALUES WILL BE FOUND FIRST, ROUGHLY. EP1, EP2 ARE CONVERGENCE TOLERANCE PARAMETERS. A ROOT IS CONSIDERED TO HAVE BEEN FOUND IF. (1) THE MODULUS OF THE RELATIVE DIFFERENCE OF TWO ITERANTS .LT. EP1 (2) MODULII OF THE FUNCTION AND MODIFIED FUNCTION VALUES BOTH .LT. EP1 (2) MODULII OF THE FUNCTION AND MODIFIED FUNCTION VALUES BUTH, E EP2.
 'MODIFIED FUNCTION' MEANS F(Z)/DNR, WHERE DNR=PRODUCT (Z-RT(I)), THE PRODUCT BEING OVER ALL PREVIOUSLY FOUND ROOTS.
 IF MXM ITERATIONS ON ROOT I DOES NOT PRODUCE CONVERGENCE, ITERATION STOPS. WHAT HAPPENS NEXT DEPENDS ON SW4 - DEFAULT PROCEDURE IS TO SET THE REST OF THE ARRAY CONV=.FALSE. AND RETURN. ITERATION SET THE REST OF THE ARRAY CONV=.FALSE. AND RETURN. IF SW1 INFORMATION ABOUT EACH ROOT FOUND IS PRINTED. IF SW2, THEN INFORMATION ABOUT EACH ITERANT OF EACH ROOT IS PRINTED. IF SW3, THEN THE CONJUGATE OF EACH NON-REAL ROOT FOUND IS CONSIDERED TO BE A ROOT ALSO. IF SW4, SOME ATTEMPT IS MADE TO CONTINUE AFTER FAILURE TO CONVERGE ON A ROOT. THE LATEST ITERANT IS ASSIGNED TO RT(I) AND CONV(I)= .FALSE. (ELSE CONV(I)=.TRUE.), AND WE CONTINUE NORMALLY. THE MODIFIED FUNCTION NOW HAS A POLE AT THIS FALSE ROOT. BUT GENERALLY WE SEEM NOW TO CONVERGE TO A ROOT, SO IT CAN BE WORTHWHILE CONTINUING, IF THERE ARE MORE ROOTS ASKED FOR. THIS IS ESPECIALLY TRUE IN CASES OF SLOW CONVERGENCE TO MULTIPLE ROOTS. WHEN NRTS (GOOD OR BAD) ROOTS HAVE BEEN FOUND, MRAF WILL GO BACK AND TRY TO RE-DO THE NONCONVERGENT ROOT ESTIMATES. ON RETURN, CONV WILL INDICATE WHICH ROOTS ARE GOOD. IF SW5, THE STARTING VALUES IN THE SEARCH FOR ROOT I WILL BE PP1 ETC. PLUS WHATEVER WAS IN RT(I) ON ENTRY. IF SW7, THE NEGATIVE CONJUGATE OF EACH ROOT FOUND IS CONSIDERED TO BE A ROOT. IF SW7, ONLY REAL ROOTS ARE FOUND, AND ONLY REAL ARGUMENTS ARE GIVEN TO F, ALTHOUGH F ITSELF IS STILL A COMPLEX FUNCTION (UNLESS YOUR COMPILER RETURNS REAL FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX FUNCTION VALUES IN THE SAME PLACE AS THE REAL PART OF COMPLEX Č C C C C C C C C REAL PART OF COMPLEX FUNCTION VALUES), IN GENERAL IT IS EASIER TO SOLVE FOR ALL ROOTS INSTEAD OF ONLY REAL ROOTS, ESPECIALLY IF THERE ARE ROOTS NEAR REAL AXIS IN REGION OF INTEREST. (IN REVISED VERSION OF PROGRAM, SWR IS TEMPORARILY MADE FALSE IF THERE IS OBVIOUSLY A COMPLEX ROOT NEARBY, IF SWR IS TRUE THEN REAL ROOTS ARE FOUND MORE QUICKLY, SO IT IS BETTER TO SET SWR TRUE INITIALLY. - MJG IF RTC.LT.O ON ENTRY, WE CONTINUE WITHOUT REINITIALIZING JUST AS THOUGH WE WERE ABOUT TO LOOK FOR ROOT ABS(RTC)+1. THE MODIFIED FUNCTION IS NOT WELL DEFINED CLOSE TO ROOTS ALREADY FOUND. THE RELATIVE SIZE OF THE REGION OF INDETERMINACY IS TAKEN TO BE EP3. A STARTING VALUE IN SUCH A REGION WILL BE MOVED OUT BEFORE CONTINUING. AN ITERANT THERE WILL BE ASSUMED TO BE ANOTHER ROOT. CCCCCCC ROOT ROOT. THE NORMAL ITERATION IS ALTERED IN SOME CASES WHERE THE METHOD OTHERWISE TENDS TO GET LOST. H IS THE DIFFERENCE BETWEEN 2 ITERANTS. THE RATIO OF TWO SUCCESSIVE H'S NEVER EXCEEDS XRMAX. IF A NEWLY FOUND H IS TOO LARGE, ITS MAGNITUDE IS DECREASED BUT ITS DIRECTION IS PRESERVED. IF THE RATIO OF THE MODIFIED FUNCTION AT TWO SUCCESSIVE ITERANTS EXCEEDS FRMAX, THE LATEST ITERATION IS DISCARDED AND THE H FOR THAT ITERATION IS NOW TAKEN TO BE THE SAME AS IN THE PRECEEDING ITERATION. IF YOU DON'T ASSIGN ANYTHING TO A CMRAF VARIABLE YOU GET THE DEFAULT SET BY THE DATA STATEMENTS. A COROLLARY IS THAT YOU SHOULDN'T TRY SETTING THEM BY DATA STATEMENTS IN YOUR PROGRAM. THÊ ΙĒ

002454 002455 002456 002458 002458 002460 002461 002463 002463 002464 002464 002464 THE PROCEDURE MAY BE THOUGHT OF AS A GENERATION OF A SEQUENCE Z(1), WHOSE LIMIT IS A ROOT OF F AND WHOSE FIRST 3 ELEMENTS ARE Z(-2)=P1, Z(-1)=P2, Z(0)=P3. A QUADRATIC IS FITTED TO THE FUNCTION AT THREE ELEMENTS Z(1-2), Z(1-1), Z(1), USING NEWTON'S DIVIDED DIFFERENCES. SINCE THE FITTED FUNCTION, A QUADRATIC, IS ANALYTIC WE SHOULD NOT BE SURPRISED WHEN THE METHOD FAILS ON NON-ANALYTIC FUNCTIONS. THE USER CAN OFTEN HELP GREATLY BY MAKING HIS FUNCTION WELL-BEHAVED IN THE REGION OF INTEREST, BY MULTIPLYING HIS ORIGINAL FUNCTION BY ANOTHER WHICH CANCELS SINGULARITIES IN THE ORIGINAL WITHOUT ADDING ANYTHING MORE. CLOSE TO A SINGLE ROOT, CONVERGENCE IS LITTLE FASTER THAN FOR THE SECANT METHOD, BECAUSE THE ONLY ADDITIONAL INFORMATION USED IS FROM X(1-2), WHICH IS A LONG WAY OFF. HOWEVER CONVERGENCE TO MULTIPLE ROOTS IS IMPROVED OVER THE SECANT OR NEWTON-RAPHSON METHODS. IN PARTICULAR CONVERGENCE TO DOUBLE ROOTS IS ABOUT AS FAST AS TO SIMPLE ROOTS. THIS METHOD IS CAPABLE OF FINDING COMPLEX ZERGES OF A FUNCTION WHICH IS REAL FOR REAL ARGUMENT, USING REAL STARTING VALUES. THIS IS NOT TRUE OF, FOR INSTANCE, NEWTON-RAPHSON OR THE SECANT METHOD. CC 002465 002466 002467 002468 002469 002470 002473 002473 002473 002476 002476 0024776 0024776 0024778 0024778 0024778 002478 002483 002483 002483 002483 002483 002483 THIS PROGRAM BEGAN AS A FORTRAN TRANSLATION OF CACM ALGORITHM 196, INCORPORATING CHANGES SUGGESTED BY J. TRAUB AND THE CERTIFICATION APPEARING IN THE CACM, JAN. 1968, P12. THEN SOME REMAINING GLITCHES WERE REMOVED AND NEW FEATURES ADDED BY A.B.LANGDON, EECS, BERKELEY, FEB. 1968. CABS2(Z)=CABS(Z)**2=REAL(Z)**2+AIMAG(Z)**2 (AVGIDS TIME WASTED IN SQRT IN CABS) ANP = NPP IZ = (ANP + 1,)/2, P1 = CMPLX(REAL(PP1),AIMAG(PP1)) P2 = CMPLX(REAL(PP2),AIMAG(PP2)) P3 = CMPLX(REAL(PP3),AIMAG(PP3)) 002486 002487 002488 002489 002490 LNCR=0 LNCR=0 NOPHNU = .FALSE. FAILC = 0 EP3 = 1.E-13 SWR1 = SWR IF (SW7 .AND. SW1) PRINT 512 512 FORMAT (61H NEGATIVE CONJUGATE OF A CONVERGED ROOT IS CONSIDERED A 002491 002492 002493 002493 002494 002495 002496 002497 ROOT, ) IF( RTC.LT.O ) GO TO 100 RTC=0 002498 002499 002500 DÖ 99 I=1,NRTS PHIZ(I) = ,TRUE, 99 CÖNV(I)=,TRUE, 100 RTC=IABS(RTC) 002501 002502 002502 002503 002504 002505 c INITIALIZE SEARCH FOR NEXT ROOT. 10 ITC=0 IF EIGV = 0, PHI(P) IS NOT CALCULATED. IF EIGV = 2, PHI(P) IS CALCULATED FOR ALL ROOTS IF EIGV = 3, PHI(P) IS CALCULATED FOR ALL ROOTS, AND PHI(X) IS PLOTTED FOR UNSTABLE ROOTS. IF EIGV = 4 PHI(X) PLOTTED FOR ALL PAGE 002506 002507 002508 002509 002510 002512 002513 002513 002515 002515 002515 002515 002515 002516 002519 002522 002522 002522 C CCCC PLOTTED FOR UNSTABLE ROOTS, FEIGV = 4, PHI(X) PLOTTED FOR ALL ROOTS, NOCONJE.FALSE, NONEG = .FALSE, SWR = SWR1 SWS, START SEARCH NEAR RT(RTC+1) IF( .NOT.SWS ) GO TO 9 P1=PP1+START P2=PP2+START P3=PP3+START CONTINUE IF C IF CONTINUE FX1=FROOT(P1) ETC. I=RTC+1 9 SĔT IF(SW2) PRINT 498, I FORMAT(5H ROOT,13,23H SEARCH STARTING POINTS) ROOT=P1 002523 002524 002525 002526 002527 498 NRET=1 GO TO 200 FX1=FROOT 002528 1 002529 X1=P1 002530 002531 002532 ) PRINT 499, NRET, ROOT, FROOT 3H Z,I1,1H=,2E22.13,19H MODIFIED FUNCTION=,2E22.13) SW2 499 FORMATCSH ROOT=P2 002533 NRET=2 G0 T0 200 002533 002534 002535 002536 002537 GO 10 200 2 FX2=FR00T X2=P2 IF( SW2 ) PRINT 499, NRET, RCOT, FROOT RCOT=P3 002537 002538 002539 002540 002541 002542 9002543 002543 002545 NRET=3 GØ TØ 200 3 FX3=FRØØT 200 X3=P3 THE FUNCTION WILL BE CONSIDERED SMALL (FOR PURPOSES OF CONVERGENCE TEST) WHEN IT IS SMALL COMPARED TO FI2, EVEN IF THIS IS NOT OF ORDER UNITY FI2 = AMIN1(CABS2(FX3), CABS2(FX2)) FI2 = AMIN1(F12, CABS2(FX1)) IF( SW2 ) PRINT 499, NRET, ROOT, FROOT PREPARE FIRST MULLER-TRAUB ITERATION. H=X3-X2 FX21=(FX2-FX1)/(X2-X1) IF P3= A ROOT WE GET FX3/FX2=0/0 LATER, SO CHECK THAT NOW. GO TO 20 X3=P3 CC 002545 002546 002547 002548 002549 002550 PREP C 002551 С IF 002552

DO MULLER-TRAUB ITERATION. SEE P212 OF TRAUB'S BOOK. 11 CONTINUE C CONTINUE NREDH = FALSE, FX32=(FX3-FX2)/H FX321=(FX32-FX21)/(X3-X1) W=FX32+H*FX321 002557 002560 002561 W IS SO BIG THAT T WILL OVERFLOW, GIVE UP ON THIS ROOT, IF(ABS(REAL(W)).GT.1.E+150,OR.ABS(AIMAG(W)).GT.1.E+150) GO TO 33 T=CSQRT(W*W-(4.,0.)*FX3*FX321) C IF 002563 002564 002565 RHO=W+1 RHO=W+T T =W-T IF( CABS2(T),GT.CABS2(RHO) ) RHO=T H=-2.*FX3/RHO EVEN IF SWR IS ORIGINALLY TRUE, IF WE SMELL AN OBVIOUS COMPLEX ROOT, WE TEMPORARILY MAKE SWR FALSE AND FIND IT. IF(.NOT.SWR) GO TO 12 RH = ABS(REAL(H)) AIH = ABS(AIMAG(H))/RH IF(RH.LE.1.E-O2.AND.AIH.GE.1.E+O2) SWR = .FALSE. IF(.NOT.SWR) NREDH = .TRUE. IF(SWR) H=REAL(H) 2 ITC=ITC+1 X1=X2 č 002573 002574 12 12 ITC=ITC+1 X1=X2 X2=X3 FX1=FX2 FX2=FX3 FX21=FX32 IF NEW H IS TOO MUCH LARGER THAN LAST H REDUCE ITS MAGNITUDE TO THE MAXIMUM ALLOWED KEEPING ITS DIRECTION THE SAME. IF(.NOT.NREDH) GO TO 35 IF(CABS2(H+X3).LT.RMX*RMX) GO TO 13 YR = CABS2(H+X3).LT.RMX*RMX) GO TO 13 YR = CABS2(H+X2).LT.RMX*RMX) 002576 002577 002578 002579 002582 002583 XR = CABS2(H)/CABS2(X2-X1)002586 002587 002588 GO TO 36 35 XR = CABS2(H)/CABS2(X2-X1) IF(XR.LT.XRMAX*XRMAX) GO TO 13 36 IF( SW2 ) PRINT 496, H 496 FORMAT(30H (X3-X2)/(X2-X1) TOO LARGE. H=,2E22.13,17H WILL BE REDUC 002591 .ED.) H=H*(XRMAX/SQRT(XR)) H=H*(XRMAX/SQRT(XR))
13 X3=X2+H
IF(CABS2(X3 - P3).LT.ERRMAX*ERRMAX) G0 T0 40
IF(SW2) PRINT 41
41 F0RMAT(19H X3 T00 FAR FR0M P3)
I = RTC + 1
G0 T0 33
CHECK F0R CONVERGENCE.
, DUE T0 CL0SE RELATIVE SPACING 0F ITERATES.
40 IF(REAL(X2).EQ.0.AND.AIMAG(X2).EQ.0.) G0 T0 45
IF(CABS2(X3/X2 - 1.).LT.EP2*EP2) G0 T0 22
45 IF(CABS2(X3/X1 - 1.).LT.EP2*EP2) G0 T0 22
IF(REAL(X1).EQ.0.AND.AIMAG(X1).EQ.0.) G0 T0 46
IF(CABS2(X3/X1 - 1.).LT.EP2*EP2) G0 T0 22
46 IF(CABS2(X3 - X1).LT.EP2*EP2) G0 T0 22
R00T=X3 002594 002595 002599 ROOT=X3 ROOT=X3 NRET=4 GG TG 200 4 FX3=FROOT IF NEW ITERATE GIVES MUCH LARGER FUNCTION MODULUS, HALVE THE STEP AND TRY AGAIN. IF(ABS(REAL(FX3)).GT.1.E+150.OR.ABS(AIMAG(FX3)).GT.1.E+150)GG TO C IF ( CABS (FX3)/CABS (FX2).LT.FRMAX ) GO TO 103 IF ( CABS (FX3)/CABS (FX2).LT.FRMAX ) GO TO 103 44 IF ( SW2 ) PRINT 497, X3,FUNC,FROOT 497 FORMAT (36H FZ3/FZ2 TOO LARGE. 23 ADJUSTED FROM,2E22.13/ 11H FUNCTION=,2E22.13,20H, MODIFIED FUNCTION=,2E22.13) 002615 002616 002617 497 FORMAT(36H F237F . 11H FUNCTION=, H=.5*H GO TO 13 ITERATION COMPLETE. 103 CONTINUE ITERATION OUTPUT I=RTC+1 IEC SU2 ) PRINT 002621 002622 С С 11ERATION COTFOI 1 =RTC+1 1F(SW2) PRINT 500, 1, ITC,X3,FUNC,FROOT 500 FORMAT(6H ROOT,13,12H, ITERATION,13,16H, ROOT ESTIMATE=,2E22.13/ 11H FUNCTION=,2E22.13,20H, MODIFIED FUNCTION=,2E22.13) CHECK FOR CONVERGENCE .DUE TO FUNCTION BEING VERY SMALL. 20 IF( CABS2(FX3)+CABS2(FUNC),LE.EP2*EP2) GO TO 22 NOT CONVERGED YET. SHOULD WE ITERATE AGAIN. IF(CABS2(X3),GT,RMX*RMX.AND.CABS2(H).LT.25.AND.ITC.GE.5)GO TO 33 IF ROOT-SEARCH SEEMS TO BE GOING OFF TO INFINITY, STOP IT. IF(CABS2(X3),GT,RMX*RMX.AND.ITC.GE.5.AND.CABS2(X2-X1).GT..25.AND. .CABS2(H).GT.CABS2(X2-X1)) GO TO 33 IF( ITC.LT.MXM ) GO TO 11 IF WE SEEM TO BE GETTING CLOSE TO A ROOT, ALLOW 10 MORE ITERATIONS BEFORE GIVING UP, IF(ITC.LT.MXM+10.AND.CABS2(FUNC).LT.1.E-06*FI2) GO TO 11 IF(ITC.LT.MXM+20.AND.CABS2(FUNC).LT.1.E-12*FI2) GO TO 11 IF FUNCTION IS PRETTY SMALL, BUT MORE THAN EP2,, SEE IF EIGENMODE IS GOOD, AND IF IT IS, CONSIDER ROOT COMVERGED. ROOT ESTIMATE=, 2E22.13/ Č 002630 С 002632 002633 C 002636 c C 

IF(EIGV.EQ.0) GO TO 33 IF (CABS2(FUNC).GT.1.E-14) GO TO 33 CALL SAVPHI IF(CABS2(PHI(NRTSV,1) - (1.,0.)).LT..0001) GO TO 22 NO. WE HAVE TRIED LONG ENOUGH. 33 IF( SW1 ) PRINT 502, 1, ITC 502 FORMAT(28H FAILURE TO CONVERGE ON ROOT,13,110,11H ITERATIONS ) CONV(I)=.FALSE. FAILC = FAILC + 1 UNLESS SW4, FAILURE TO CONVERGE MEANS WE MUST STOP AFTER SETTING THE REST OF THE ARRAY CONV =.FALSE.. IF( SW4 ) GO TO 22 RTC=RTC+1 43 RT(RTC)=X3 DO 21 I=RTC,NRTS 21 CONV(I)=.FALSE. RETURN 002642 002643 002644 002645 002646 002647 002648 C 002649 002650 002651 002652 002653 č 002654 002655 002656 002657 002658 002659 RECORD THE FINDING OF A ROOT. 22 CONTINUE C 002659 002660 002661 002662 002663 002663 002664 002665 22 CGNTINUE INSTITUTE NUMBER IF(NOT.SW7.GR.,NGT.SW3.GR.NGT.CGNV(1)) GG TG 30 L = 0 IF (NOT.SW7.GR.ADSTREAM STREAM IF(.NOT.SW7 .OR. ,NOT.SW3 ,OR..NOT. CONV(1)) GO TO 30 L = 00026669 002667 002669 002670 002672 002673 002673 002673 002674 002675 002675 002675 002675 002679 002680 002680 002682 002684 002685 002686 002687 002688 C 002689 002690 002691 002692 C 002694 002695 002697 002698 002698 002700 002701 002702 002703 002703 002705 002705 002706 002707 002708 002708 26 CONTINUE PHIZ = .TRUE. IFF PHI(0) = 0, THE PURPOSE OF PHIZ IS EXPLAINED IN *ROOTS* FOLLOWING STATEMENT 76 IF(CABS2(PHJ)/SUM.GT.1.E-11) PHIZ(RTC) = .FALSE. PHPLUS = CABS2(PHI(NRTSV,IZ+1)) PHMIN = CABS2(PHI(NRTSV,IZ-1)) IF(CABS2(PHJ).GT.1.E-04*PHPLUS) PHIZ(RTC) = .FALSE. IF(CABS2(PHJ).GT.1.E-04*PHPLUS) PHIZ(RTC) = .FALSE. ABNM = SOBT(FUM) C C IF(CABS2(PHJ).GT.1.E-04*PHPLUS) PHIZ(RTC) = .FALSE. IF(CABS2(PHJ).GT.1.E-04*PHMIN) PHIZ(RTC) = .FALSE. ABNM = SQRT(SUM) I = IZ ARNM = ARPHI(I) IF (ABPHI(I) .LE. .1E-03) ARNM = .5*ARPHI(NPP) DO 31 I = 1, NPP ABPHI(I) = ABPHI(I)/ABNM ARPHI(I) = ABPHI(I)/ABNM ARPHI(I) = ARPHI(I) - ARNM IF (ARPHI(I) .LE.-180.) ARPHI(I) = ARPHI(I) + 360. IF (ARPHI(I) .GT. 180.) ARPHI(I) = ARPHI(I) - 360. IF (ARPHI(I) .GT. 180.) ARPHI(I) = ARPHI(I) - 360. IF (NTS.GT.2) GO TO 31 DON'T CALCULATE PHINEW UNLESS THIS IS THE FIRST GOOD ROOT WE HAVE FOUND. IF(NOPHNU) GO TO 31 IF(PHZP.NE.SGNL(1, PHIZ(RTC))) GO TO 31 IF(PHZP.NE.SGNL(1, PHIZ(RTC))) GO TO 31 IF(RALAG(RT(RTC)).LT.-1.E-06.AND.SW31) GO TO 31 IF( REAL(RT(RTC)).LT.-1.E-06.AND.SW31) GO TO 31 IF( REAL(RT(RTC) 002709 002710 002712 002713 002713 002714 002715 002716 002716 002717 002718 002718 002720 002721 002722 C 002722 002723 002724 002725 002726 002727 C 002727 002728 002729 002730 002731 002732 002733 31 CONTINUE IF(.NOT.SW1) GO TO 5304 PRINT 507, (ABPHI(I), I 507 FORMAT (10E12.4/10E12.4) PRINT 506, (ARPHI(I), I 506 FORMAT (10F12.2/10F12.2) 5304 IF(.NOT.SW5) GO TO 47 = 1, NPP) = 1, NPP) 002734

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START = ABS(REAL(X3)) IF(START.GE.RMX) START = 1.5 47 IF(ITC.GT.O) GO TO 5305 IF ONE OF THE STARTING POINTS WAS A ROOT (I.E. ITC = O), THEN CHANGE THE STARTING POINTS, OTHERWISE WE WILL GET THE SAME ROOT AGAIN NEXT CHANGE = .15*(P2 - P1) P1 = P1 + CHANGE P2 = P2 + CHANGE P3 = P3 + CHANGE S05 CONTINUE QUIT IF ENOUGH ROOTS ARE FOUND, OR IF WE ARE ALMOST OUT OF TIME. IF(WARN(O)) RETURN IF(RTC.GE.NRTS) GO TO 300 IF(FAILC.GT.NFL) RETURN IF SW7, THE NEGATIVE CONJUGATE OF A CONVERGENT ROOT FOUND IS TO BE REGARDED AS A ROOT. 28 IF (.NOT, SW7 .OR, NONEG .OR. ABS(REAL(X3)).LT.EP1) GO TO 29 IF(SW1) PRINT 514 514 FORMAT(62H NEGATIVE CONJUGATE OF LAST ROOT WILL BE CONSIDERED NEXT C 5305 C C IF(SW1) PRINT 514 514 FORMAT(62H NEGATIVE CONJUGATE OF LAST ROOT WILL BE CONSIDERED NEX ROOT. ) ITC = 0 NONEG = .TRUE. X3 = -1.*CONJG(X3) ROOT = X3 IF(.NOT,CONV(RTC)) CONV(RTC+1) = .FALSE. PHIZ(RTC+1) = PHIZ(RTC) GG TO 30 29 NONEG = .FALSE. IF SW3, THE CONJUGATE OF A CONVERGENT NON-REAL ROOT IS TO BE REGARDED AS A ROOT ALSO, IF NOT ALREADY USED. 23 IF( .NOT.SW3. OR. ABS(AIMAG(X3)).LT.EP1 .OR. NOCONJ ) GO TO 10 IF( SW1 ) PRINT 503 503 FORMAT(53H CONJUGATE OF LAST ROOT WILL BE CONSIDERED NEXT ROOT.) ITC=0 NOCONJ=.TRUE. X3=CONJG(X3) ROOT=X3 IF(.NOT.CONV(RTC)) CONV(RTC+1) = .FALSE. PHIZ(RTC+1) = PHIZ(RTC) NRET=5 GO TO 30 EVALUATES ERGOT 514 FORMAT (62H NEGATIVE CONJUGATE OF LAST ROOT WILL BE CONSIDERED NEXT С 5 G0 10 30 EVALUATES FR00T WHERE (R00T-RT(1))*...*(R00T-RT(RTC))*FR00T=F(R00T) RETURNS T0 STATEMENT NUMBER NRET. SPECIAL TREATMENT WHEN R00T IS CLOSE T0 AN EARLIER FOUND R00T IS DESCRIBED AT HEAD OF PROGRAM. 200 CONTINUE IF(WARN(0)) RETURN FUNC=F(R00T) IF(SWR) FUNC=REAL(FUNC) DNR=(1.,0.) IF(RTC.EQ.0) G0 T0 202 TEST=EP3*EP3*CABS2(R00T) D0 203 I=1,RTC IF( CABS2(R00T-RT(I)),GE.TEST ) G0 T0 201 T=.33333333*(P1+P2+P3)+EP1 IF( NRET.NE.1 ) G0 T0 210 P1=2.*P1-T R00T=P1 G0 T0 200 210 IF( NRET.NE.2 ) G0 T0 211 P2=2.*P2-T R00T=P2 G0 T0 200 CCCC Ĉ P2=2.*P2-T R00T=P2 g0 T0 200 211 IF( NRET.NE.3 ) G0 T0 212 P3=2.*P3+T R00T=P3 g0 T0 200 201 DNR=DNR*(R00T-RT(I)) 203 C0NTINUE 202 FR00T=FUNC/DNR g0 T0 (1,2,3,4,5), NRET 212 IF( SW1 ) PRINT 504 504 F0RMAT(71H ITERANT CLOSE TO R00T F0UND BEF0RE. ACCEPT IT AS A R00T , WITHOUT TEST.) FR00T=0, g0 T0 22 ŏŏ28ŏŏ 002801 002803 002804 002806 002807 002808 002809 002810 002811 GO TO 22 002813 002815 002816 002817 002818 TERMINATION PROCEDURE, INCLUDES CAPABILITY OF REDOING 'BAD ROOTS' -ONES WHICH HADN'T CCCC CONVERGED. HOW MANY GOOD ROOTS ARE THERE .. HOW MANY GOOD ROOTS ARE INERE., 300 NCR=0 DG 301 I=1,NRTS IF ( CGNV(I) ) NCR=NCR+1 301 CONTINUE IF ALL ROOTS ARE GOOD, OR IF WE HAVEN'T BEEN ASKED TO REDO BAD ONES, IF ALL ROOTS ARE GOOD, OR IF WE HAVEN'T BEEN ASKED TO REDO BAD ONES, WE ALL ROOTS ARE GOOD, OR IF WE HAVEN'T BEEN ASKED TO REDO BAD ONES, WE ALL ROOTS ARE GOOD, OR IF WE HAVEN'T BEEN ASKED TO REDO BAD ONES, WE QUIT, WE QUIT, WE QUIT, WE QUIT, WE AND FO NOTS OF NOT SWA OR NOR IF INCR ) RETURN 002819 č 002820 002821 002822 002823 002824 002825 Č 002826 IF ( NCR.EQ.NRTS . OR. . NOT.SW4 . OR. NCR.LE.LNCR ) RETURN

FIND FIRST BAD ROOT. DO 302 I=1,NRTS IF( .NOT.CONV(I) ) GO TO 303 302 CONTINUE IF IS THE LAST ROOT WE WOULD ONLY BE REPEATING THE SAME NON-CONVERGENT ITERATIONS IF WE TRIED AGAIN. 303 IF( I.EQ.NRTS ) RETURN MOVE LAST ROOT INTO 1ST BAD ROOTS PLACE. GO BACK AND TRY TO FIND A NEW LAST ROOT. GOOD LUCK. IF( SW1 ) PRINT 505, I 505 FORMAT(48H WILL ATTEMPT TO RE-DO UNCONVERGED ROOT ESTIMATE,I3) X3=RT(NRTS) RT(NRTS) =RT(I) RT(NRTS)=RT(I) С C č ç 002838 002839 002840 002841 002842 002843 002843 002844 RT(I)=X3 CONV(I)=CONV(NRTS) RTC=NRTS-1 LNCR=NCR CONV(NRTS)=.TRUE. ĠØ TO 28 Ĉ END SUBROUTINE EIGPLT PLGTS EIGENFUNCTIONS COMPLEX CMEGA, PHI(18,22), A(20,20), F(100), IPXK0 INTEGER PMAX REAL KYAI LGGICAL GOOD(18) DIMENSION X(100), Y(100) DATA KYAI/0./ COMMON/PHIC/ A,NRTSV,PHI COMMON/PHIC/ A,NRTSV,PHI COMMON/PHIC/ A,NRTSV,PHI COMMON/TWX/NP COMMON/TVNE/LPENÓN,LPENOFF,ITALICS,IWINK,INTENSE,IRIGHT,IUP NDY1 = NDY - 1 CALL PAGE PRINT 10 NRTSV =,14) PRINT 10 NRTSV =,14) PRINT 10 NRTSV =,14) PRINT 11, GOOD 11 FORMAT(2015) IF(NRTSV.E0.1) RETURN NRTSU = NRTSV - 1 PMAX = (NP-1)/2 PI = 3.141592653589 DO 3 N = 1, NRTSU IF(NGT.GOD(N)) GG TO 3 GMEGA = PHI(N ,NP+1) KYAI = REAL(PHI(N,NP+2)) SUM = 0. DO 6 J = 1,NP 6 SUM = SUM + CABS2(PHI(N,J)) SUM = SGRT(SUM) DO 7 J = 1, NP 7 PHI(N,J) = PHI(N,J)/SUM DO 6 I = = 1, 100 AI = 1, 100, YUO = 2 PHIX(I) C 002854 002856 002857 002860 002873 002877 002878 002879 PHI(N,J) = PHI(N,J)/SUM DG 1 I = 1, 100 AI = I X(I) = AI/100, XKO = 2.*PI*X(I) F(I) = (0.,0.) DG 2 J = 1, NP P = J - PMAX - 1 IPXKO = CMPLX(0.,XKO*P) F(I) = F(I) + PHI(N,J)*CEXP(IPXKO) Y(I) = F(I)*CGNJG(F(I)) YM = Y(1) 002888 002891 002892 002893 002894 IPARC 2 F(I) = F(I) *CONJG(F(1,, YM = Y(1) D0 5 I = 2,100 5 YM = AMAX1(YM,Y(I)) XMAX = 1. YMIN = 0. XMIN = 0. CALL MAPS(XMIN,XMAX,YMIN,YM,0.11328,1.0,0.15,1.0) CALL SETCH(1:,2.,1,0,1,0,0) WRITE(IGRAPH,4) OMEGA 4 FORMAT(34H PLGT OF PHI**2 VS. X FOR OMEGA = ,2E22.13) WRITE(IGRAPH,8) KYAI, IBRANCH(N) 8 FORMAT(8H KYAI = ,E12.6,10X,7H BRANCH,I4) CALL FRAME(1) D0 PLGT OF PHI(X,Y) FM = SQRT(YM) YM = NDY CALL MAPS(XMIN,XMAX,YMIN,YM,0.11328,1.0,0.15,1.0) D0 12 K = 1, NDY1 FLK = K 002898 002898 002901 002902 002906 002907 C 002912 ŏŏ2913 

002915 002916 002917 002919 002920 002920 002922 002923 002923 002924 002925 002925 002926 002927 002928	13 12 14 9 3	PHASE = 2.*3.1415926535*FLK/YM DØ 13 I = 1,100 Y(I) = FLK + REAL(F(I)*CEXP((0.,1.)*PHASE))/FM CALL TRACE(X,Y,100) CALL SETCH(1.,1.,1,0,1,0,0) WRITE(IGRAPH,14) FØRMAT(51H PLØT ØF PHI(X,Y) FØR ØNE WAVELENGTH IN Y DIRECTIØN) CALL FRAME(1) PRINT 9, ØMEGA, KYAI, IBRANCH(N) FØRMAT(35H EIGENFUNCTIØN PLØTTED FØR ØMEGA = ,2E10.2,8H KYAI = ,E1 .0.3,10X,7H BRANCH,I4) CØNTINUE RETURN END
002929 002930 002931 002932 0029334 002935 002935 002936 002937 002938 002939 002941 0022941 0022944 0022944 0022945 0022946 0022946 0022946 0022946 0022951 0022953 0022953 0022954	901 902 903 904 905 9800	SUBROUTINE PRNPLT(X,Y,XMAX,XINCR,YMAX,YINCR,ISX,ISY,NPTS) CCMMON/CIO/ ICR,IHSP,IGRAPH DIMENSION X(NPTS),Y(NPTS),IGRID(105),XAXIS(11) INTEGER BLANK,DOT,STAR,PLUS / 1H,1H.,1H*,1H+ / FORMAT(14X,105Å1) FORMAT(14X,105Å1) FORMAT(15X,103(1H.)) FORMAT(15X,103(1H.)) FORMAT(16X,11(1H+9X)) FORMAT(16X,11(1H+9X)) FORMAT(46HISCALING ERROR IN PRNPLT, EXECUTION TERMINATED ) IF(XINCR,EG.0., 0R.YINCR.EQ.0.) GO TO 800 YAXMIN=0.01*YINCR IZERO=YMAX/YINCR+1.5 JZERO=103.5-XMAX/XINCR IF(JZERO,GT.103.0R.JZERO.LT.4) JZERO=2 PRINT 905 PRINT 903 DO 10 I=1,51 IF (I.NE.IZERO) GO TO 16 DO 14 J=1,105 IGRID(J)=FLUS GO TO 15 DO 11 J=1,105
002955 002957 002957 002958 002959 002960 002961 002962 002963 002964 002965	11 15	IGRID(J)=BLANK IGRID(JZERØ)=PLUS IGRID(104)=DØT IGRID(2)=DØT DØ 12 K=1,NPTS ITEST = (YMAX-Y(K))/YINCR+1.5 IF(ITEST ,NE.I) GØ TØ 12 J=103.5-(XMAX-X(K))/XINCR IF(J,GT.103) J=105 IF(J,LT.3) J=1 IGRID(J)=STAR CØNTINUE
002966 002967 002968 002969	12	IF(MOD(I,10).EQ.1) GO TO 13 PRINT 901,IGRID GO TO 10
002970 002971 002972	13	YAXIS=YMAX-(I-1)*YINCR IF(ABS(YAXIS).LT,YAXMIN) YAXIS=0. PRINT 902,YAXIS,(IGRID(J),J=1,105)
002973 002974 002975 002976 002977 002978 002979 002980 002980 002981	10 20 800	CONTINUE PRINT 903 PRINT 905 DØ 20 M=1,11 XAXIS(M)=XMAX-XINCR*(FLØAT(11-M))*10.0 IF(ABS(XAXIS(M)).LT.XAXMIN)XAXIS(M)=0. CØNTINUE PRINT 904,XAXIS,NPTS RETURN PRINT 9800
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002985 002986 002987 002988 002989 002990 002991 002991 002992	SUBROUTINE TRACET(TEXTURE,X,Y,NPT) REAL TEXTURE,X(1),Y(1) INTEGER NPT IF(TEXTURE.EQ.0.0) CALL TRACE(X,Y,NPT) IF(TEXTURE.NE.0.0) CALL TRACEP(X,Y,NPT,3) RETURN END
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FUNCTION SGNL(A,L) LOGICAL L IF(L) B = -1. IF(.NOT.L) B = 1. SGNL = SIGN(A,B) RETURN END

LCGICAL WARN FUNCTION WARN(N) WARN = .FALSE. RETURN END

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An outline of the program LOCAL and its subroutines is shown in the flow chart. LOCAL was used to find the normal modes and frequencies using the local method. In contrast to ROOTS, which was used for the nonlocal method, all branches  $\omega(k_y)$  were not calculated, and "smart" root-finders and branch-followers were not used in LOCAL. Instead, a Newton-Raphson iteration scheme was used to find the solutions  $x = x_L$ and  $\omega = \omega_L$  to the simultaneous equations  $D(x,\omega,k_x = 0) = 0$  and  $\partial D/\partial x = 0$ . Only the mode with the highest growth rate was sought. The Newton-Raphson iteration method worked well only if initial guesses of  $x_L$  and  $\omega_L$  were reasonably close to the correct solutions. The mode found as the fastest-growing mode was really only the local maximum of Im  $\omega(k_y)$  closest to the initial value of  $k_y$  chosen. The limitations in LOCAL were tolerable because ROOTS could be used to make good estimates of  $x_L$ ,  $\omega_L$  and  $k_y$  for initial guesses in LOCAL.

The routine QDERIV shown in the flow chart was similar to the subroutine DISP used with ROOTS, but instead of calculating det A p,p' it calculated

 $D(x,\omega) = G_0(\omega) + G_{-1}(\omega)\exp(-ik_0x) + G_{+1}(\omega)\exp(ik_0x)$ 

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(with  $k_x = 0$  assumed). Since the G's do not depend on x, the derivatives  $\partial D/\partial x$  and  $\partial^2 D/\partial x^2$ , needed for the Newton-Raphson method, could be easily found analytically, and were also returned by QDERIV. On the other hand, the derivatives  $\partial D/\partial \omega$  and  $\partial^2 D/\partial x \partial \omega$  were found by calling QDERIV twice using slightly different values of  $\omega$ .

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### FLOW CHART OF LOCAL



# FLOW CHART OF SEARCH (KY, XI, WI, XL, WL)

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KY is  $k_{ya_{i}}$ . XI and WI are the initial guesses for  $x_{L}$  AND  $\omega_{L}$  (normalized to  $\omega_{ci}$  and  $2\pi/k_{0}$ ). XL and WL are final values of  $x_{L}$  and  $\omega_{L}$  returned by SEARCH



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## Listing of LOCAL and Subroutines

Like ROOTS, LOCAL runs on the A-machine of the National Magnetic Fusion Energy Computer Center. However, it is not believed to contain any nonstandard Fortran, except for the PROGRAM, READ, FORMAT, and FUNCTION statements.

The subroutines QFKCAL, QCALC, BESSI, BESSJ, CABS2, ZEE, ABCALC, MAXW, RING, and COLD, used by LOCAL, have been already listed with ROOTS, so they are not shown here.

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05 18.55.58 09/16A 1976 PROGRAM LOCAL (INPUT, HSP, OUTPUT, TAPE10=OUTPUT) REAL KY, KOAI, KOSQ, KYI, KYNEW, KYI, KY2 COMPLEX XI, WI, XL1, WL1, XL2, WL2, WIDTH , XLNEW, WLNEW COMPLEX SLPX1, SLPX2, SLPW1, SLPW2, D2XDK2, D2WDK2 COMPLEX SLDX, D2GDX2, DQDW, XL, WL, QO, DELW, DQDKXS INTEGER QMAX LOGICAL FAIL LCGICAL UNMAG COMMON/CMAG/UNMAG COMMON/CMAG/UNMAG COMMON/CDFU/ DFU(4), YYY(4) COMMON/CDFE/ RLARM(4), AMASS(4), CHARG(4), JAY(4), DENSE(4), NSPEC COMMON/CRESJ/ QMAX(4), NMASS(4) COMMON/CGDER/ QO, DQDX, D2QDX2, DQDW, DQDKXS COMMON/CRESJ/ QMAX(4), NMAX(4) DATA AEAI/O./, DFU/4*0./, JAY/4*0/ DATA RLARM/4*1./, AMASS/4*1./, CHARG/4*1./, DENSE/4*1./, AMASQ ./4*1./ BOX BO5 DATA RLARM/4*1./,AMASS/4*1./,CHARG/4*1./,DENSE/ ./4*1./ DATA ICR/2/ PRINT 17 17 FORMAT(1H1) GO TO 18 15 CONTINUE 18 READ (ICR,16) AIL, R, CMPSQ,XI,WI,KYI,UNMAG 16 FORMAT(8F5.0,11) IF(AIL.EQ.0.) CALL EXIT IF(XI.NE.0..OR.WI.NE.O..OR.KYI.NE.0.) GO TO 20 IF(FAIL) GO TO 15 WI = WL WI = WL XI = XL KYI = XL IONS = 5H MAG. IF(UNMAG) IONS = 6H UNMAG IF(UNMAG) IONS - ON C. GAM = 1. IF(R,LE,1,) GAM = 0. RLARM(2) = 1./SQRT(R) RLARM(3) = 0. AMASS(3) = 1./3700. CHARG(3) = -1. DENSE(1) = R/(R - GAM DENSE(2) = -GAM/(R-GAM NSPEC = 3 AMASS(3) = AMASS(3)** = 0./3700. = 1./3700. = R/(R - GAM) = -GAM/(R-GAM) CHARG(3) = -1, DENSE(2) = -GAM/(R-GAM) DENSE(2) = -GAM/(R-GAM) NAMEC 3 = AMASS(3)*2 19 (DG1 2 = AMASS(3)*2 19 (DG1 2 = 3) = -GAM/(R-GAM) CONTRACT 19 (CONTRACT 19 (CONTR

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IF(ITER.GT.99) GG TG 13
SLPX1 = (XL1-XL)/DK1
SLPX2 = (XL2-XL1)/DK2
D2XDK2 = 2.*(SLPX2 - SLPX1)/(DK1+DK2)
SLEW1 = (WL1 - WL1/DK1
SLPW2 = (WL2 - WL1)/DK1
SLPW2 = (WL2 - WL1)/DK1
SLPW2 = (XL2 - KL1)/DK1
SLPW2 = (XL2 - KL1)/DK2
D2WDK2 = 2.*(SLPW2-SLPW1)/(DK1+DK2)
XLNEW = .5*(XL+XL1) + SLPX1*(KYNEW - .5*(KY+KY1)) + .5*D2XDK2*
(KYNEW - .5*(KY+KY1))**2
IF(CABS(XLNEW-XL))GT.1)/GG TG 6
WLNEW = .5*(KY+KY1))**2
IF(CABS(XLNEW-XL))GT.2)/GG TG 6
KY1 = KLNEW
WL = .5*(KY+KY1))**2
IF(SLOPE2.GT.0.)/KYNEW = .9*KY
IF(SLOPE2.GT.0.)/KYNEW = .1.*KY
GG TG 7
WL1 = WL + .5*DELW
WL2 = WL1 + DELW
XCGRR = CABS(WIDTH)
XCGRR = CABS(WIDTH)
XCGRR = CABS(WIDTH)
XCGRR = CABS(WIDTH)
XCGRR = CABS(WIDTH)
XCGRR = CABS(WIDTH)
ACCGRR = CABS(WIDTH)
CCGRR = CABS(WIDTH)
XCGRR 00107 000107 000108 000109 000117 000118 000119 000120 000121 ŏŏŏiīzż 000123 000124 000125 000126 000127 000128 000129 000132 000133 000134 000135 000135 000136 ËŇD 000140 000141 000142 000143 000144 000145 000146 SUBROUTINE SEARCH (KY,XI,WI,XL,WL) LOGICAL FAIL,NDPOLE COMPLEX XI,WI,XL,WL,X,W,QO,QI,Q2,DQDX,DQDW,D2QDX2,D2QDXW,DQDX1,DET COMPLEX XDIFF,WDIFF REAL KY,KXSQ COMMON/CFAIL/ FAIL COMMON/CQDER/ QO,DQDX,D2QDX2,DQDW,DQDKXS COMMON/CNDP/ NDPOLE COMMON/CNDP/ NDPOLE COMMON/CKX/ KXSQ NDPOLE = .FALSE. KXSQ = 0. FAIL = .FALSE. ITER = 0 X = XI W = WI CALL QDERIV(X,W,KY,QO,DQDX,D2QDX2) 000149 000150 000151 000152 000153 000154 11ER = 0 X = XI W = WI 2 CALL GDERIV(X,W,KY,QO,DQDX,D2QDX2) CALL GDERIV(X,W+.OO1,KY,Q1,DQDX1,Q2) DQDW = 1000.*(G1-QO) D2GDXW = 1000.*(CDGDX1 - DQDX) DET = DQDX*D2QDXW - D2QDX2*DQDW XDIFF = (-D2QDXW*QO + DQDW*DQDX)/DET XL = X + XDIFF WDIFF = (D2QDX2*QO - DQDX*DQDX)/DET WL = X + WDIFF IF(CABS(XL - X).GT..005) GG TG 1 IF(CABS(WL - W).GT..005) GG TG 1 NDPGLE = .TRUE. CALL GDERIV(X,W,KY,QO,DQDX,D2QDX2) CALL GDERIV(X,W,KY,QC,DQDX,D2QDX2) CALL GDERIV(X,W,KY,Q2,DQDX1,DQDX1) KXSQ = 0.000 CALL GDERIV(X,W,KY,Q2,DQDX1,DQDX1) DQDKXS = 10000.*(G1-Q0) DQDKXS = 10000.*(G2-Q0) KXSQ = 0. NDPGLE = .FALSE. RETURN 1 ITER = ITER + 1 IF(ITER.GT.99) GG TG 4 IF(CABS(WL/W - 1.).GT..2) GG TG 3 X = XL W = WL GG TG 2 3 XL = .5*(XL + X) WL = .5*(WL + W) GG TG 1 4 FAIL = .TRUE. PRINT 5. KY XL WL GO,DQDX END END 000157 000158 000160 000160 000160 000163 000163 000164 000165 000166 000165 000166 000167 000168 000169 000170 000171 Ð 000173 000174 000175 000176 000177 000178 000180 000181 000182 000183 000187 000188 000189 000190 000191 RETURN END

SUBROUTINE GDERIV(X,W,KY,G,DGDX,D2GDX2) COMPLEX X,W,OMEGA,GO,GPLUS,GMINUS, APP,BP,AP,BP,IX,DEPOLE,OMEGD,U COMPLEX G,DGDX,D2GDX2,EIKOX INTEGER PMAX,P REAL KY,KYSG,KYAI,KOAI,KOSG,KXSQ LOGICAL UNMAG,NDPOLE COMMON/CMAG/UNMAG COMMON/DIS/GF,KYSG, KOSG,EPSF(4),XXX,KYAI,KOAI,XXXX,PMAX, OMPSG,AEAI,KMAX(4),AMASG(4),XXXXX(4),KMIN(4) COMMON/CDFU/ DFU(4),YYY(4) COMMON/CSPEC/ RLARM(4),AMASS(4),CHARG(4),JAY(4),DENSE(4),NSPEC COMMON/CSPEC/ RLARM(4),AMASS(4),CHARG(4),JAY(4),DENSE(4),NSPEC COMMON/CSPEC/ RLARM(4),AMASS(4),CHARG(4),JAY(4),BPP(4) COMMON/CNDP/ NDPOLE COMMON/CNDP/ NDPOLE COMMON/CNDP/ NDPOLE COMMON/CKX/KXSG DATA PI /3.1415926538/ KYAI = KY KYSG = KY*KY OMEGA = W 000193 000194 000195 000196 000197 000198 000199 000200 OMEGA = W 000213 000214 000217 000218 000219 000223 000223 000224 000225 C 000226 000227 000228 000229 000229 000231 000234 000235 000237 000238 000239 000240 000241 С 000245 000245 M = 1 CALL ABCALC GO = KYSQ + KXSQ DO 3 I = 1, NSPEC 3 GO = GO + AP(I)*OMPSQ*CHARG(I)**2*DENSE(I)/AMASS(I) DU 3 I = 1, NSPEC 3 G0 = G0 + AP(I)*CMPSQ*CHARG(I)**2*DENSE(I)/AMASS( M = 2 CALL ABCALC GPLUS = 0. D0 4 I = 1, NSPEC GPLUS = GPLUS + EPSF(I)*(APP(I) + BPP(I)) 4 GMINUS = GMINUS + EPSF(I)*(AP(I) - BP(I)) IX = (0, 2.)*3.1415926535*X EIKOX = CEXP(IX) Q = G0 + EIKOX*GPLUS + GMINUS/EIKOX DQDX = 2.*PI*(EIKOX*GPLUS - GMINUS/EIKOX)*(0,,1.) D2QDX2 = -4.*PI*PI*(Q - GO) Q = Q*DEPGLE DQDX = DQDX*DEPGLE D2QDX2 = D2QDX2*DEPGLE RETURN END 000252 000253 000253 000264 000265 000265

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The characteristics of the most unstable mode are shown for various values of density  $\omega_{pi}^2/\omega_{ci}^2$  and density gradient  $k_0a_i$ , using the local method (shown in Table I) and the nonlocal method (shown in Table II). In all cases the mirror ratio R = 3 and the mass ratio  $m_1/m_2 = 3700$ . In the table headings,  $x_{peak}$  is the value of x at which  $|\phi(x)|$  is greatest;  $\Delta x$  is the half-width of  $|\phi(x)|^2$ ; k is the value of k x at which  $|\tilde{\phi}(k_x)|^2$  is greatest; and "p_{max} needed" is the minimum value of Pmax which is needed in order for the nonlocal method to be valid [defined by requiring  $|\tilde{\phi}(\pm p_{\max} k_0)| < |\tilde{\phi}(k_x \text{ peak})|/4$ ]. The results in the two tables are seen to be in good agreement when inequality (21) is wellsatisfied. The one case in which the agreement is not very good is for  $\omega_{pi}^2/\omega_{ci}^2 = 1000$  and  $k_0 a_i^2/2\pi = 0.40$ . This set of parameters is near the border between regions D and E in Figure 6. Using the local method, the "region E instability" (at higher Re  $\omega$  and  $k_{_{\bf V}})$  has a slightly higher growth rate than the "region D instability," while the reverse is true using the nonlocal method; hence the huge discrepancy between the two tables for Re  $\omega$  and k a with this set of parameters.

TABLE I

ω <u>pi</u> 2 ωci	^k 0 ^a i 2π	Imω _L ^ω ci	Reω _L ^ω ci	^k y ^a i	k0 <mark>xpeak</mark> 2π	k <u>Δx</u> 02π	^k x peak ^k 0	P _{max} needed	
10000	$\begin{array}{c} 0.015\\ 0.03\\ 0.12\\ 0.18\\ 0.21\\ 0.24\\ 0.27\\ 0.30\\ 0.35\\ 0.40\\ 0.50\\ 0.60\\ 0.70\end{array}$	3.55 5.94 9.39 12.26 11.62 13.65 12.50 6.75 3.10 3.18 6.10 10.23 9.35 7.12	6.75 10.65 14.88 14.67 7.83 0. 0. 0. 0. 0. 30.70 24.75 17.79	19.4 24.5 31.4 36.3 34.0 22.1 24.6 27.7 0. 0. 27.3 29.5 29.0 27.1	0.56 0.57 0.57 0.56 0.50 0.50 0.50 0.50 0. 0. 0. 17 0.17 0.18 0.20	0.010 0.015 0.023 0.044 0.070 0.085 0.086 0.092 0.188 0.230 0.117 0.138 0.169	$29.1 \\ -3.2 \\ -15.9 \\ -13.8 \\ -7.0 \\ -0.7 \\ 4.1 \\ 9.1 \\ \pm 12.2 \\ \pm 10.7 \\ 5.5 \\ 3.4 \\ 2.6 \\ 2.2 \end{bmatrix}$	74 36 39 27 14 6 9 14 14 13 9 7 5	
1000	0.015 0.3 0.06 0.12 0.18 0.21 0.24 0.27 0.30 0.35 0.40 0.50 0.60	1.35         2.27         3.73         5.47         5.70         6.85         6.47         3.57         1.66         1.63         3.50         5.56         5.06	2.65 4.31 6.45 7.18 4.20 0. 0. 0. 0. 0. 16.32 13.16 9.53	9.8 11.9 14.8 17.7 17.4 11.1 12.8 14.8 0. 0. 15.2 16.1 15.7	0.20 0.61 0.62 0.63 0.62 0.58 0.50 0.50 0.50 0.50 0. 0. 0. 0. 18 0.18 0.19	0.207 0.019 0.026 0.036 0.063 0.101 0.126 0.124 0.131 0.258 0.327 0.158 0.190 0.231	25.8 7.2 -2.1 -5.2 -3.3 -0.3 -2.0 4.6 ±6.5 ±5.5 3.0 1.8 1.6	4 49 25 15 14 8 4 6 8 8 7 6 4 4 4	
100	0.015 0.03 0.06 0.12 0.18 0.21 0.24 0.27	0.49 0.72 1.23 1.83 1.95 1.78 2.24 1.25	0.86 1.49 2.18 2.47 1.48 0.52 0. 0.	3.4 4.7 5.2 6.2 6.0 5.8 4.4 5.2	0.63 0.65 0.64 0.63 0.58 0.54 0.50 0.50	0.035 0.045 0.063 0.107 0.174 0.210 0.214 0.224	10.4 3.9 -0.2 -1.6 -1.1 -0.7 0.7 1.6	23 14 8 7 4 3 3 4	•
10	0.03 0.06 0.12	0.15 0.42 0.62	0.65 0.79 0.84	2.19 1.99 1.88	0.70 0.67 0.65	0.077 0.109 0.149	0.7 -0.2 -0.7	6 4 4	

TABLE II

$\frac{\frac{\omega_{pi}^{2}}{2}}{\omega_{ci}}$	^k 0 <mark>^a1</mark> 2π	<u>Im ω</u> ^ω ci	<u>Re ω</u> ^ω ci	y ^k y ^a i	0 ^x peak 2π	^k 0 <u>Δx</u> 2 π	$\frac{\frac{k_{x \text{ peak}}}{k_0}}{k_0}$	p _{max} needed
10000	0.21 0.24 0.40 0.50 0.60 0.70 0.85 1.00	12.48 11.63 3.48 7.75 6.75 4.41 2.31 1.16	0. 0. 29.19 23.21 16.25 10.72 0. 0.	22.0 24.7 29.3 30.0 29.7 28.4 26.9 10.6	0.50 0.50 0.22 0.21 0.22 0.26 0.	0.10 0.11 0.15 0.19 0.24 0.36 0.27 0.42	-1 3 5 3 2 2 4 1	6 7 8 5 4 3 5 2
1000	0.18 0.21 0.24 0.27 0.30 0.35 0.40 0.50 0.60 0.70 0.85 1.00	5.43 5.63 5.63 3.18 1.95 1.81 1.76 3.19 2.50 1.16 1.25 0.82	4.46 0. 0. 0. 2.13 11.91 7.99 0. 0. 0. 0.31	17.8 10.7 13.1 15.2 0. 0. 7.5 17.6 16.9 4.2 6.6 7.9	0.59 0.50 0.50 0. 0. 0. 0. 0. 24 0.25 0. 0. 0. 95	0.08 0.12 0.14 0.18 0.05 0.09 0.14 0.24 0.33 0.22 0.31 0.42	-3 -1 2 3 ±5 ±5 -6 2 1 0 1 1	9 5 4 6 8 9 4 - 3 - 3 2 2
100	0.06 0.12 0.18 0.21 0.24 0.27 0.30 0.35 0.40 0.50 0.60 0.70 0.85 1.00	1.19 1.71 1.71 1.50 1.32 0.84 0.36 0.39 0.42 0.38 0.44 0.54 0.54 0.42 0.35	2.17 2.52 1.69 1.04 0. 0. 1.53 1.68 0.83 0.78 0.63 0. 0.35 0.45	5.4 6.6 6.4 6.2 4.5 5.3 3.8 3.1 4.6 2.3 3.6 3.6 4.6 6.3	0.65 0.66 0.63 0.50 0.50 0.40 0.33 0.07 0.22 0.16 0. 0.95 0.93	0.06 0.09 0.13 0.16 0.22 0.27 0.34 0.38 0.25 0.18 0.23 0.32 0.39 0.38	0 -1 -1 -1 0 1 1 2,-3 -1 -1 0 1 1	7 6 5 4 2 2 3 2 5 3 3 1 1 1 1
10	0.06 0.12 0.18 0.24 0.30 0.50	0.38 0.52 0.47 0.15 stable stable	0.75 0.80 0.64 0.41	2.13 2.17 2.20 2.08	0.68 0.68 0.68 0.66	0.15 0.16 0.20 0.30	0 0 0 0	3 3 3 2
3	0.15	stable						

### APPENDIX K

Derivation of Eq. (38):

 $G_{\pm 1,s}(k_x,\omega)$  is given by Eq. (35) for ions in the straight-line orbit approximation and by Eq. (36) for cold electrons. We note that  $G_{\pm 1,s} = G_{-1,s}$  for the ions if  $k_y = 0$ , and this is also true for the electrons if  $k_y = 0$  and  $k_0 << k_x$ . Then, using Eqs. (23) and (25), we have

$$D(x,k_{x},\omega) = -k_{x}^{2} + \sum_{s} G_{0,s}(k_{x},\omega) + 2\cos k_{0} \times \sum_{s} G_{+1,s}(k_{x},\omega)$$
(K1)

The condition  $\partial D/\partial x = 0$  from Eq. (16) can be satisfied if and only if  $k_0 x = 0$  or  $\pi$  in Eq. (K1). We take  $k_0 x = 0$ . Then Eq. (16) becomes

$$-k_{x}^{2} + \sum_{s} [G_{0,s}(k_{x},\omega) + 2 G_{+1,s}(k_{x},\omega)] = 0$$
 (K2)

$$-2k_{x} + (\partial/\partial k_{x}) \sum_{s} [G_{0,s} + 2G_{+1,s}] = 0$$
 (K3)

evaluated at  $k_x = k_L, \omega = \omega_L$ .

For a loss cone ion distribution we have, from Eq. (11), two ion "species", with

$$n_{0,1} = R(R-1)^{-1} n_{0,1}$$
  
 $n_{0,2} = -(R-1)^{-1} n_{0,1}$   
 $v_1 = v_1$   
 $v_2 = v_1 R^{-1/2}$ 

(K4)

(K8)

$$G_{0,e}^{0,e} = -w_{pe}^{0}k_{x}^{k}^{\lambda}w_{ce}^{0}^{k}^{k}$$

$$= -w_{pe}^{0}k_{x}^{k}^{\lambda}a_{e}^{0}(1 + \Delta_{e})/w_{ce}^{0}] + (w_{p1}^{2}/v_{1}^{2})R(R-1)^{-1}$$

$$= -w_{pe}^{0}k_{x}^{k}^{\lambda}\Delta_{e}/\omega_{ce}^{0}$$

$$(K7) \text{ sud } (K3) \text{ become}$$

$$(K7)$$

Eq. (36) gives

For the electrons, assuming  $\omega \ll \omega_{ce}$ ,  $k_0 \ll k_x$ , and  $k_y = 0$ ,

$$[exp(-k_0^{2}a_{1}^{2}/v_{2}) - R^{1}/2exp(-k_0^{2}a_{1}^{2}/2R)] \}$$

$$[exp(-k_0^{2}a_{1}^{2}/v_{2})R(R-1)^{-1}(1-R^{1}/2a_{2}^{2}/2R)] + I(\pi/2)^{1}/2(m/k_{1}v_{1})$$

$$[I + I(\pi/2) w/kv_{3}] R(R-1)^{-1}(1-R^{1}/2)I(\pi/2)^{1}/2 w/k_{1}v_{1}$$

$$[I + I(\pi/2) w/kv_{3}]$$

$$(K5)$$

$$[exp(-k_0^{2}a_{1}^{2}/2R)] + I(\pi/2)^{-1}(1-R^{1}/2a_{1}^{2}/2R)$$

$$(K5)$$

$$[exp(-k_0^{2}a_{1}^{2}/2R)] + I(\pi/2)^{-1}(1-R^{1}/2a_{1}^{2}/2R)$$

$$(K5)$$

$$(K5)$$

$$(K5)$$

$$(K5)$$

$$(K5)$$

$$(K5)$$

For  $\omega << k_x v_1$  (justified <u>a posteriori</u>), we can take  $\zeta_0$ ,  $\zeta_{\pm 1} << 1$  in Eq. (35), which becomes

$$-2k_{\rm L}[1 + \omega_{\rm pe}^{2}(1 + \Delta_{\rm e})/\omega_{\rm ce}^{2}] - (\omega_{\rm pi}^{2}/v_{\rm i}^{2})R(R-1)^{-1}$$
$$i(\pi/2)^{1/2}(\omega_{\rm L}/k_{\rm L}^{2}v_{\rm i})[R^{1/2}(e_{2} + 1) - (e_{1} + 1)] = 0$$
(K9)

where  $e_1 \equiv \exp(-k_0^2 a_i^2/2)$ ,  $e_2 \equiv \exp(-k_0^2 a_i^2/2R)$ , and Eqs. (14a) and (K4) can be used to find  $\Delta_e$ 

$$\Delta_{e} = (Re_{1} - e_{2})(R - 1)^{-1}$$
(K10)

Simultaneously solving Eqs. (K8) and (K9) for  $\omega_L$  and  $k_L$  yields Eq. (38).

We now justify the assumption that  $\omega << k_x v_i$  by using Eq. (38) to find

$$\omega_{L}/k_{L}v_{i} = 0.532i(e_{2} - e_{1}) [R^{1/2}(e_{2} + 1) - (e_{1} + 1)]^{-1}$$
 (K11)

For R = 3 and  $k_{0a_{1}} = 2$  (typical parameters for region C in Figure 6),

Eq. (K11) yields  $\omega_L/k_L v = 0.135$ . The next term in the expansion of the Z function is²⁰

$$Z(\zeta) = i\pi^{1/2} - 2\zeta + \dots$$
 (K12)

For  $k_y = 0$ ,  $\zeta_0 = \zeta_{\pm 1} = \omega/\sqrt{2} k_x v_i$  in Eq. (35), so the relative error in Z( $\zeta$ ) due to finite  $\omega_L/k_L v_i$  is

Relative error =  $2\pi^{-1/2} \zeta = (2/\pi)^{1/2} \omega_L / k_L v_i = 0.108$ 

Since the approximation  $Z(\zeta) \approx i\pi^{1/2}$  is good to within about 10%, Eq. (38) probably good to within about 10% as well. For  $m_i/m_e \leq 3700$ , this error is comparable to or less than the error due to nonlocal effects (i.e. due to finite  $k_0/k_x$ ), so there is no point in using Eq. (K12) to approximate the Z function.

We consider two different configurations of plasma, a slab with ion guiding center density

$$n_i(x,y) = n_0 \exp(-x^2/L^2)$$
 (L1)

and a column with

$$n_{i}(x,y) = n_{0} \exp(-x^{2}/L^{2} - y^{2}/L^{2})$$
 (L2)

The ion distribution function is

$$f_{01}(x, y, v_{x}, v_{y}, v_{z}) = g_{1}(v_{1}, v_{z}, v_{gc}, y_{gc})$$
  
=  $g_{1}(v_{1}, v_{z}, x + v_{y}/\omega_{ci}, y - v_{x}/\omega_{ci})$  (L3)

To apply the Penrose criterion with  $k_y = k_z = 0$ , at x = 0, we will need

$$F(v_{x}) = \int dv_{y} dv_{z} f_{0i}(0, 0, v_{x}, v_{y}, v_{z})$$

$$= \int dv_{y} g_{1i}(v_{1}, v_{y}/\omega_{ci}, -v_{x}/\omega_{ci}) \qquad (L4)$$
where  $g_{1i}(v_{1}, x, y) \equiv \int dv_{z} g_{1}(v_{1}, v_{z}, x, y).$ 

We take  $g_{i}$  to be a loss cone distribution of the form given by Eq. (11). Then Eq. (L4) becomes

$$F(v_{x}) = [2\pi v_{1}^{2}(1 - R^{-1})]^{-1} \int dv_{y} n_{1}(v_{y}/\omega_{c1}, -v_{x}/\omega_{c1})$$
  
$$[exp(-v_{1}^{2}/2v_{1}^{2}) - exp(-Rv_{1}^{2}/2v_{1}^{2})]$$
(L5)

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For a slab,  $n_i(x,y)$  is given by Eq. (L1), and Eq. (L5) becomes

$$F_{slab}(v_{x}) = (2\pi)^{-1/2} [v_{i}(1 - R^{-1})]^{-1} n_{0}$$

$$[(1 + 2v_{i}^{2}/L^{2}\omega_{ci}^{2})^{-1/2} \exp(-v_{x}^{2}/2v_{i}^{2})$$

$$-(R + 2v_{i}^{2}/L^{2}\omega_{ci}^{2})^{-1/2} \exp(-Rv_{x}^{2}/2v_{i}^{2})]$$
(L6)

For a column,  $n_i(x,y)$  in Eq. (L5) is given by Eq. (L2), and

$$F_{col}(v_{x}) = (2\pi)^{-1/2} [v_{i}(1 - R^{-1})]^{-1} n_{0}$$

$$[(1 + 2v_{i}^{2}/L^{2}\omega_{ci}^{2})^{-1/2} \exp(-v_{x}^{2}/2v_{i}^{2} - v_{x}^{2}/L^{2}\omega_{ci}^{2})$$

$$-(R + 2v_{i}^{2}/L^{2}\omega_{ci}^{2})^{-1/2} \exp(-Rv_{x}^{2}/2v_{i}^{2} - v_{x}^{2}/L^{2}\omega_{ci}^{2})] \qquad (L7)$$

According to the Penrose criterion, an instability occurs when

$$\int_{-\infty}^{\infty} \frac{F(0) - F(v_{x})}{v_{x}^{2}} dv_{x} < 0$$
 (L8)

(The electrons are ignored since they are unimportant for this instability.) Both Eqs. (L6) and (L7) are of the form

$$F(v_x) = A \exp(-av_x^2) - B \exp(-bv_x^2)$$
 (L9)

Putting Eq. (L9) into inequality (L8) yields

$$A^2a < B^2b$$
 (L10)

for instability. Thus the column is marginally stable, and the slab is unstable, to electrostatic perturbations with k in the x-direction, at x = 0, according to the Penrose criterion.

This analysis is cruder than the analysis in Appendix K since it does not find the normal modes but assumes there is a normal mode with  $\phi(\underline{x}) \approx \exp(-i\underline{k}\cdot\underline{x})$  at  $\underline{x} = 0$ . However, it makes the physical mechanism of the instability more clear; it is due to a relative deepening of the loss cone in  $F(v_x)$  as a result of  $d^2n_i/dx^2 < 0$ , and  $d^2n_i/dx^2 < d^2n_i/dy^2$ .

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